

Interferometría... ¿Magia o Ciencia?



Rubén Herrero-Illana

Introduction

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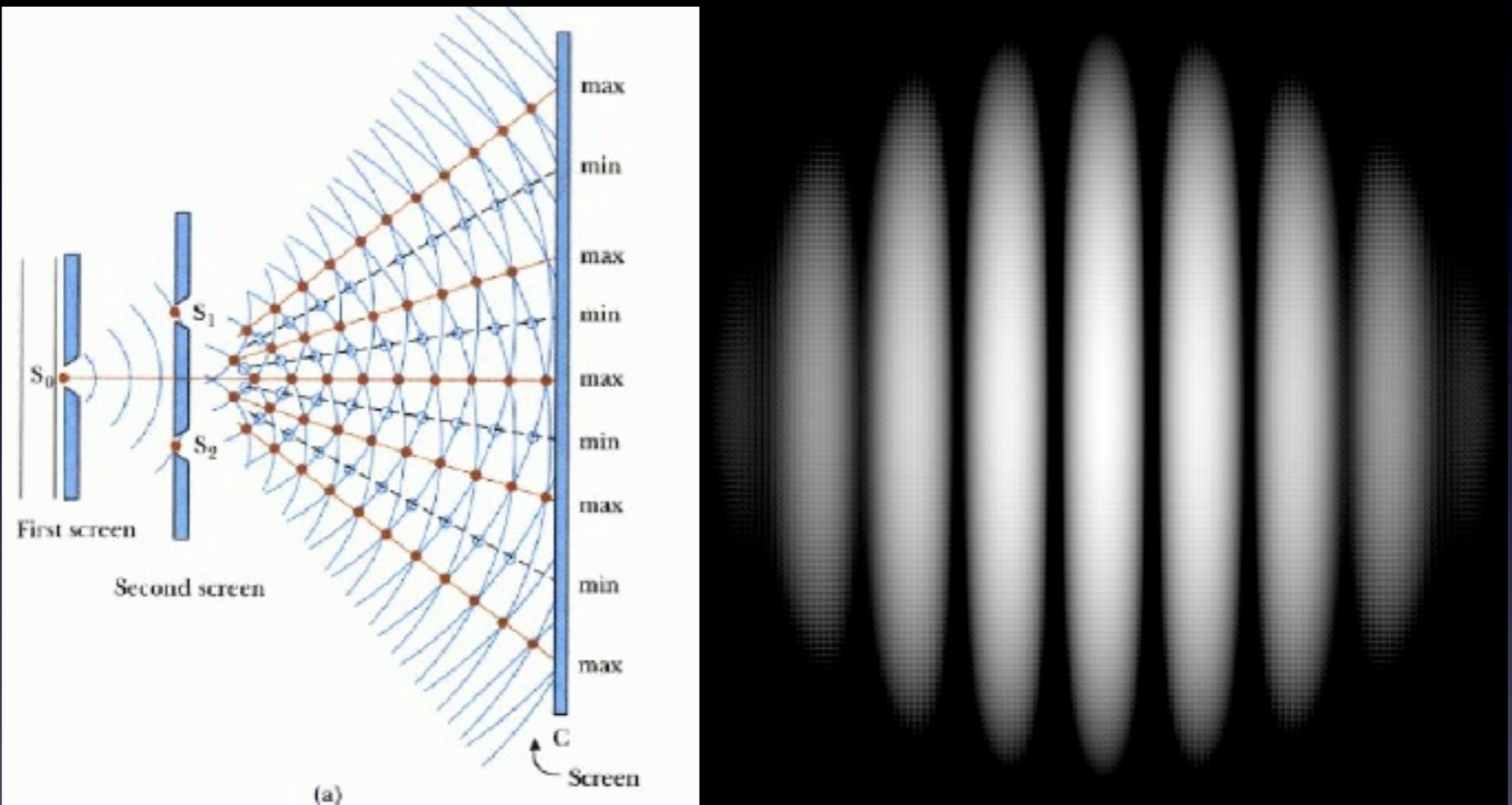
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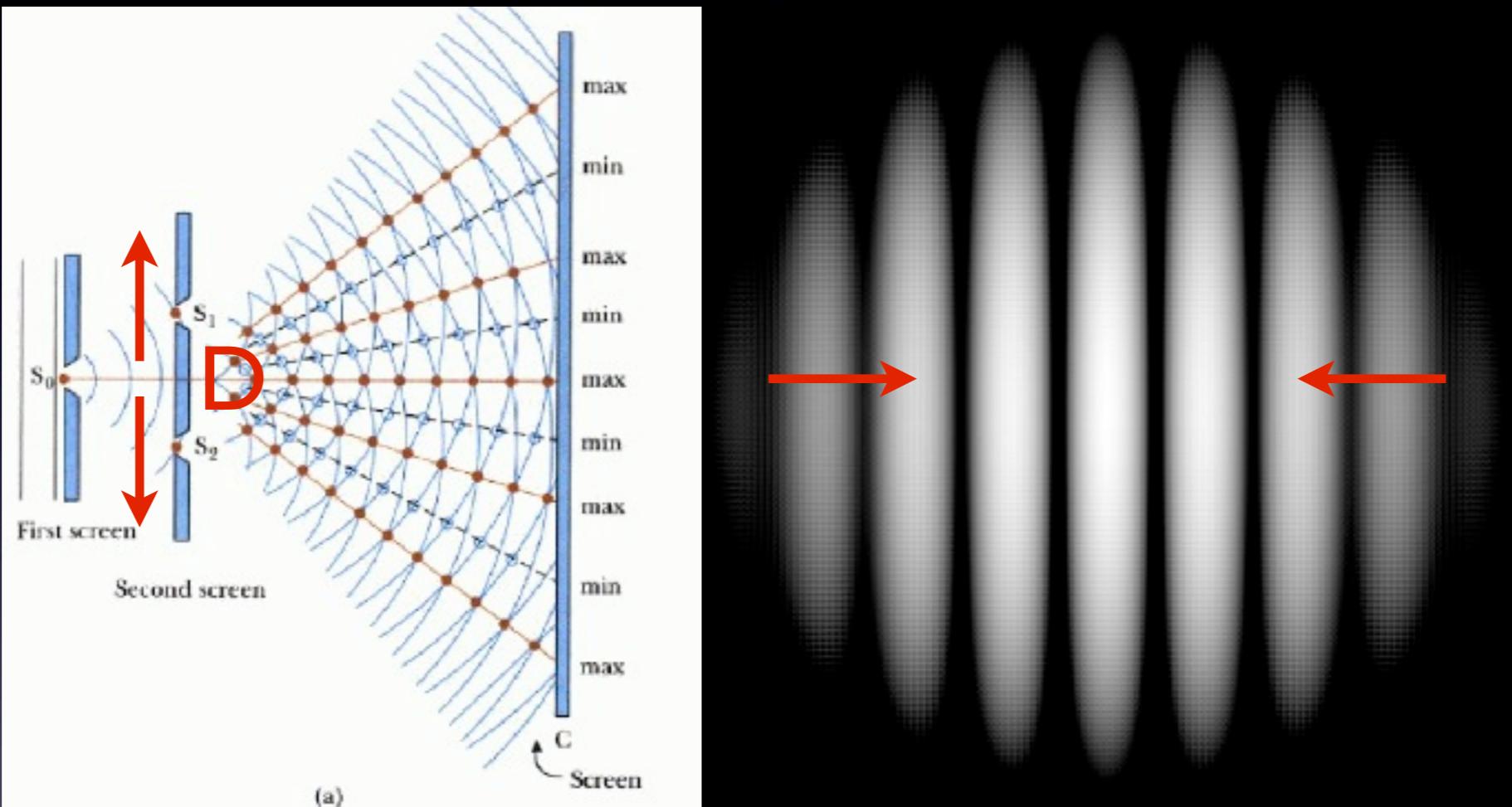


Basics

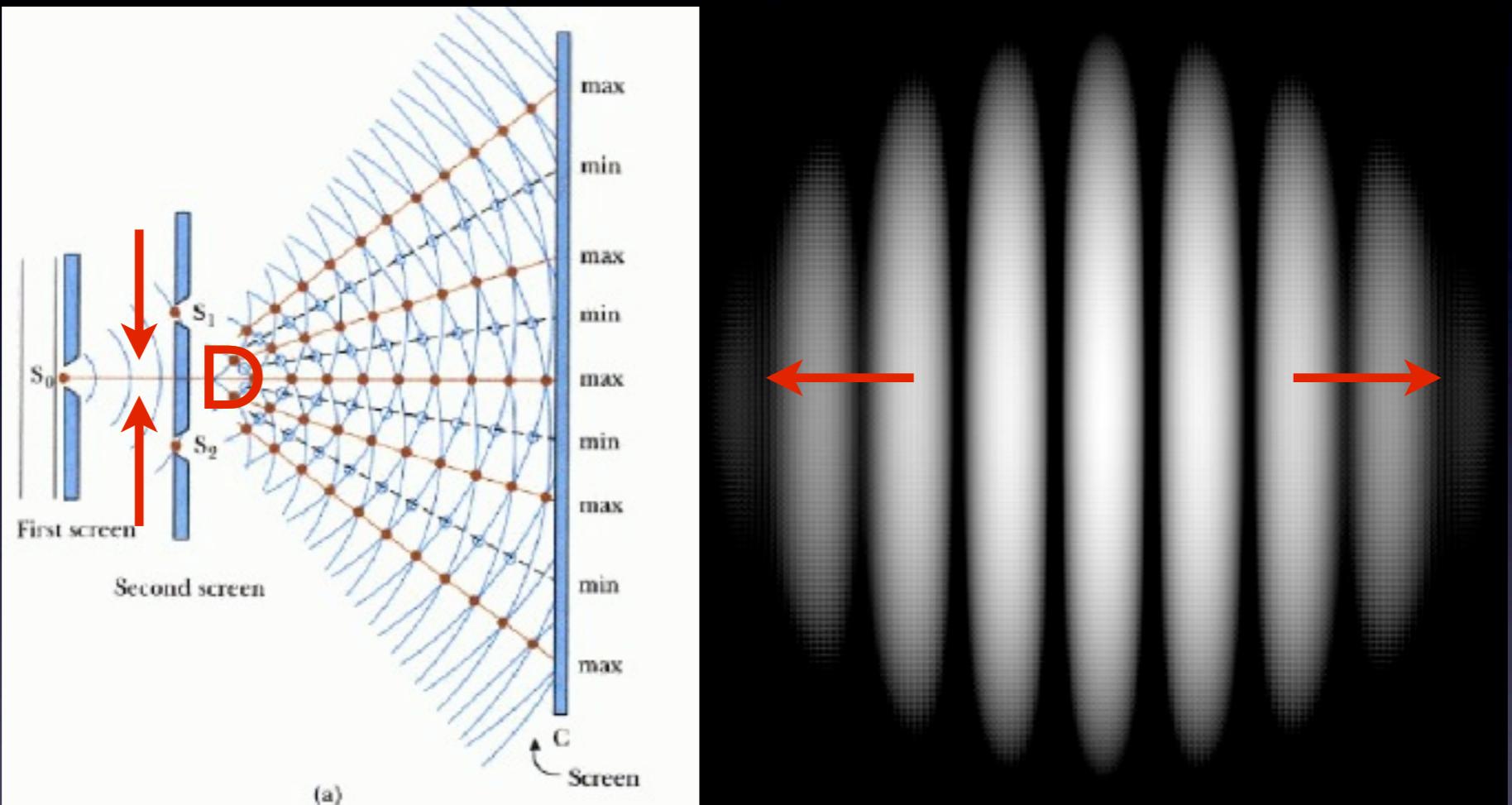
Double slit experiment



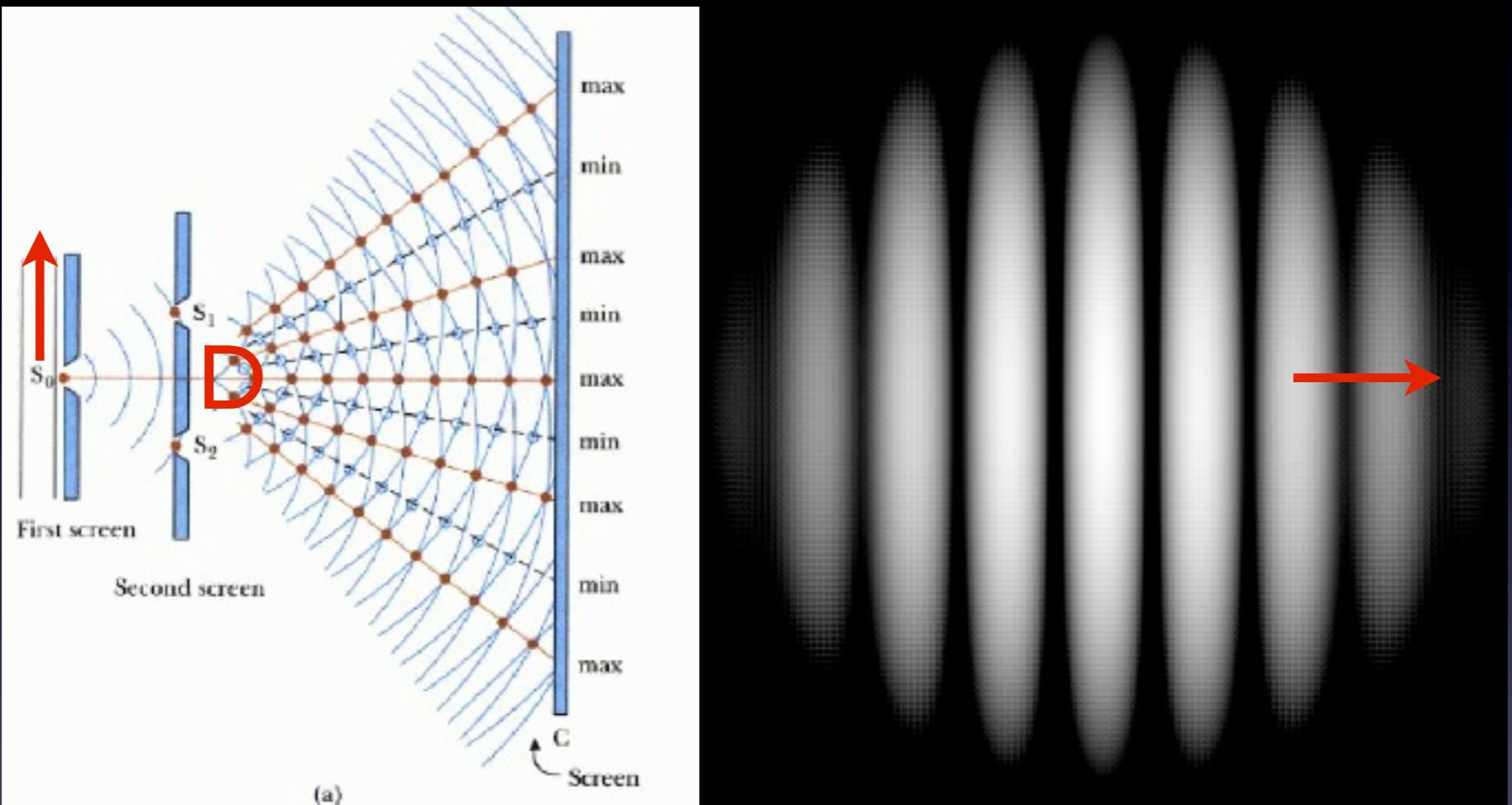
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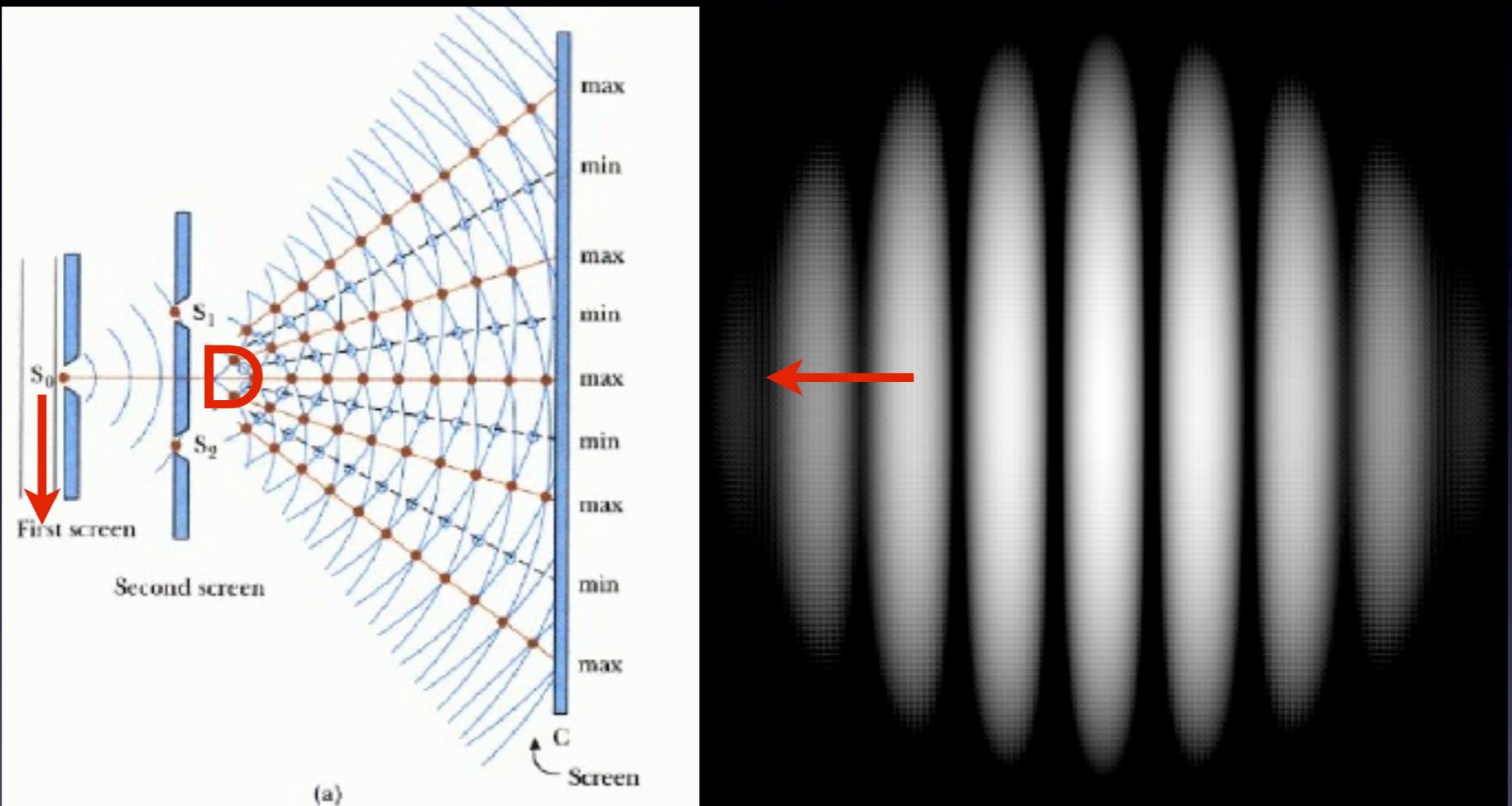
Double slit experiment



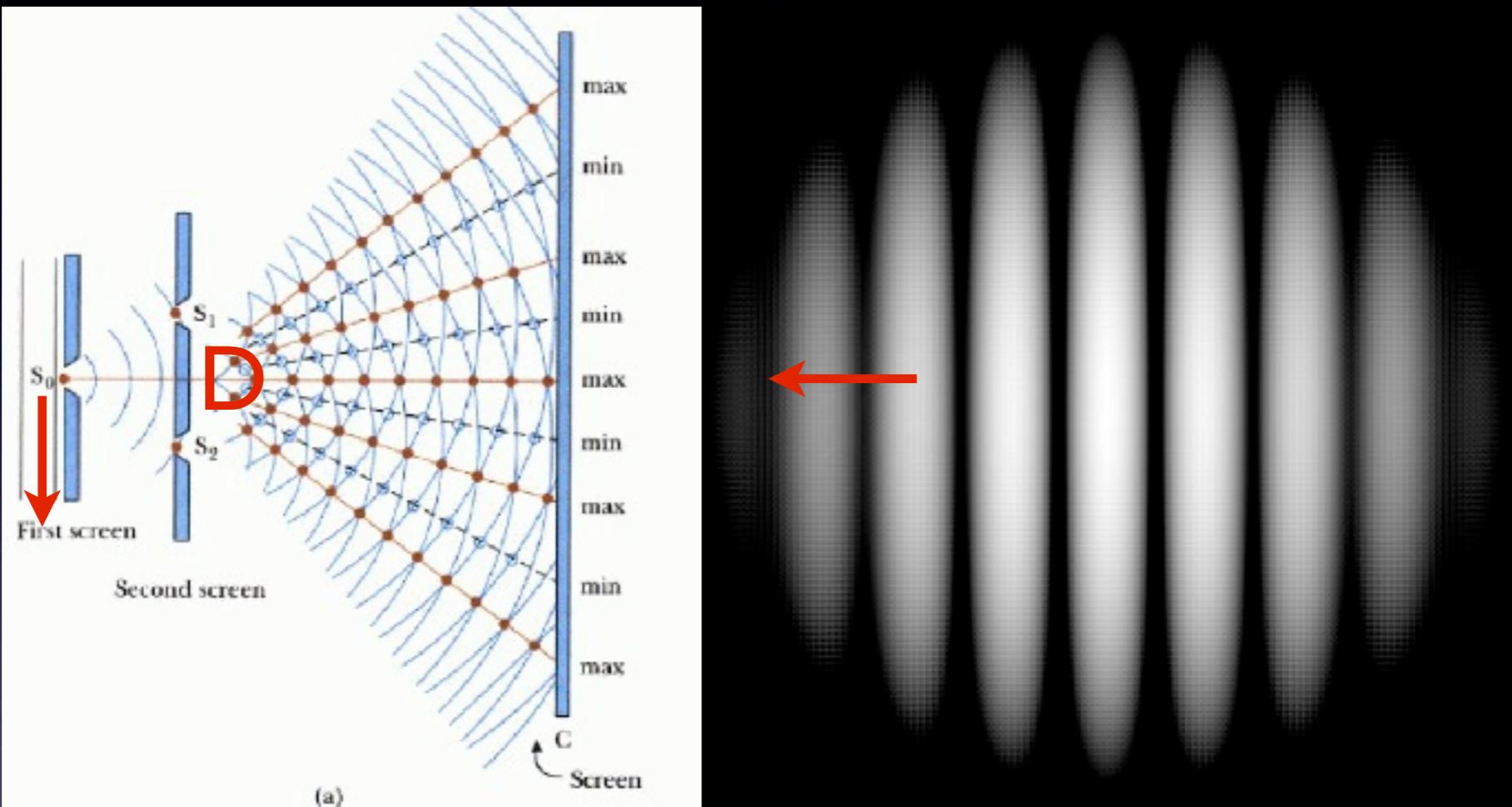
Double slit experiment



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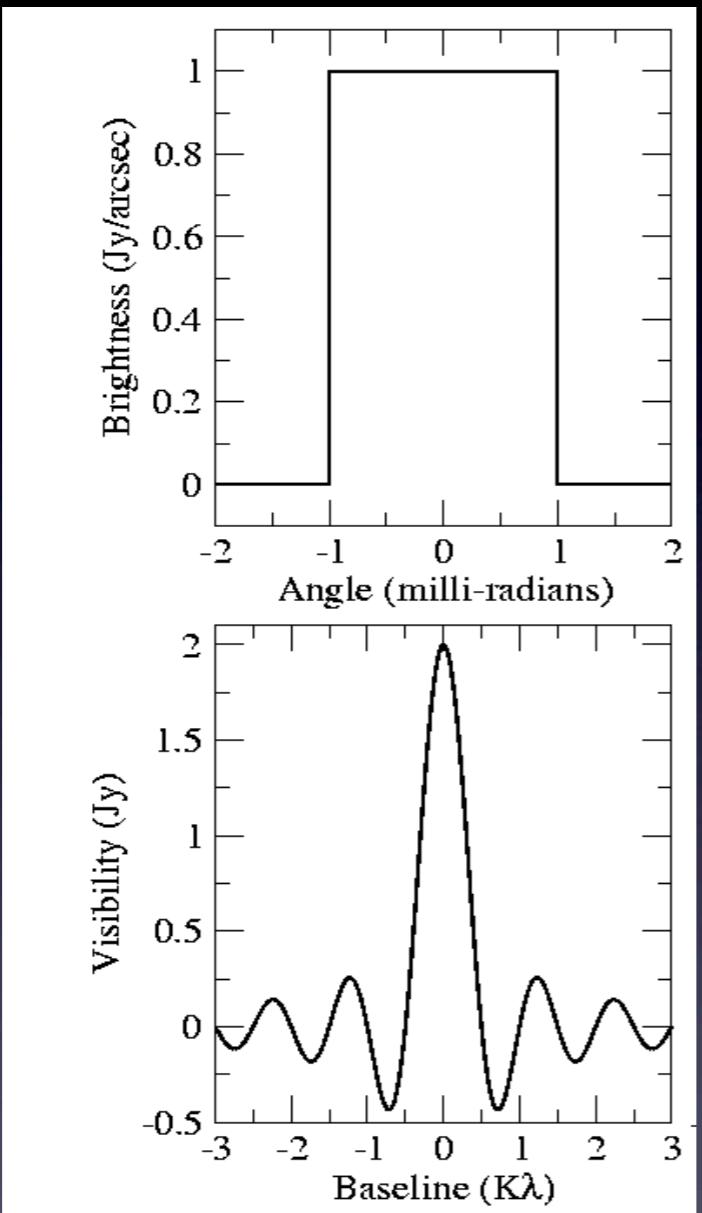
Double slit experiment



$$V = |V|e^{i\phi}$$

$$|V| = \frac{I_{max} - I_{min}}{I_{max} + I_{min}}$$

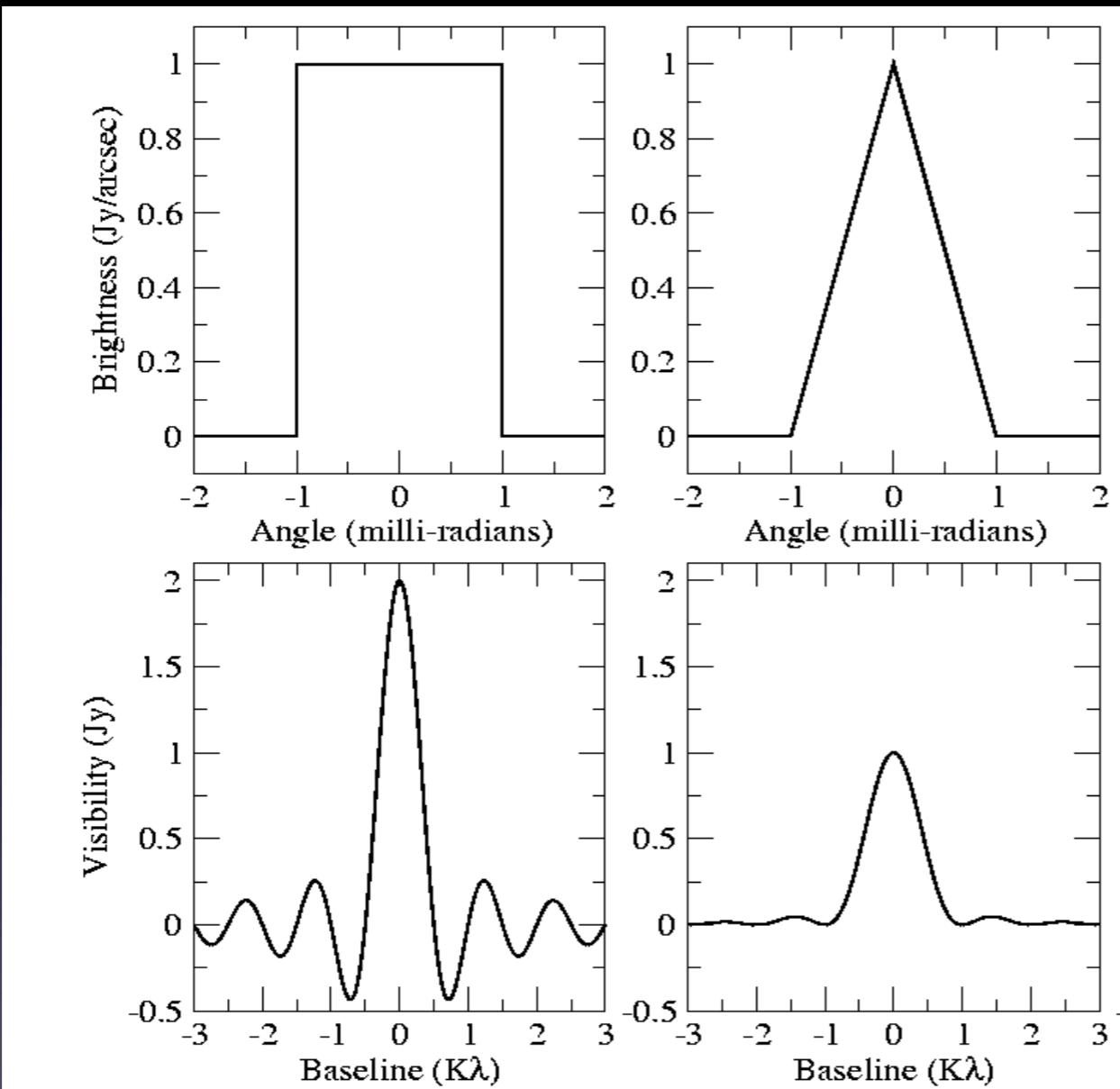
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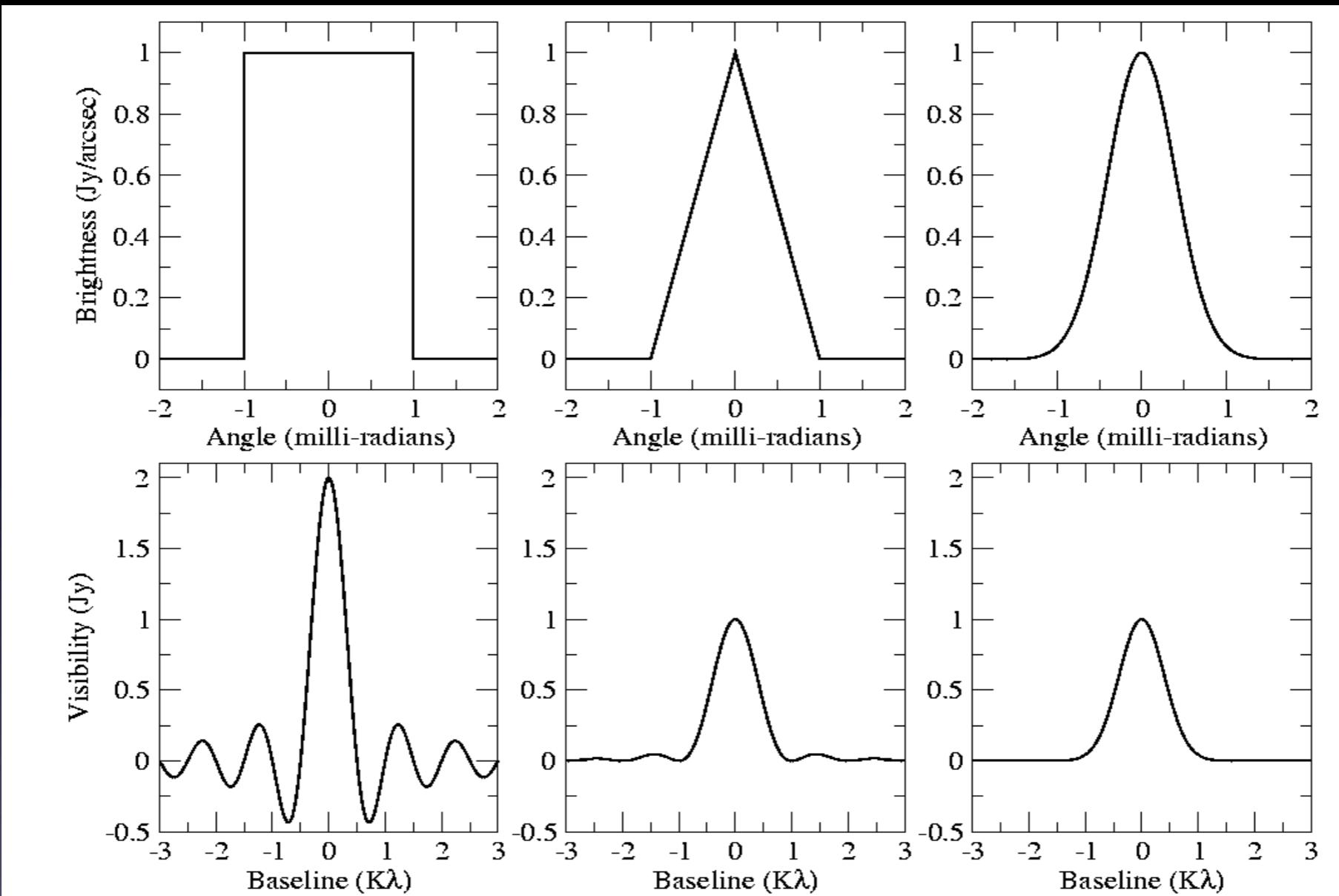
Double slit experiment



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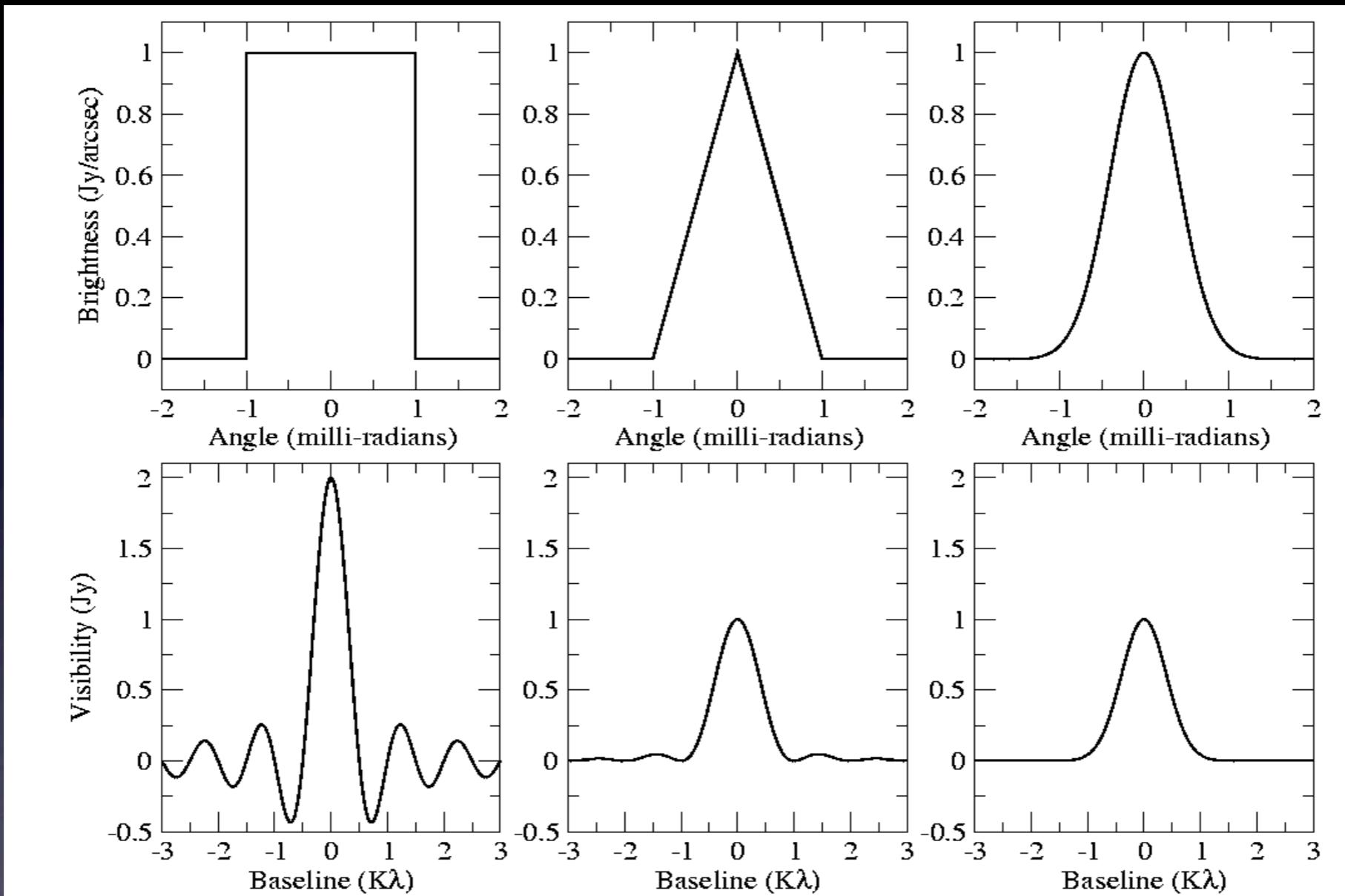
Double slit experiment



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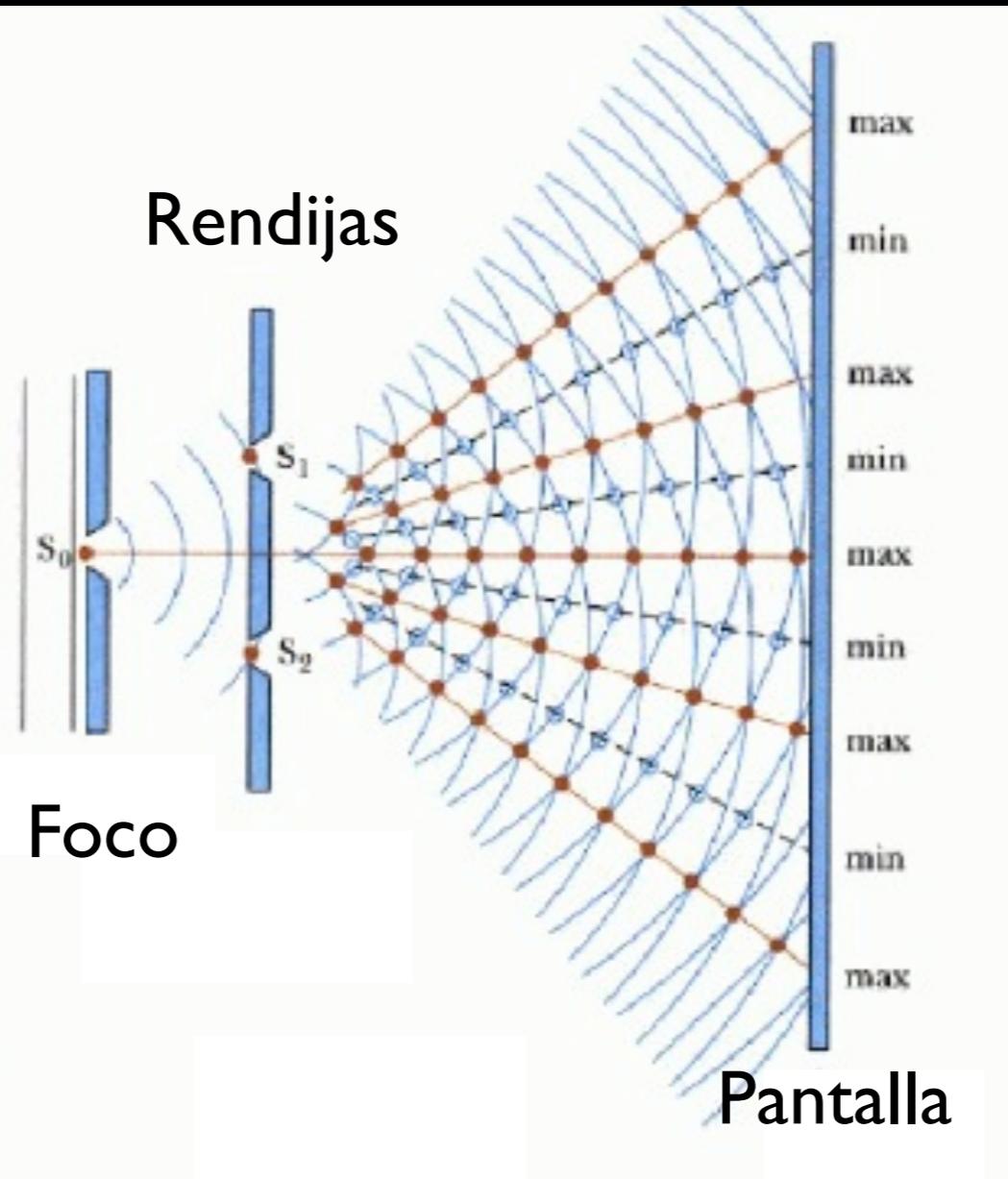


FT!!

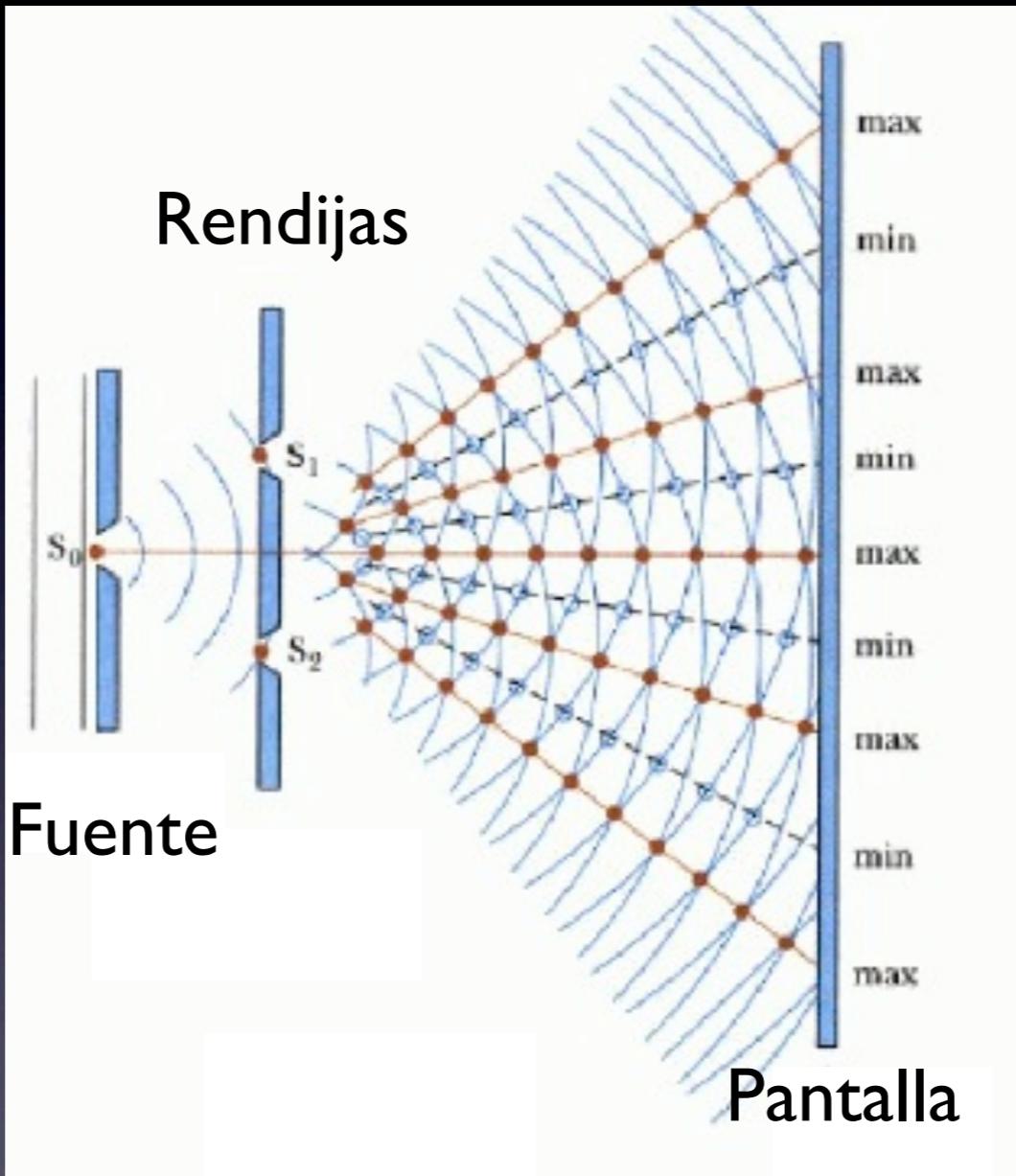
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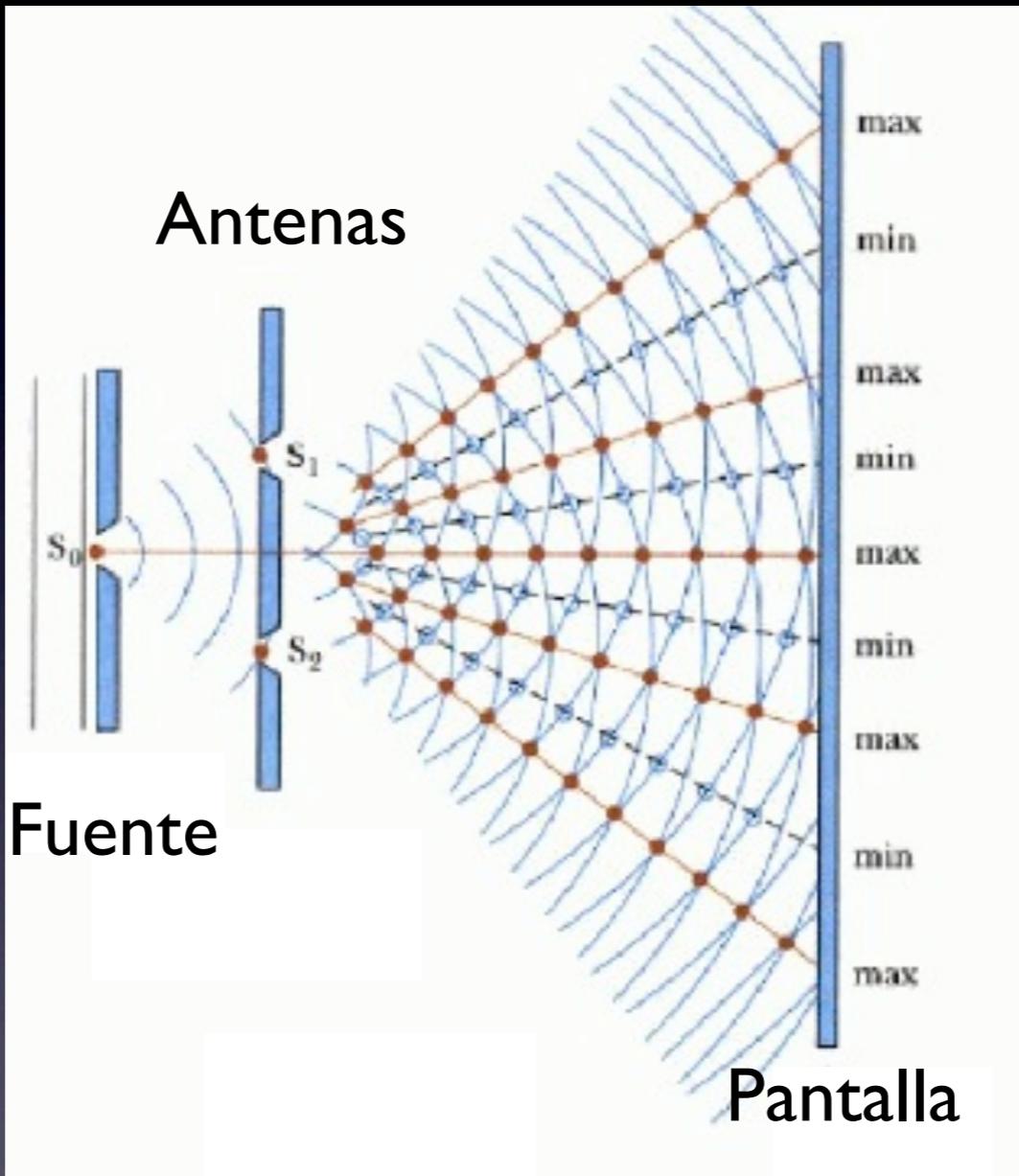
Double slit experiment



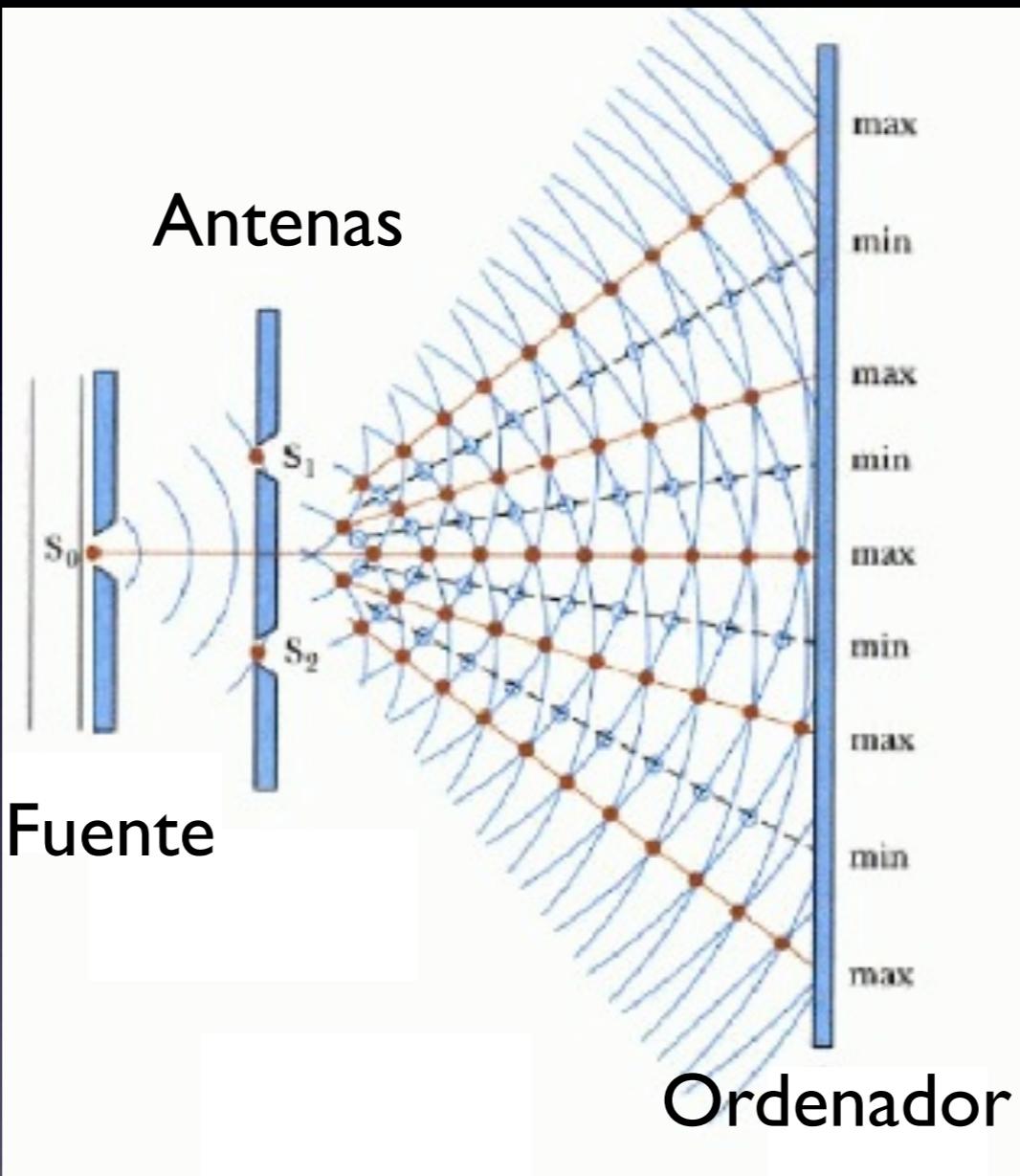
Double slit experiment



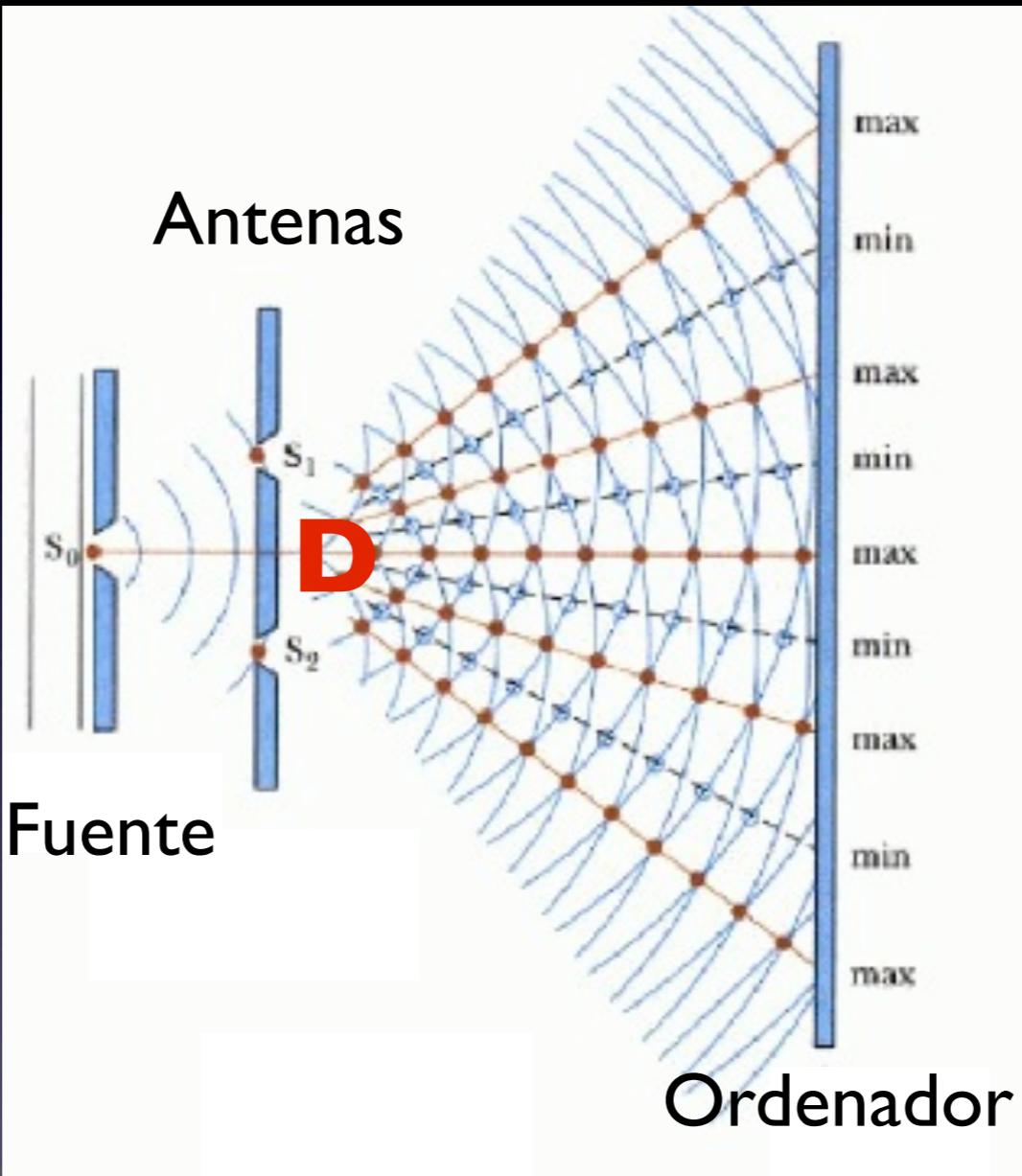
Double slit experiment



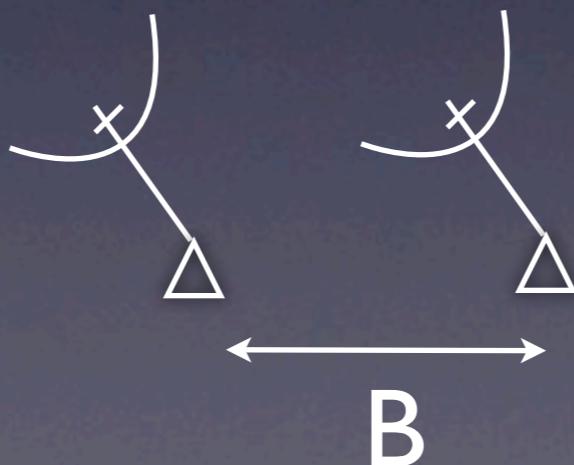
Double slit experiment



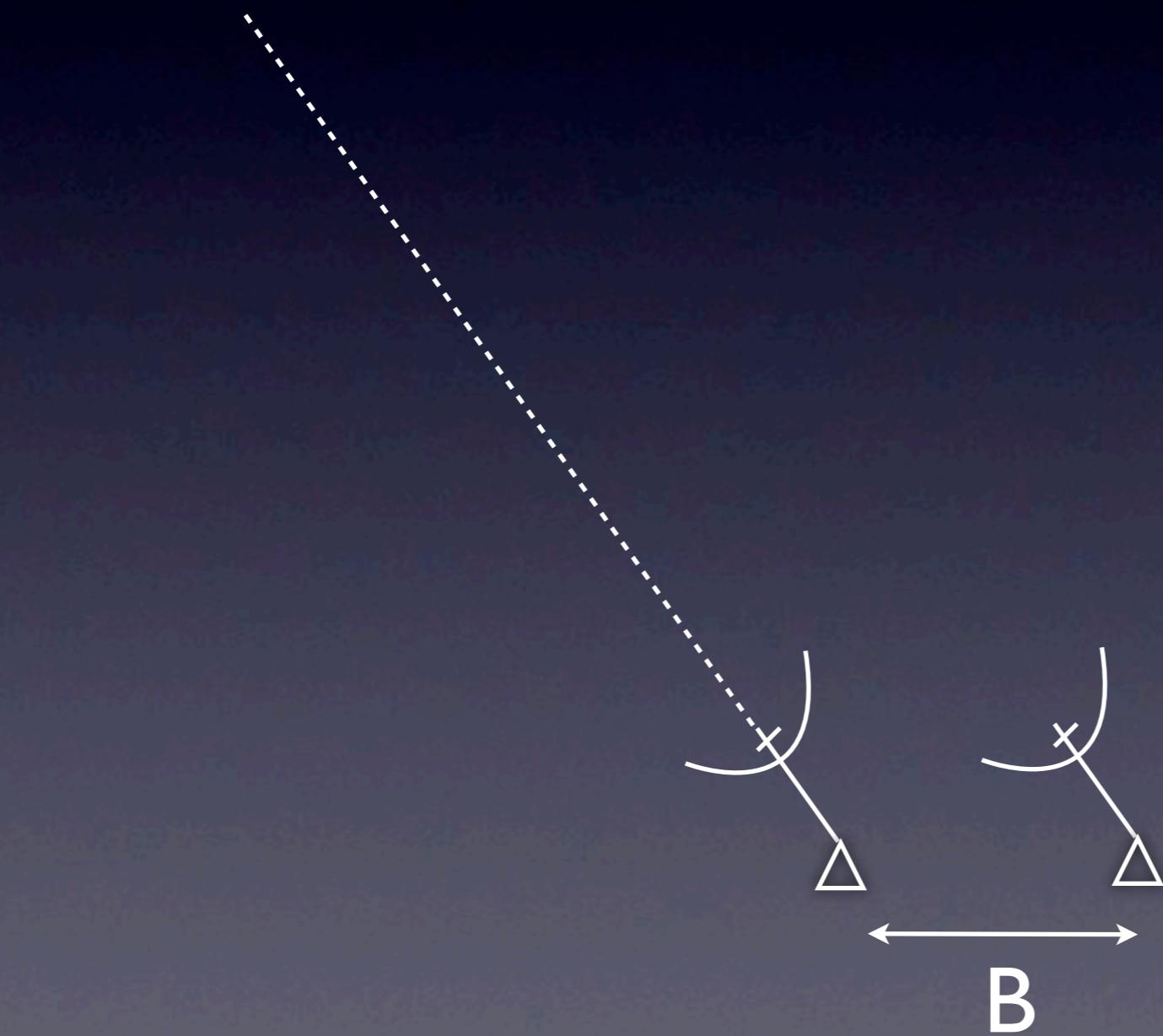
Double slit experiment



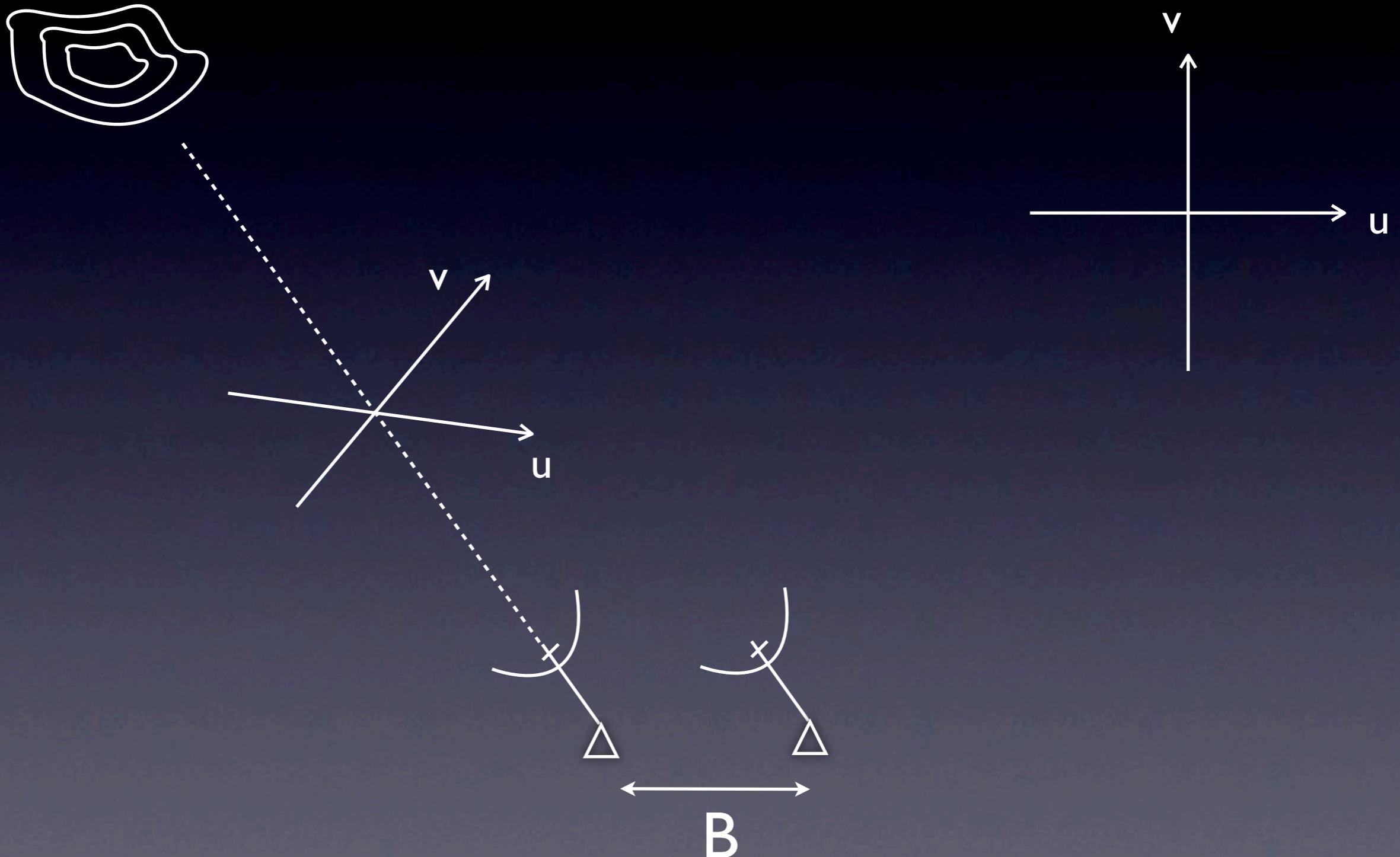
u-v plane



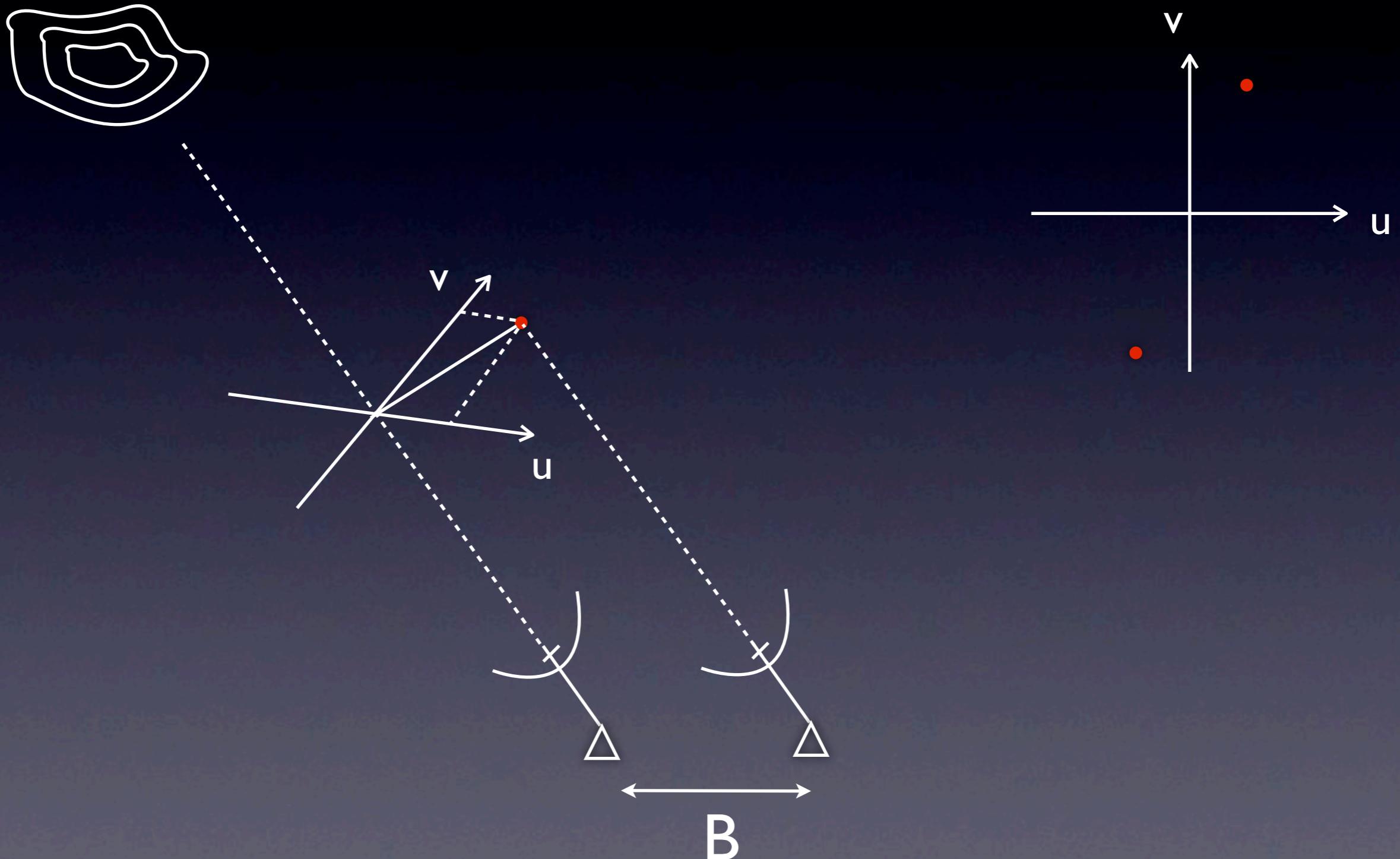
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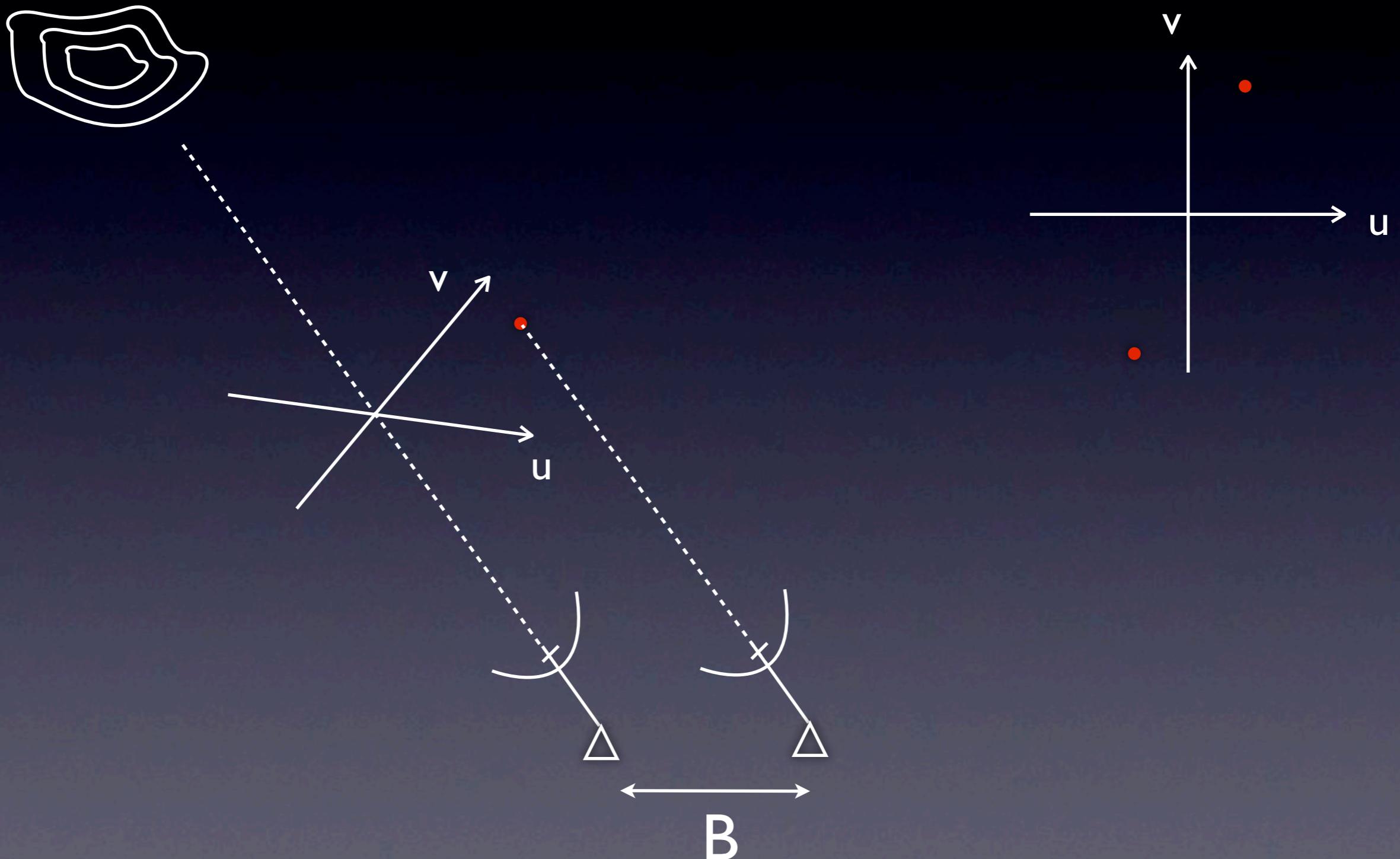
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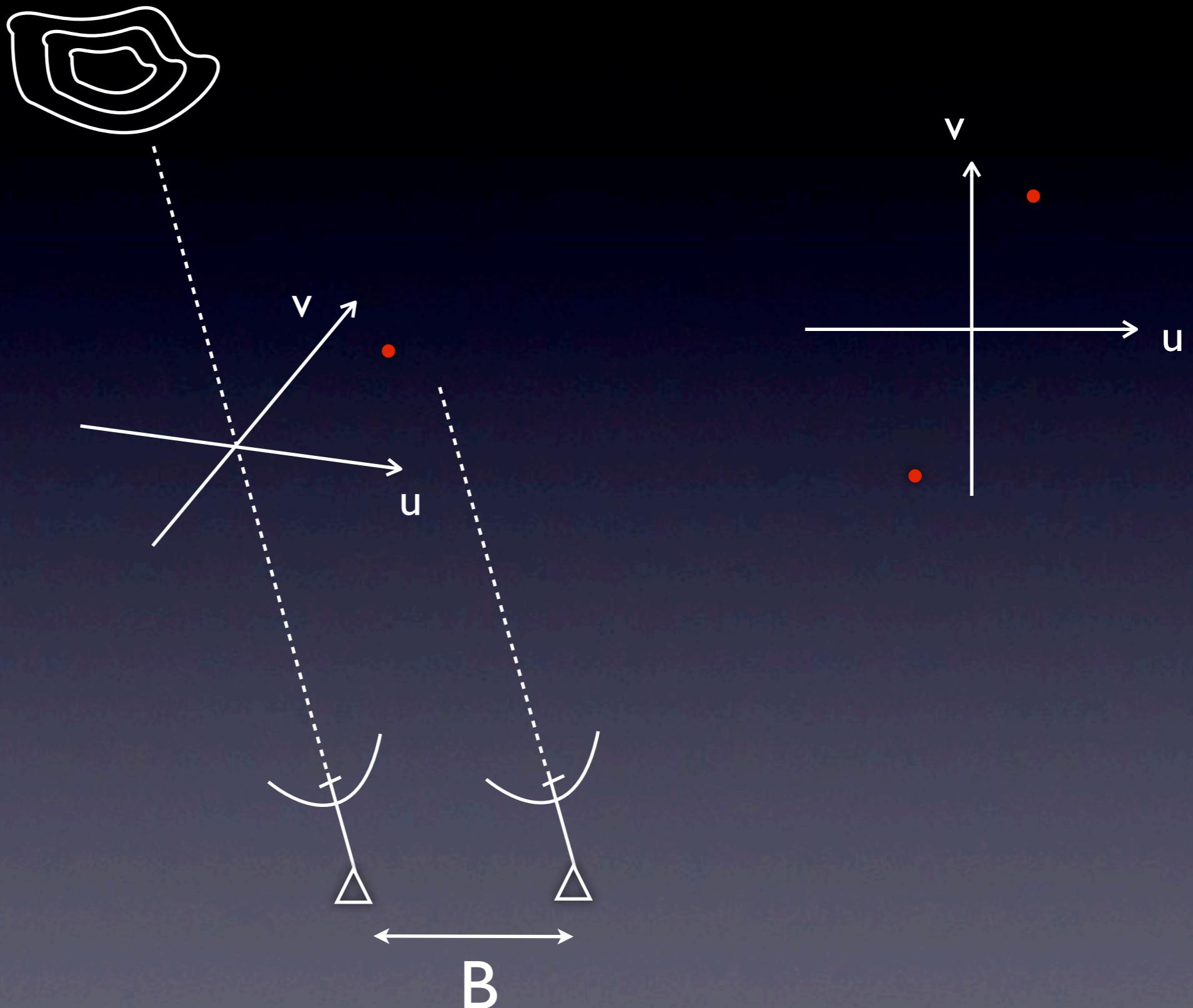
u-v plane



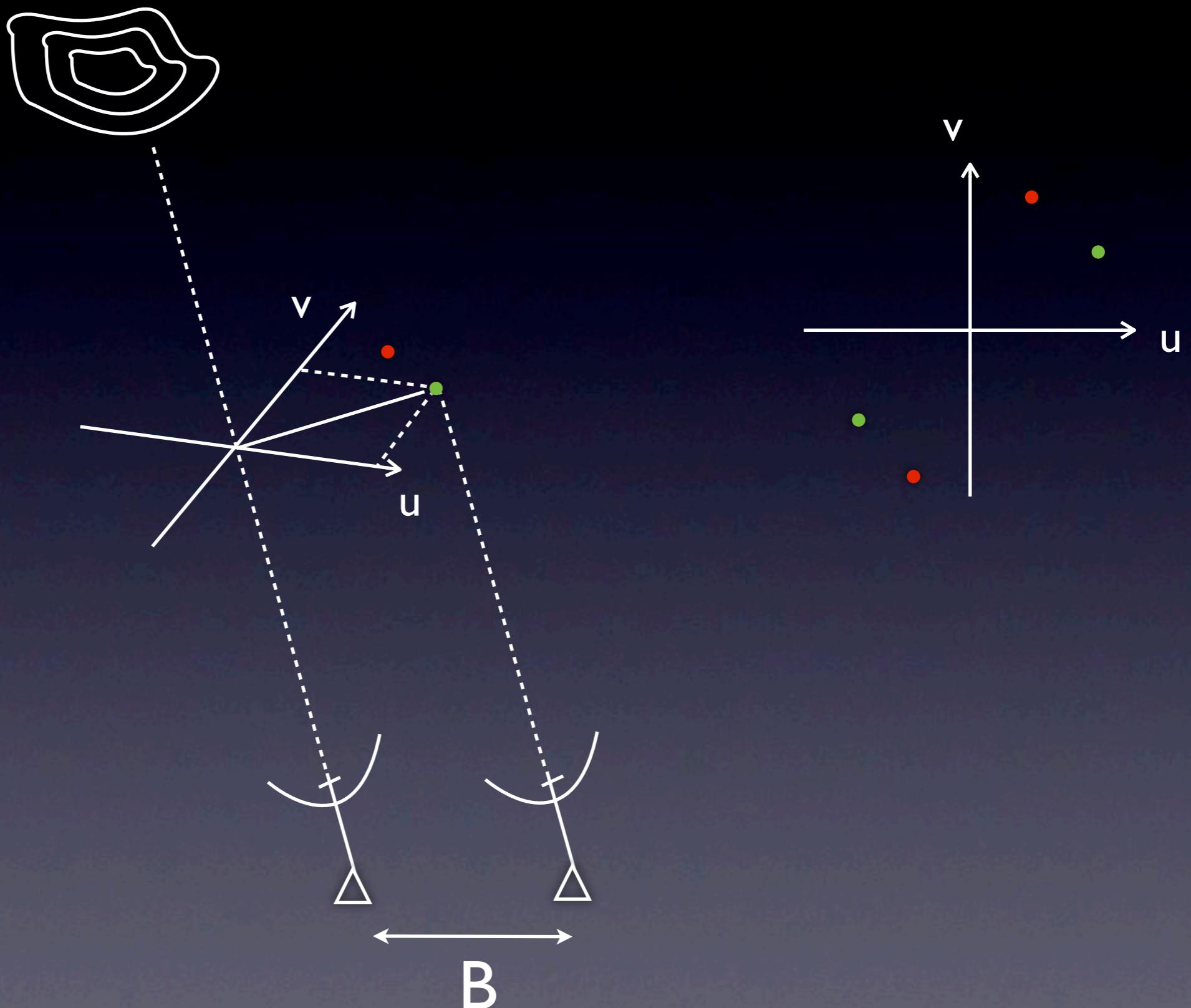
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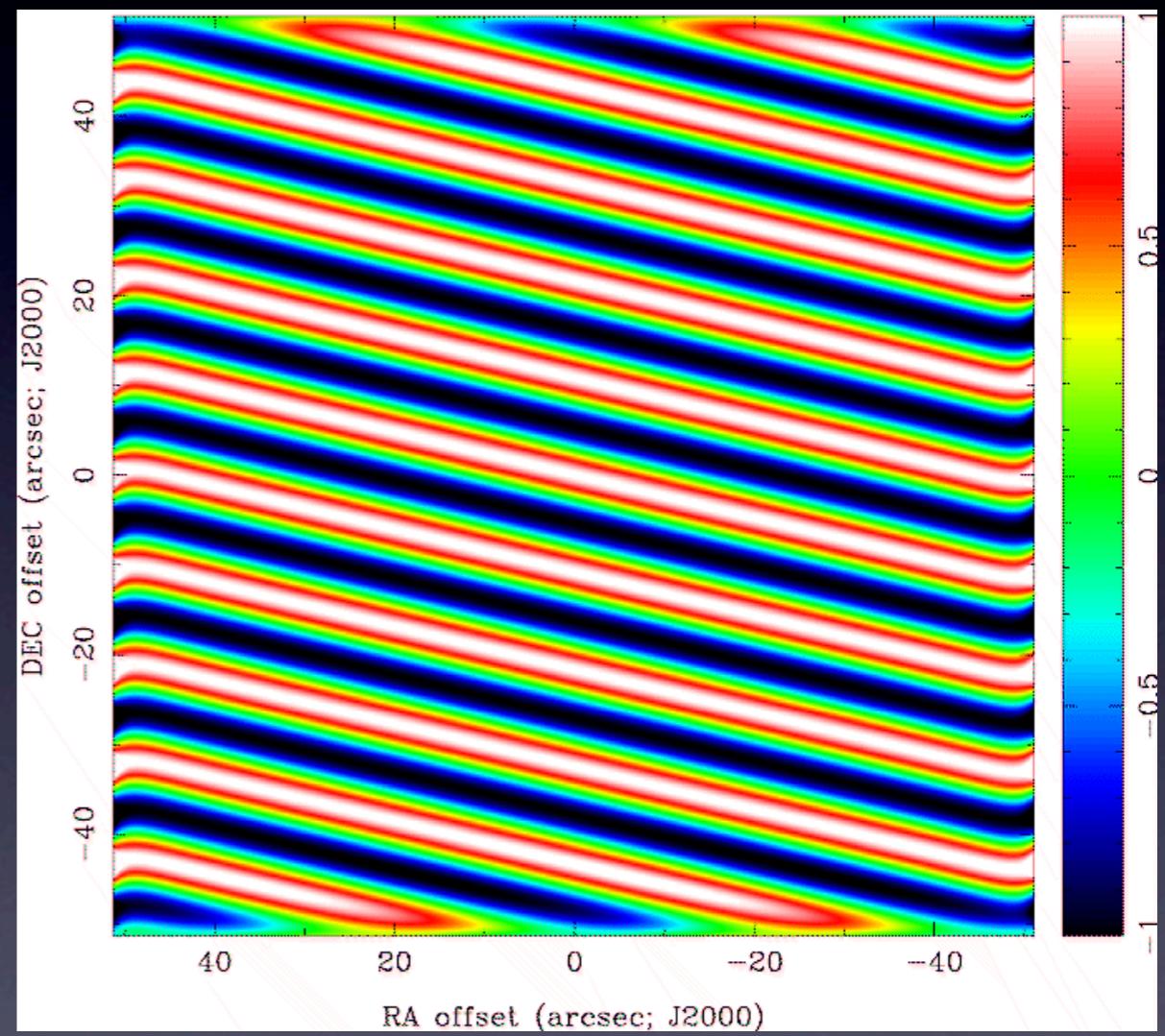
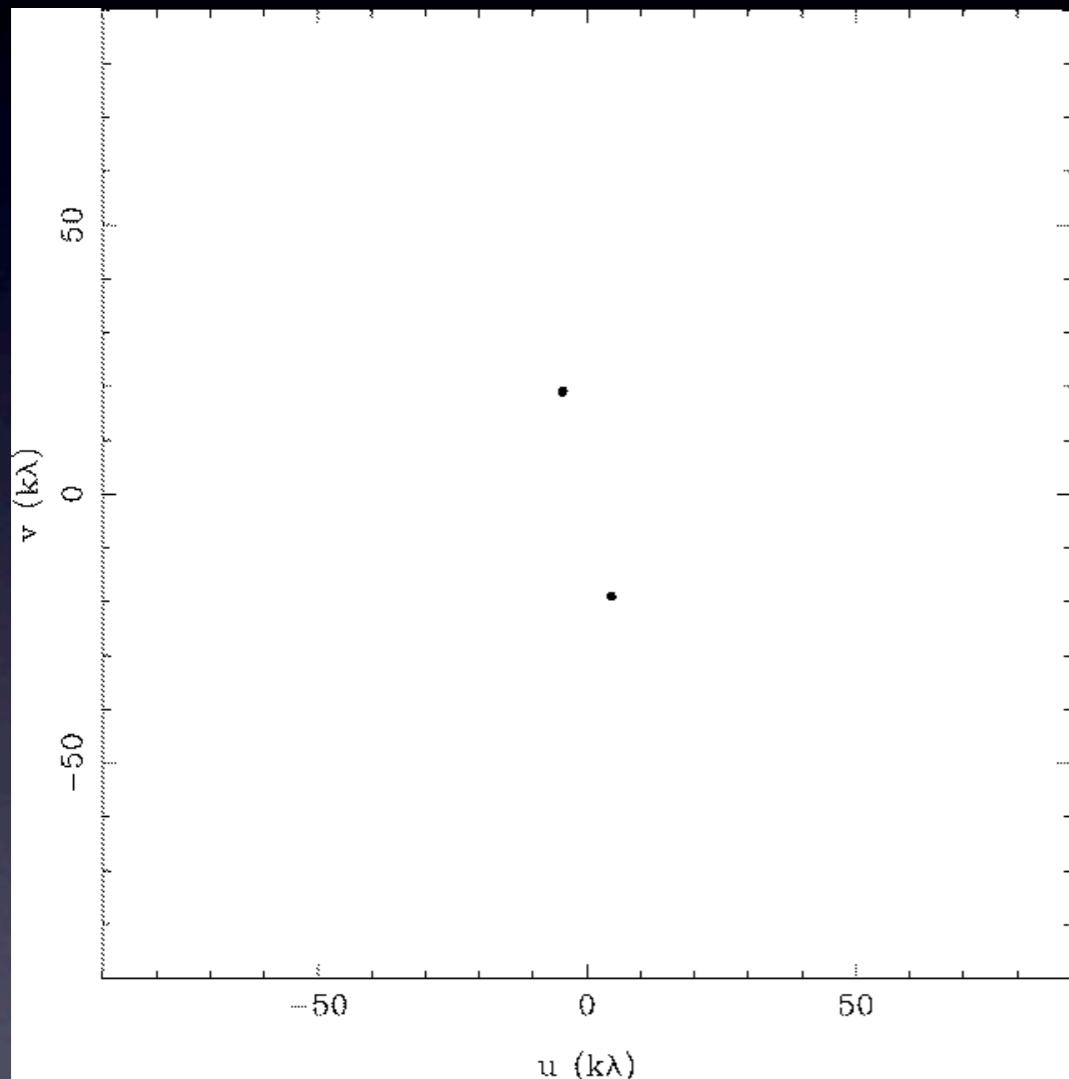


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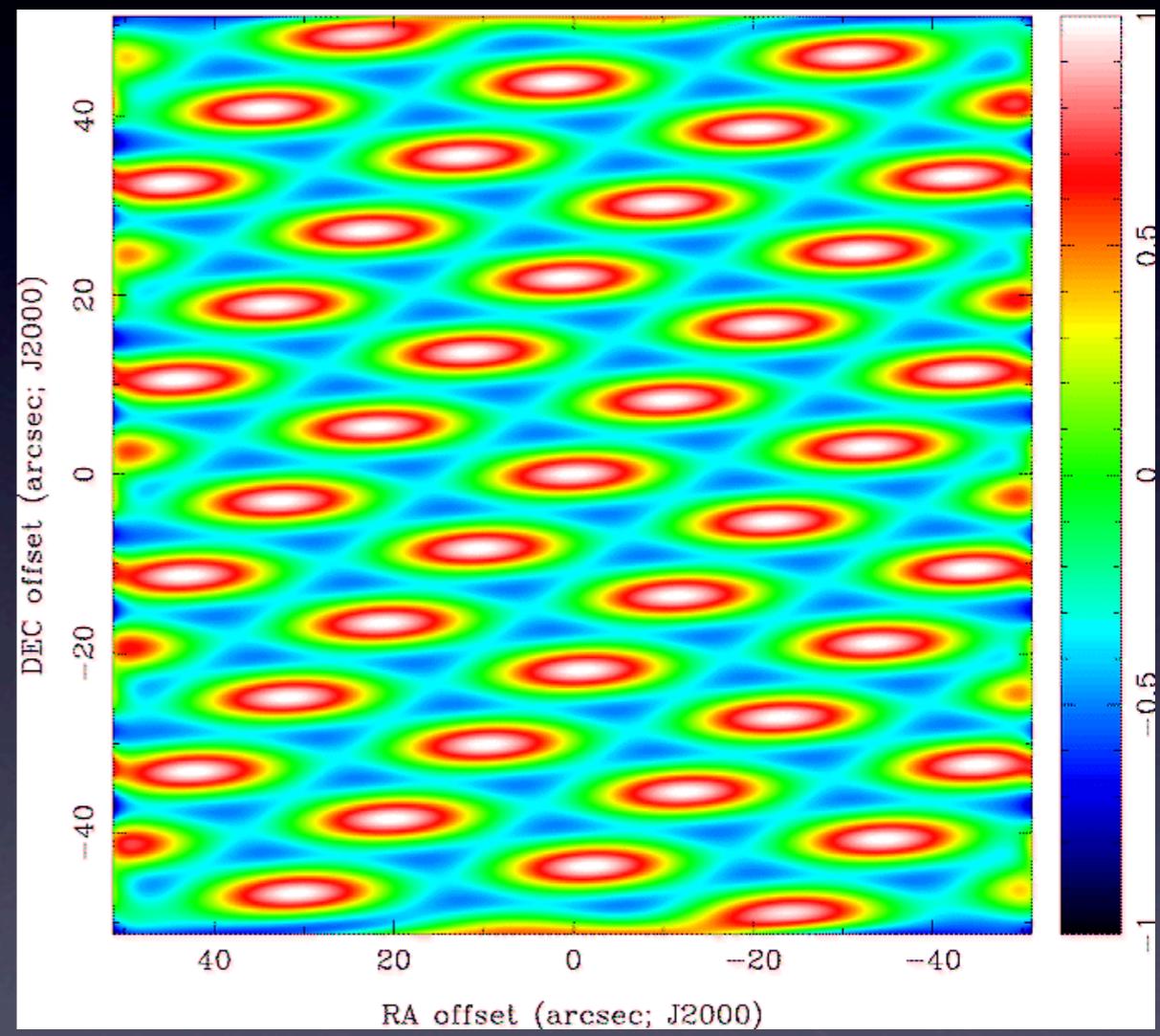
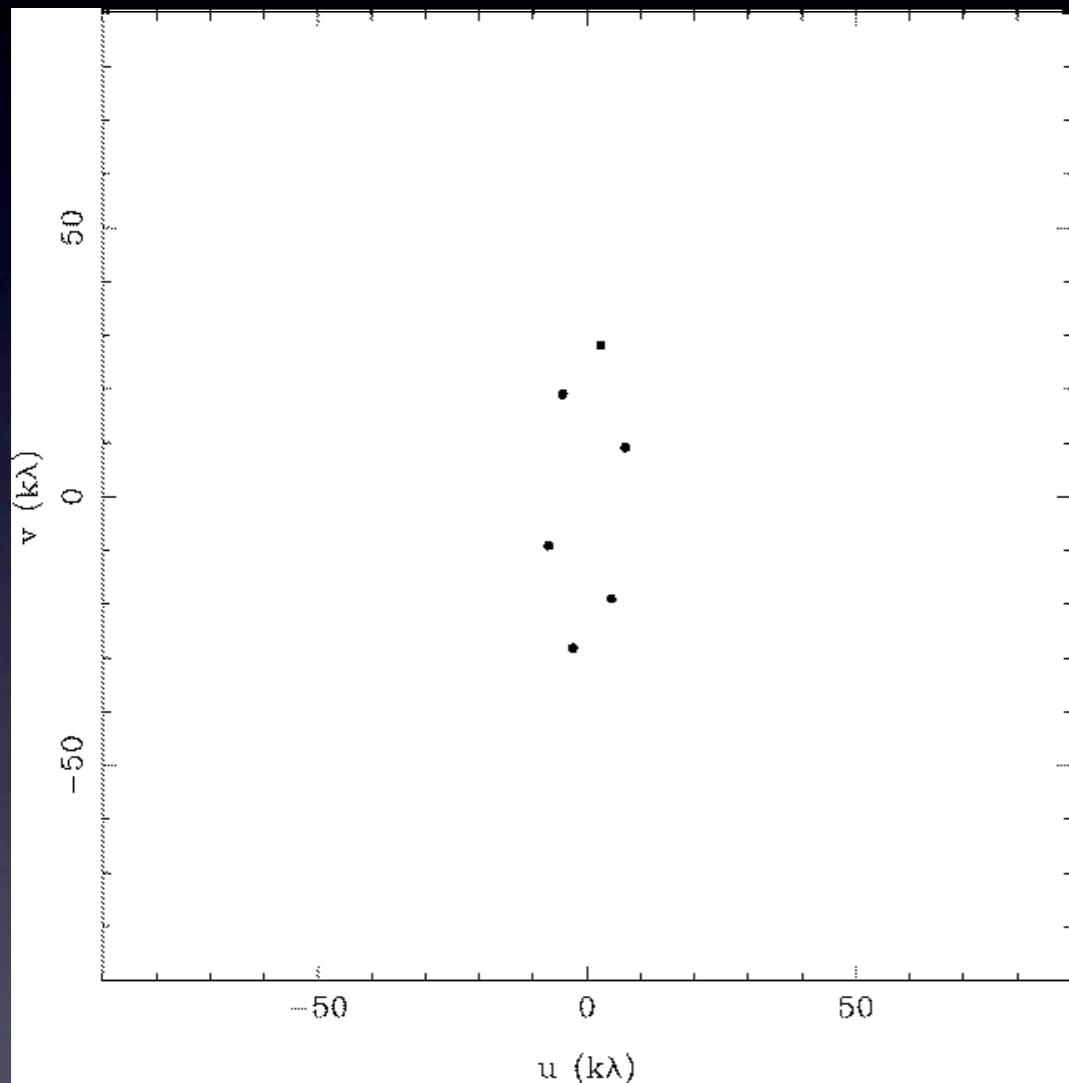
u-v coverage: Dirty beam

2 antennas



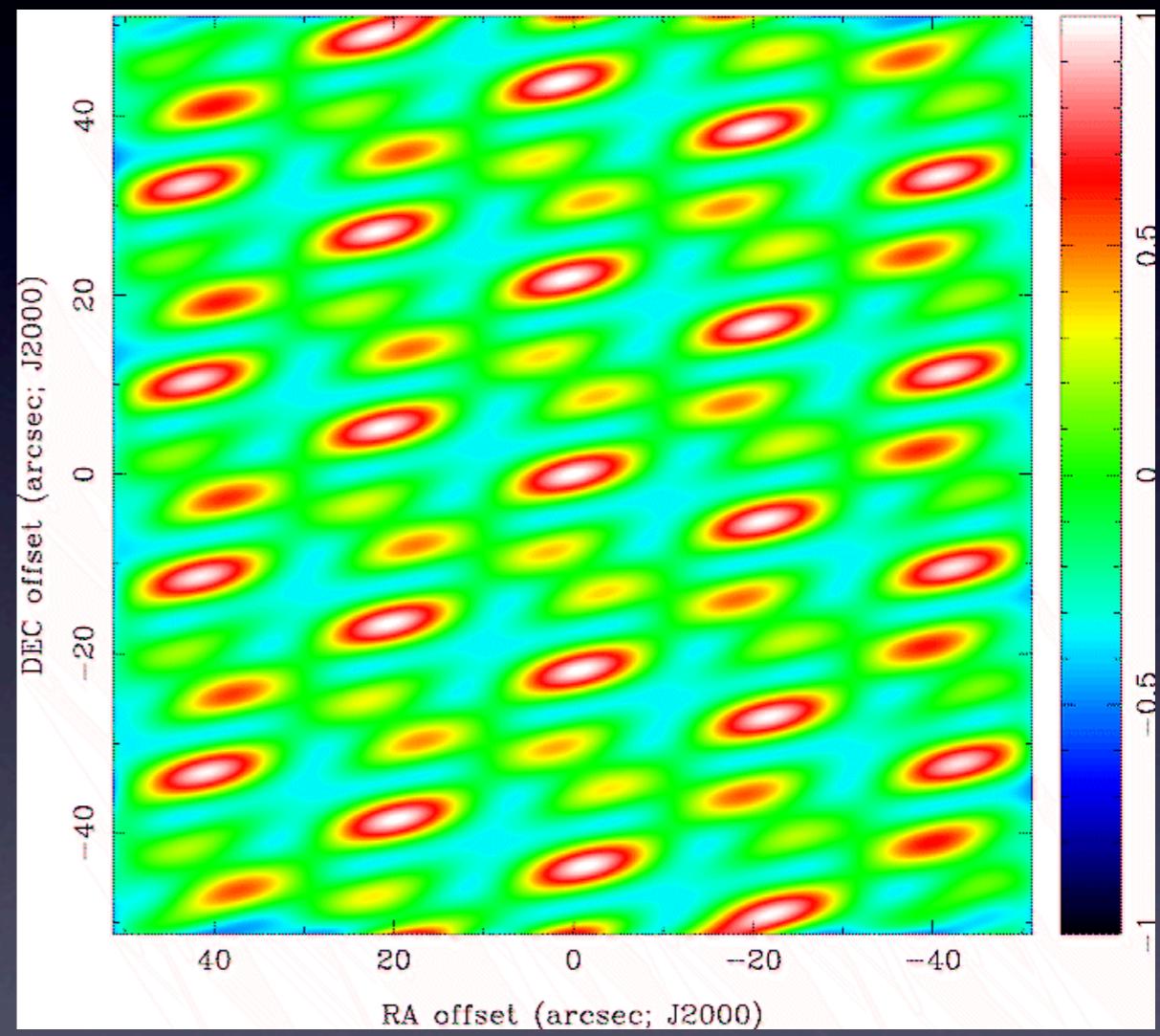
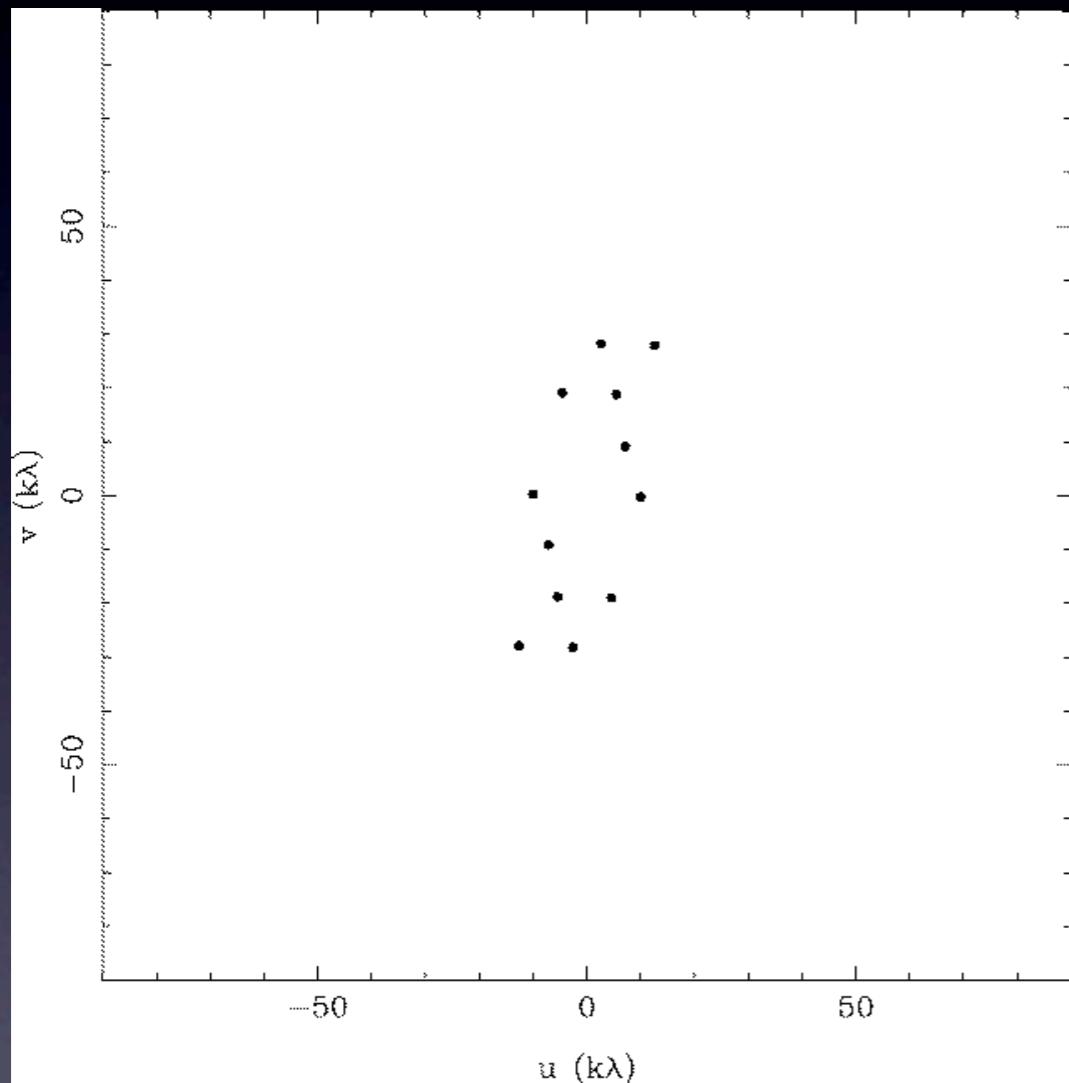
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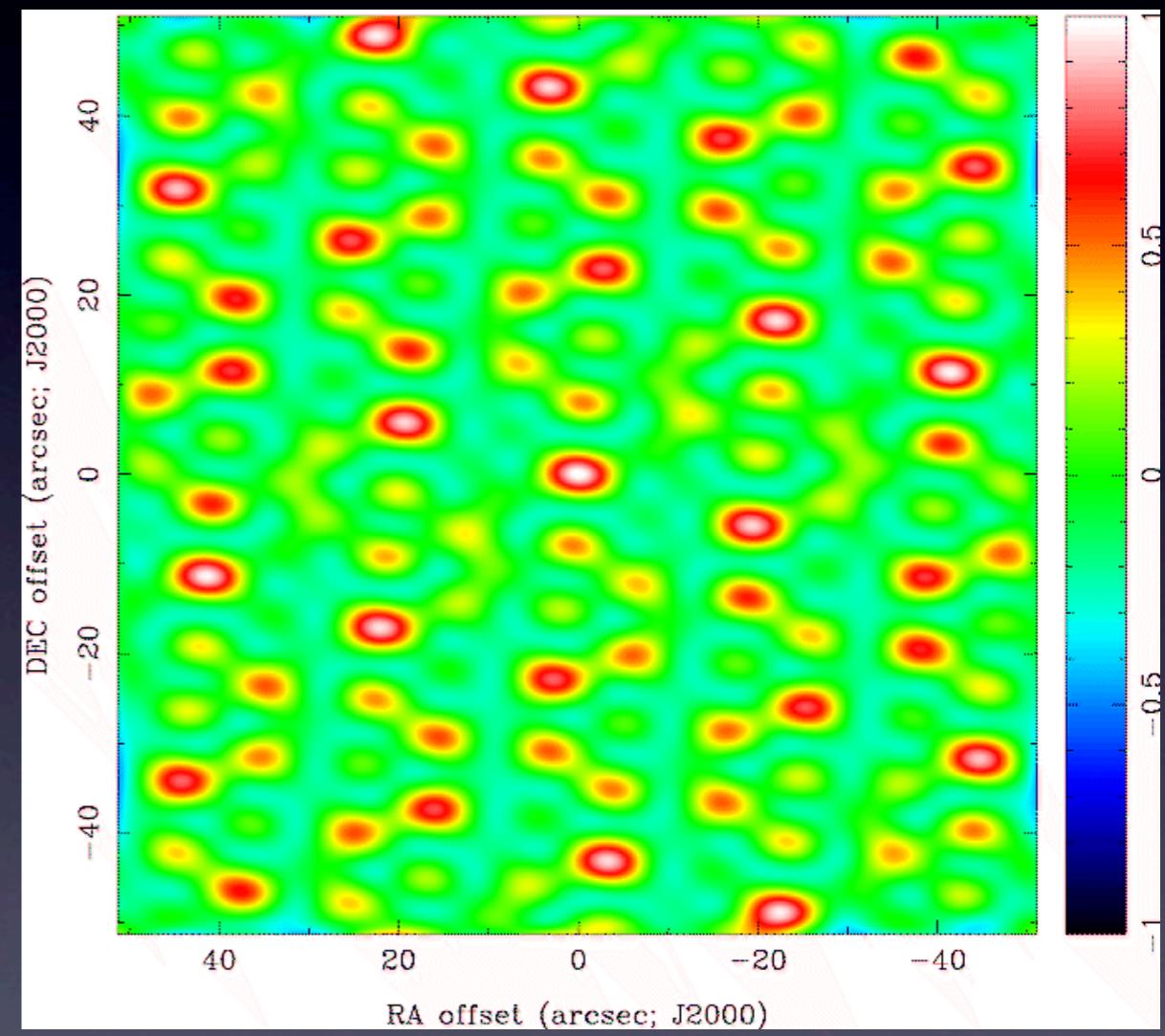
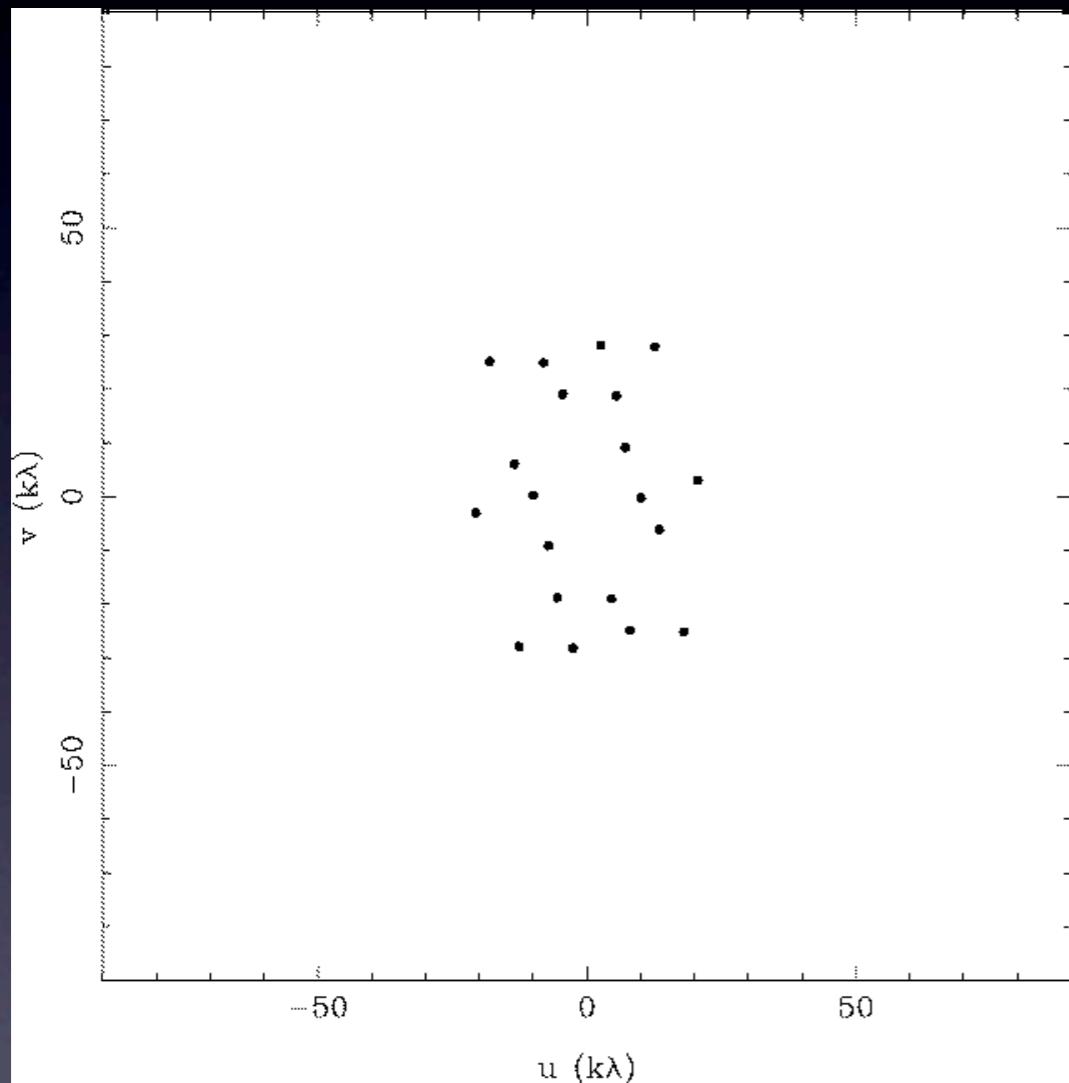
u-v coverage: Dirty beam

4 antennas



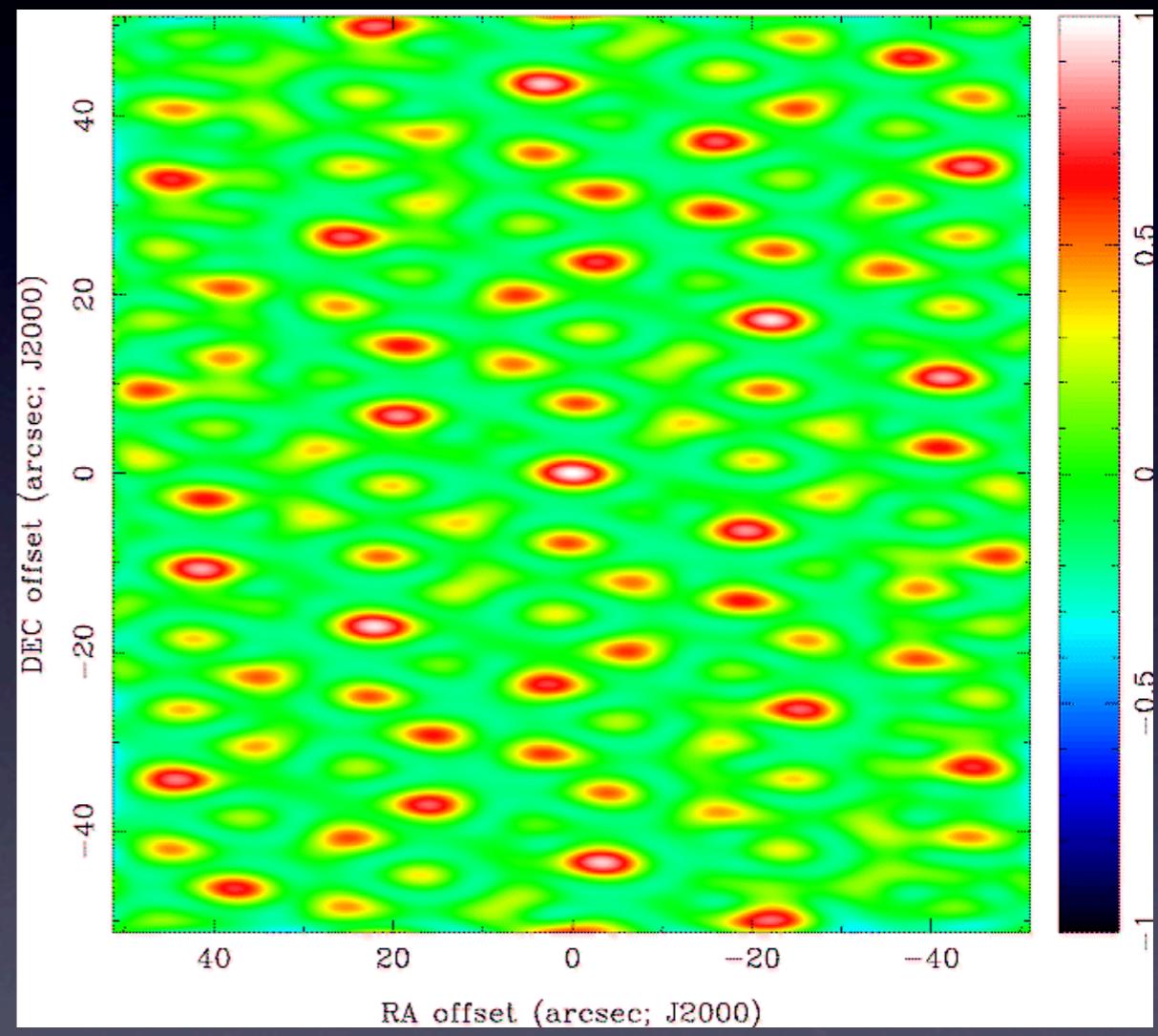
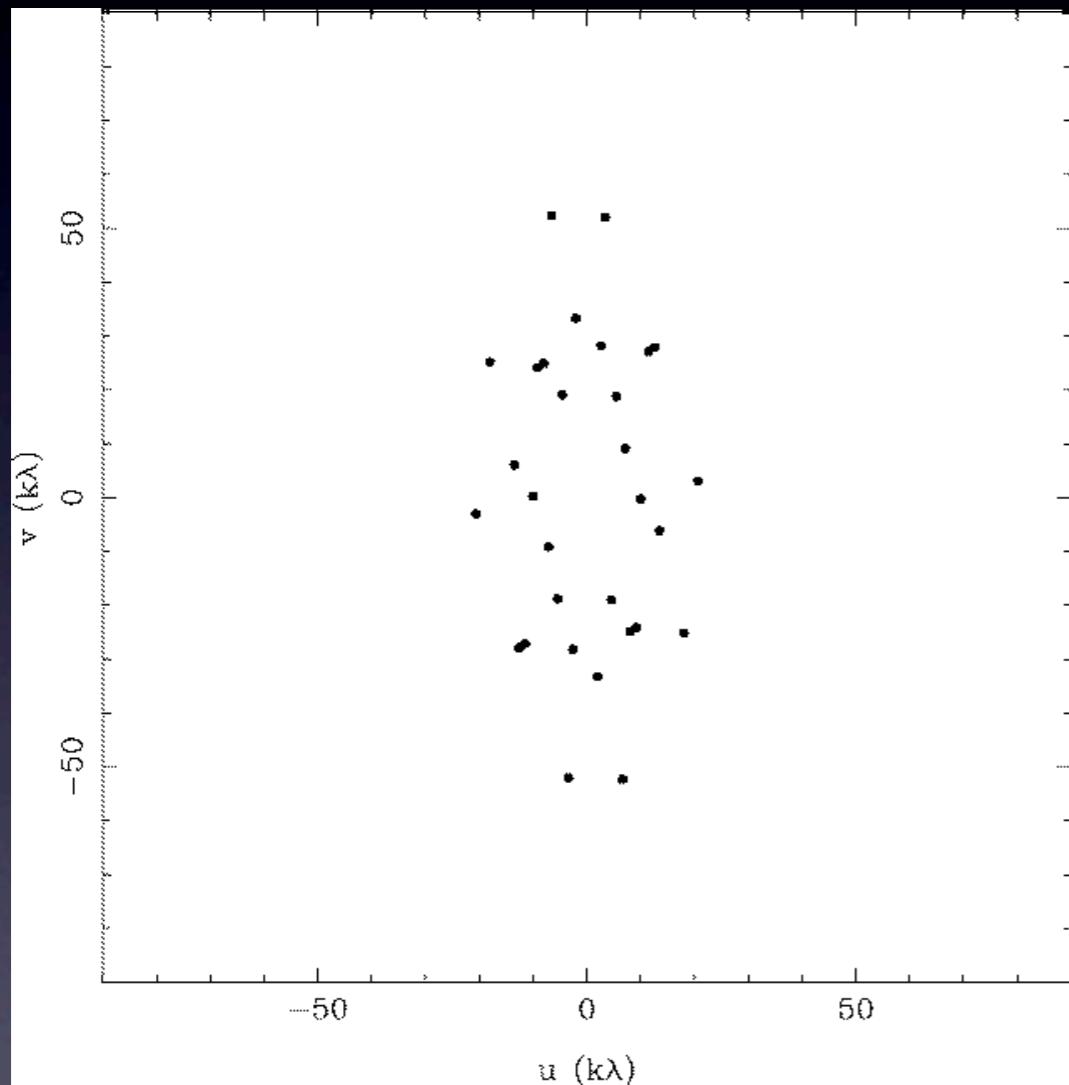
u-v coverage: Dirty beam

5 antennas



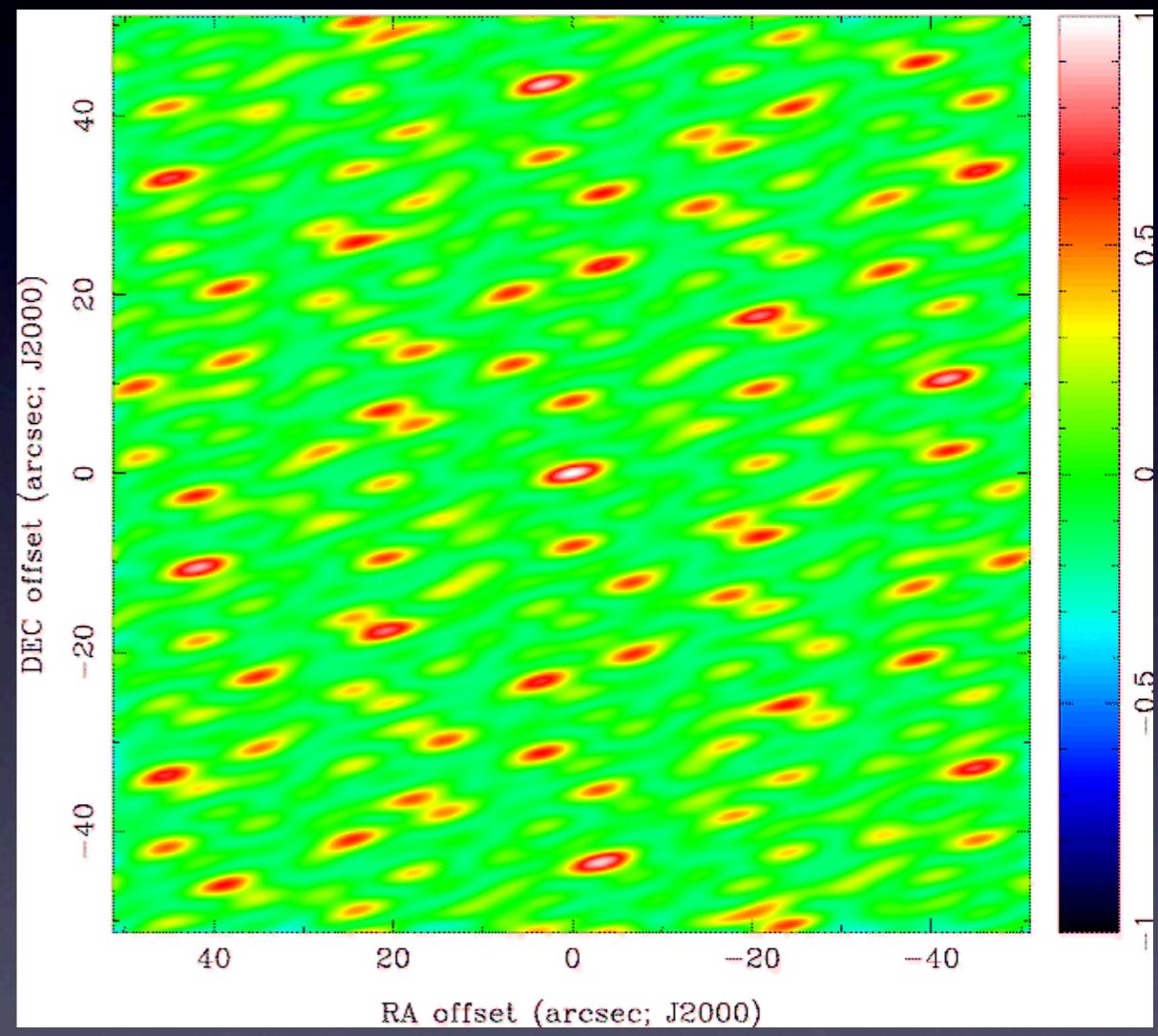
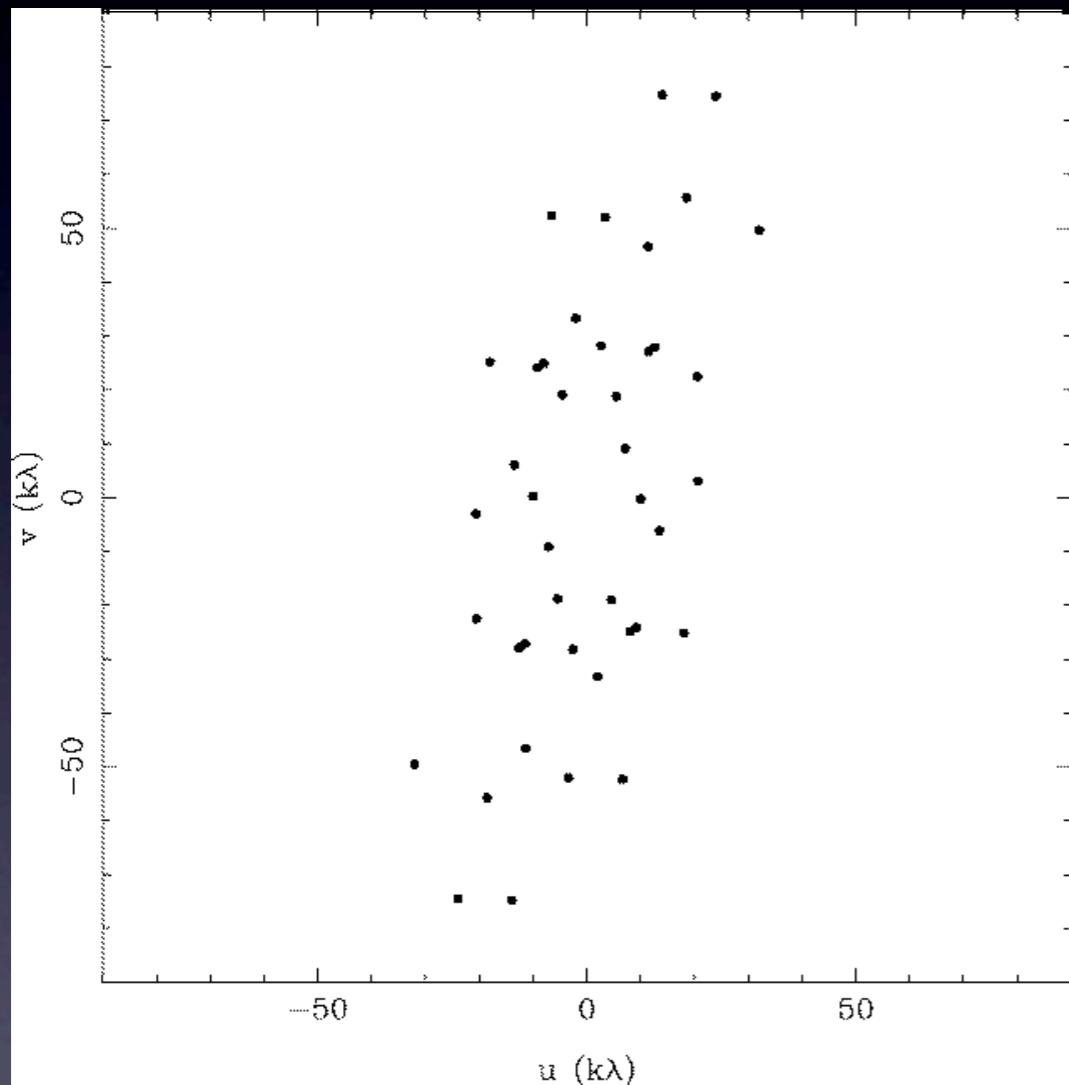
u-v coverage: Dirty beam

6 antennas



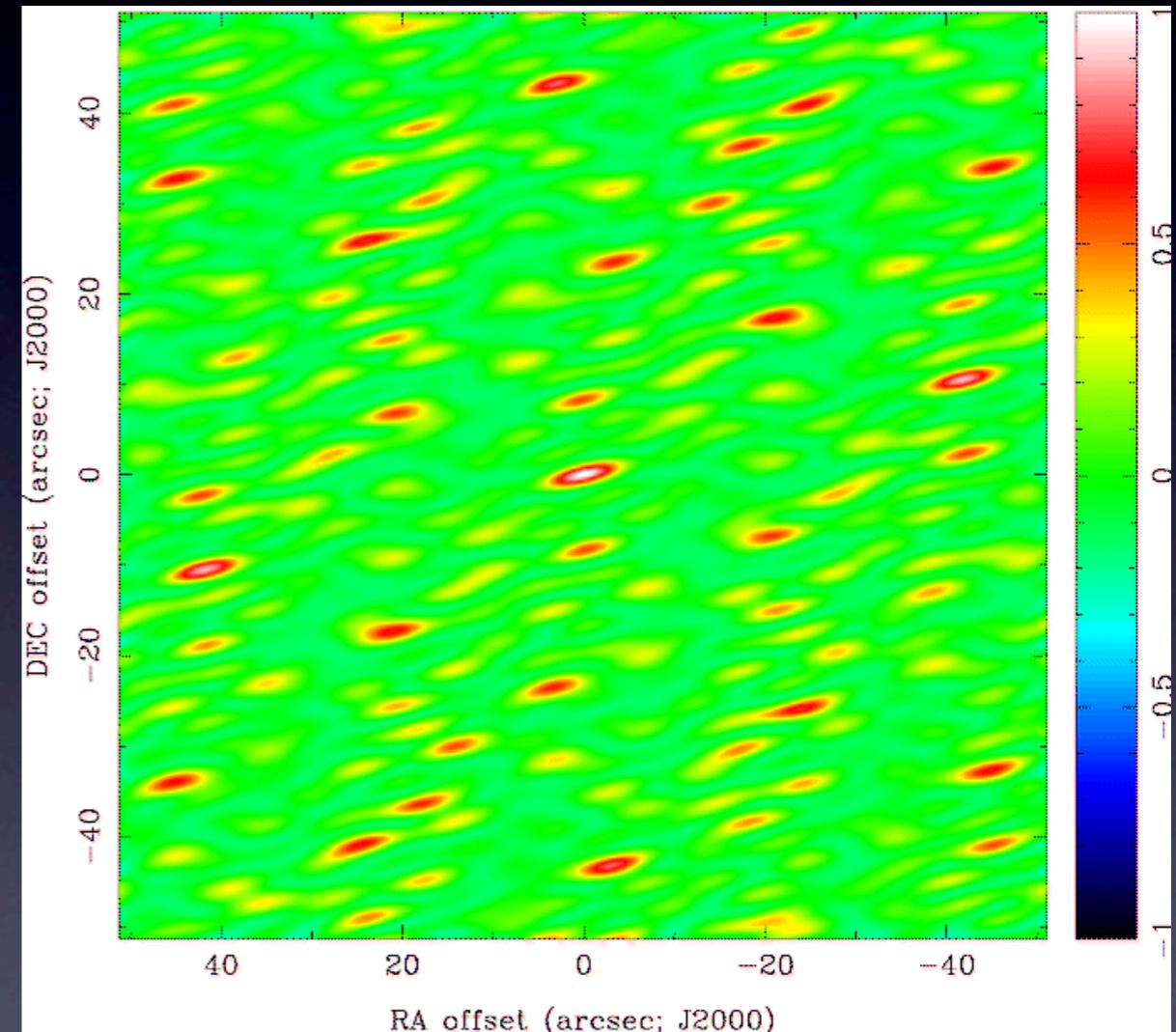
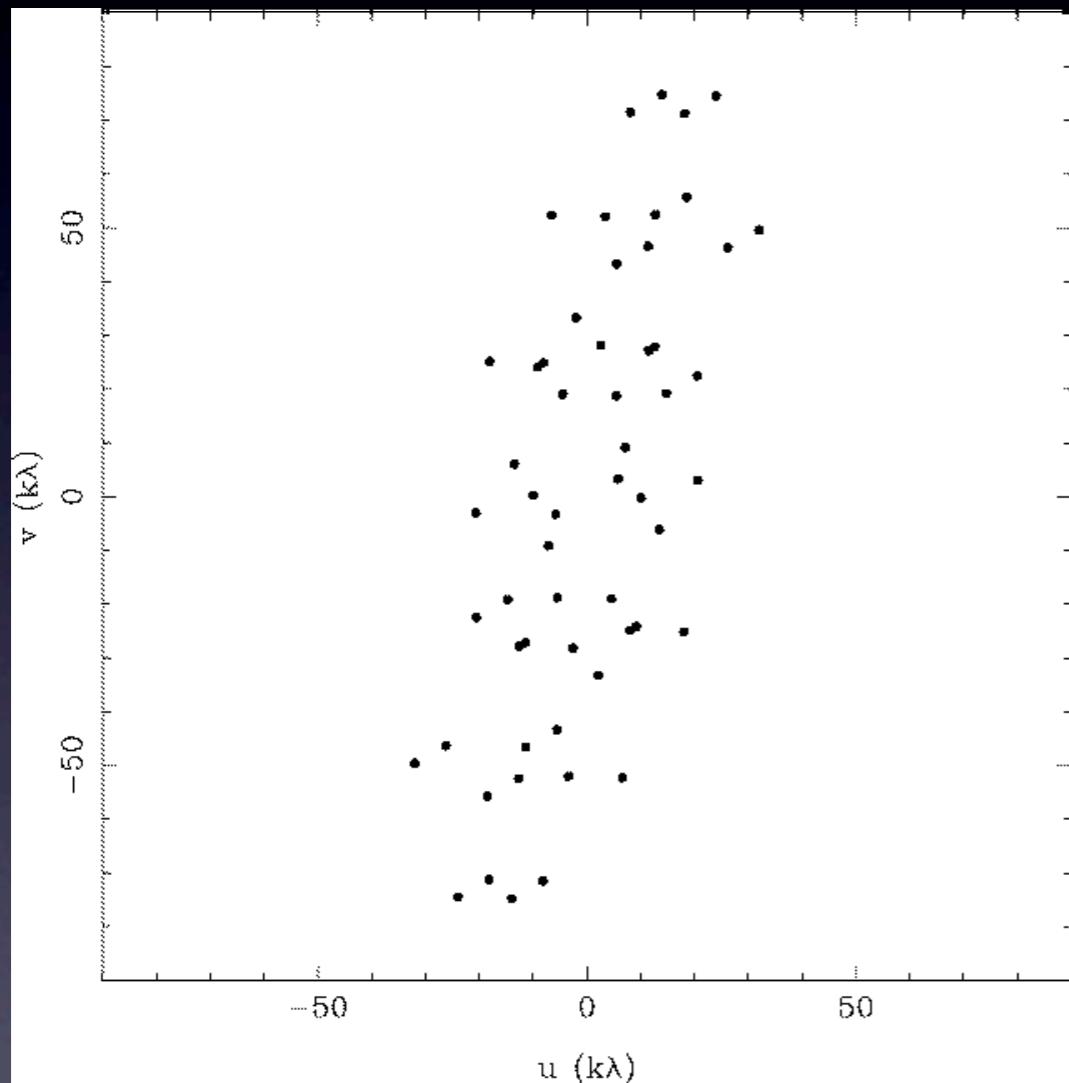
u-v coverage: Dirty beam

7 antennas



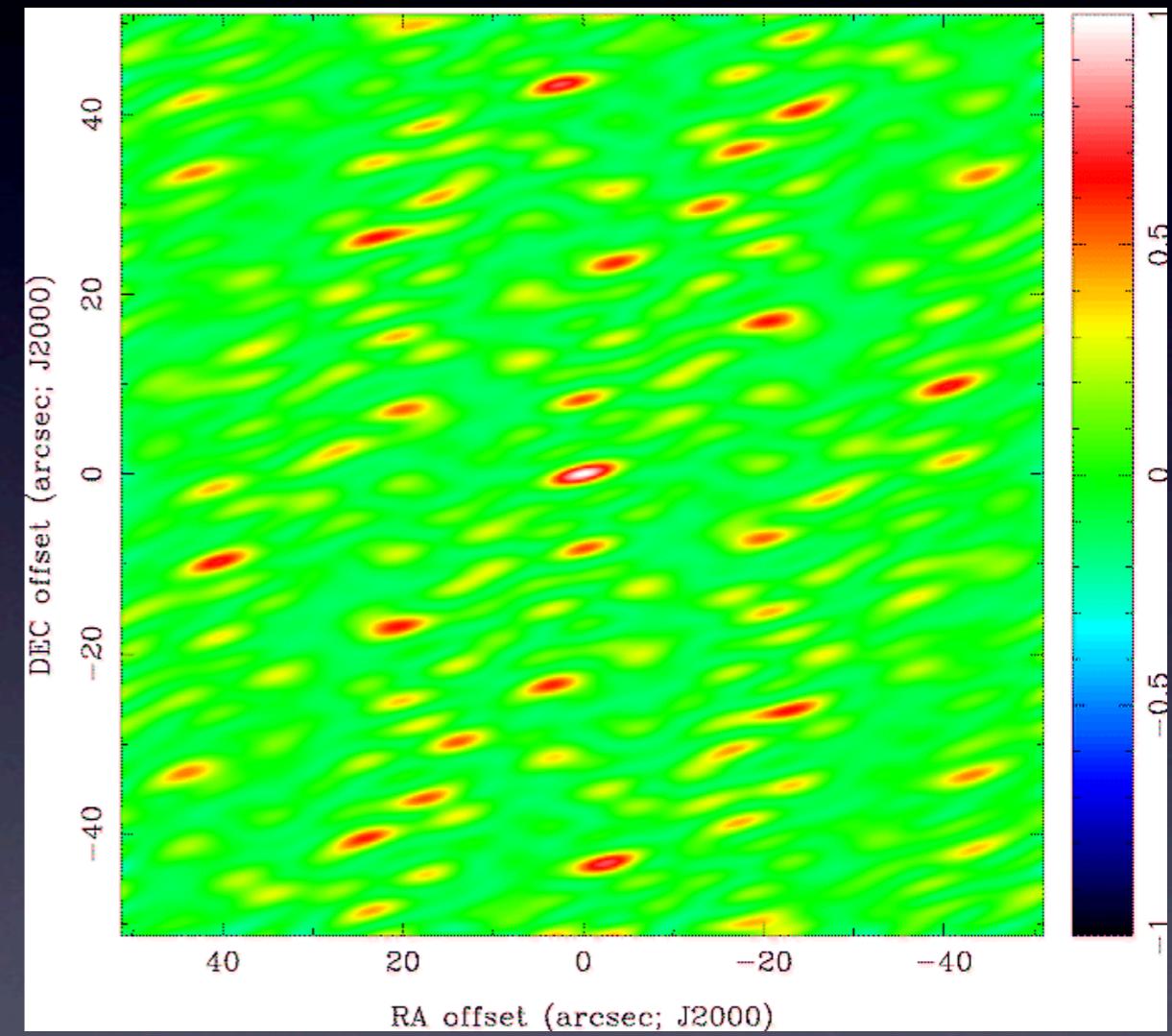
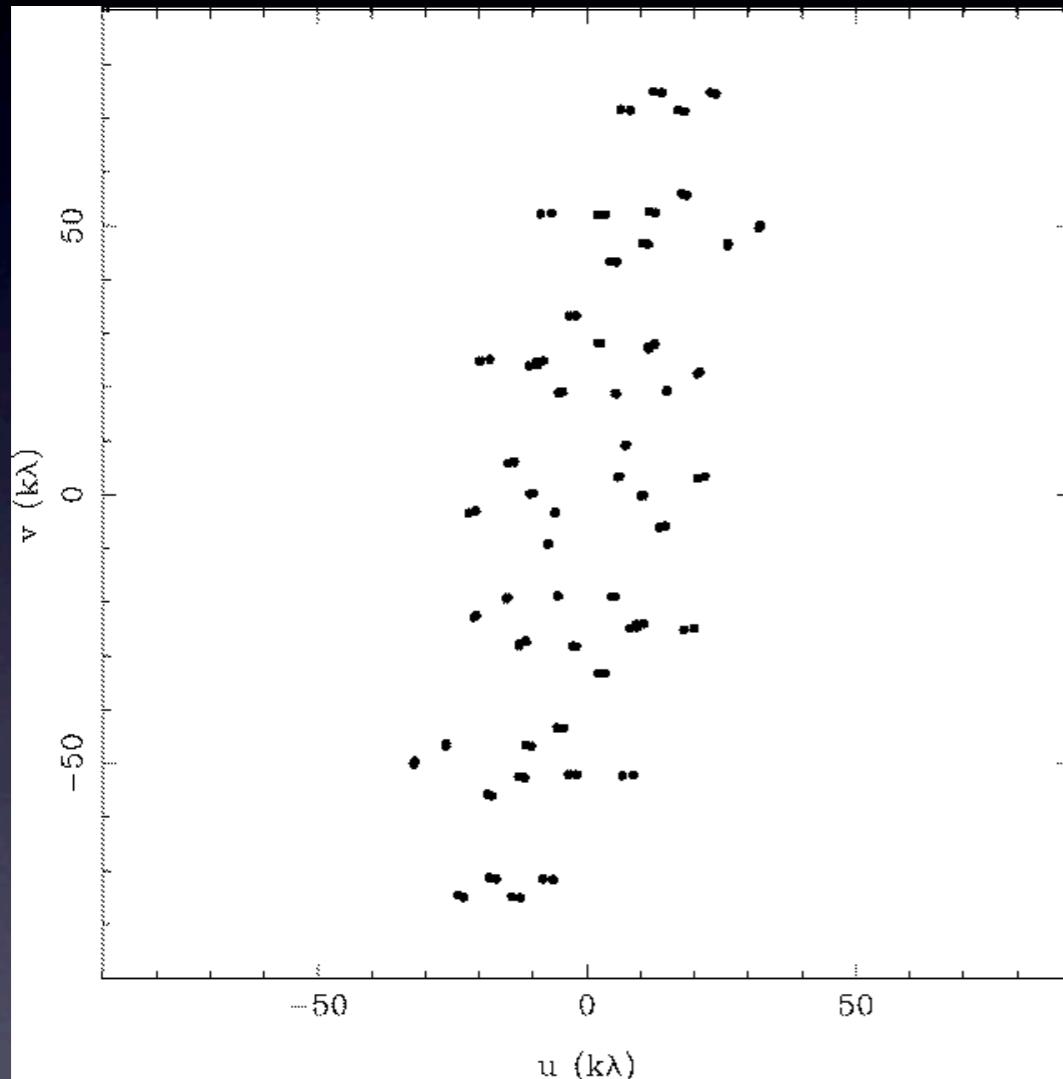
u-v coverage: Dirty beam

8 antennas



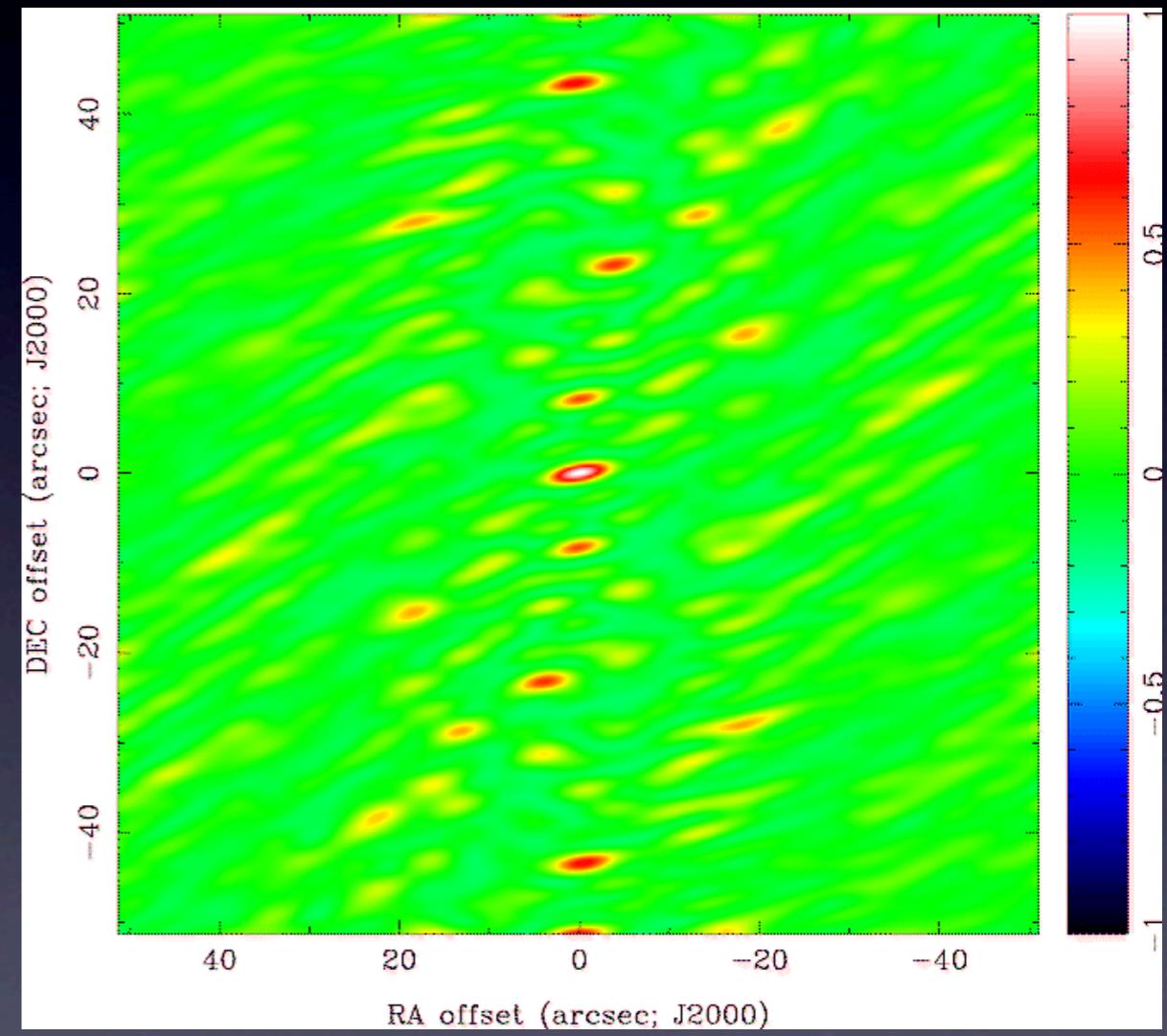
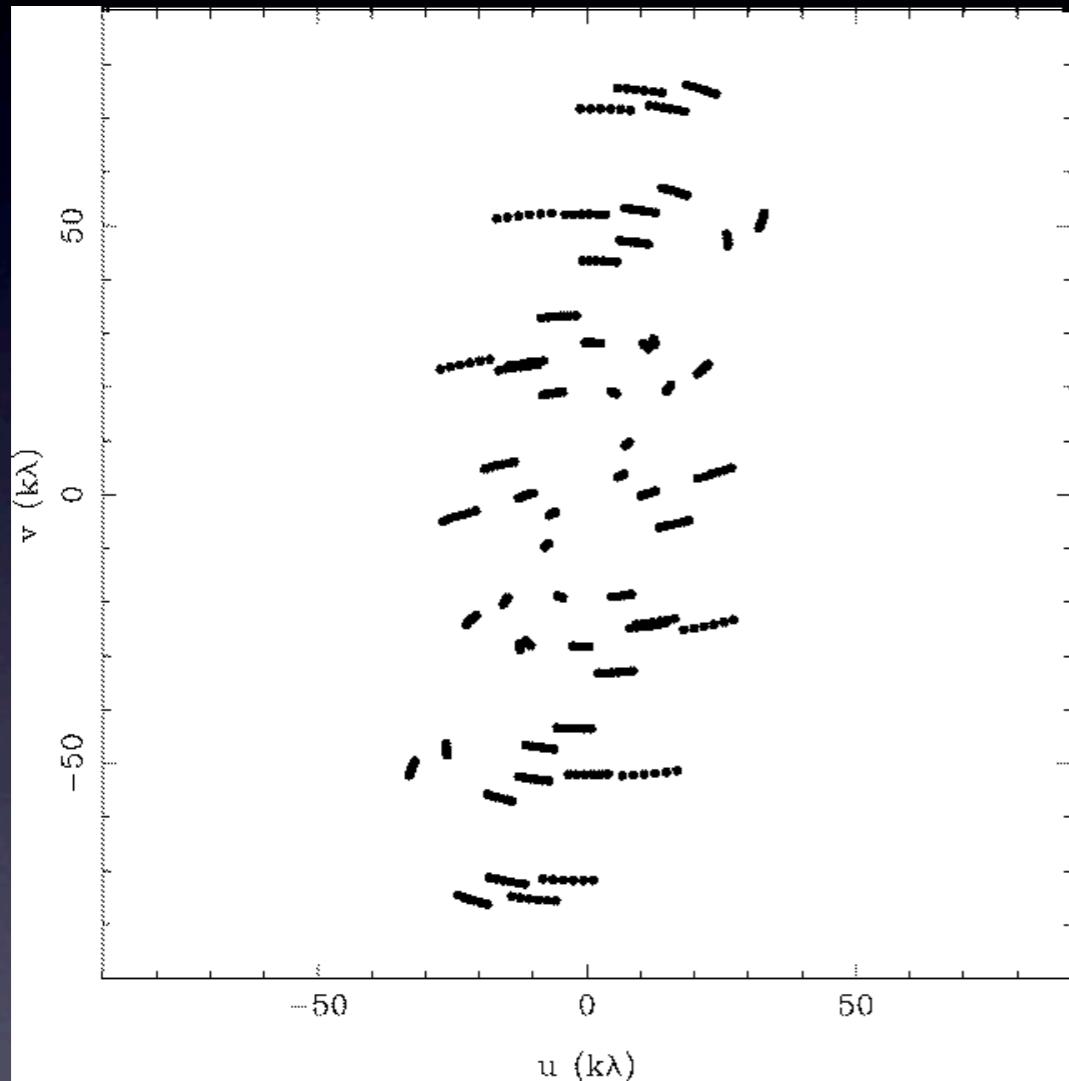
u-v coverage: Dirty beam

8 antennas x 2 samples



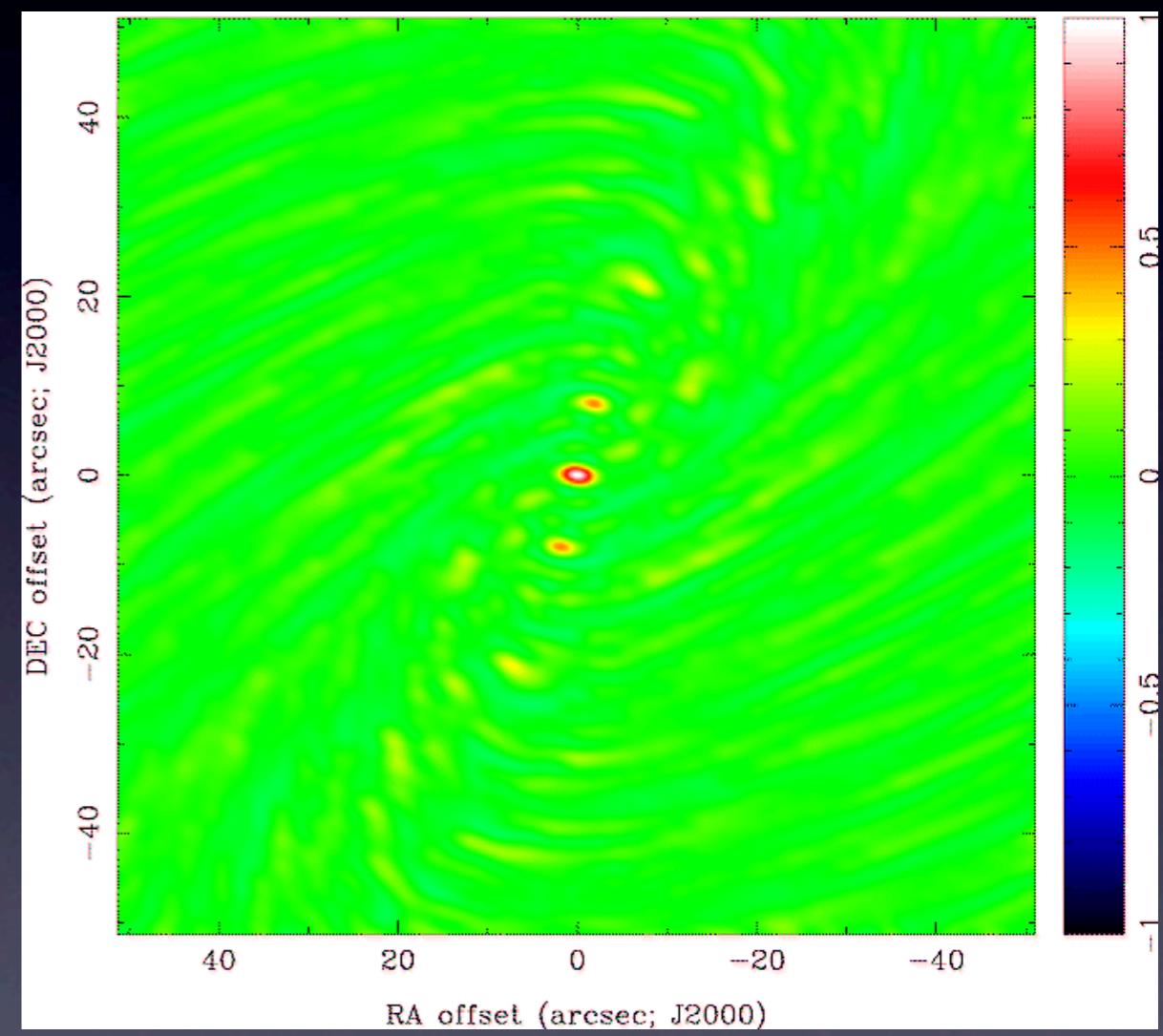
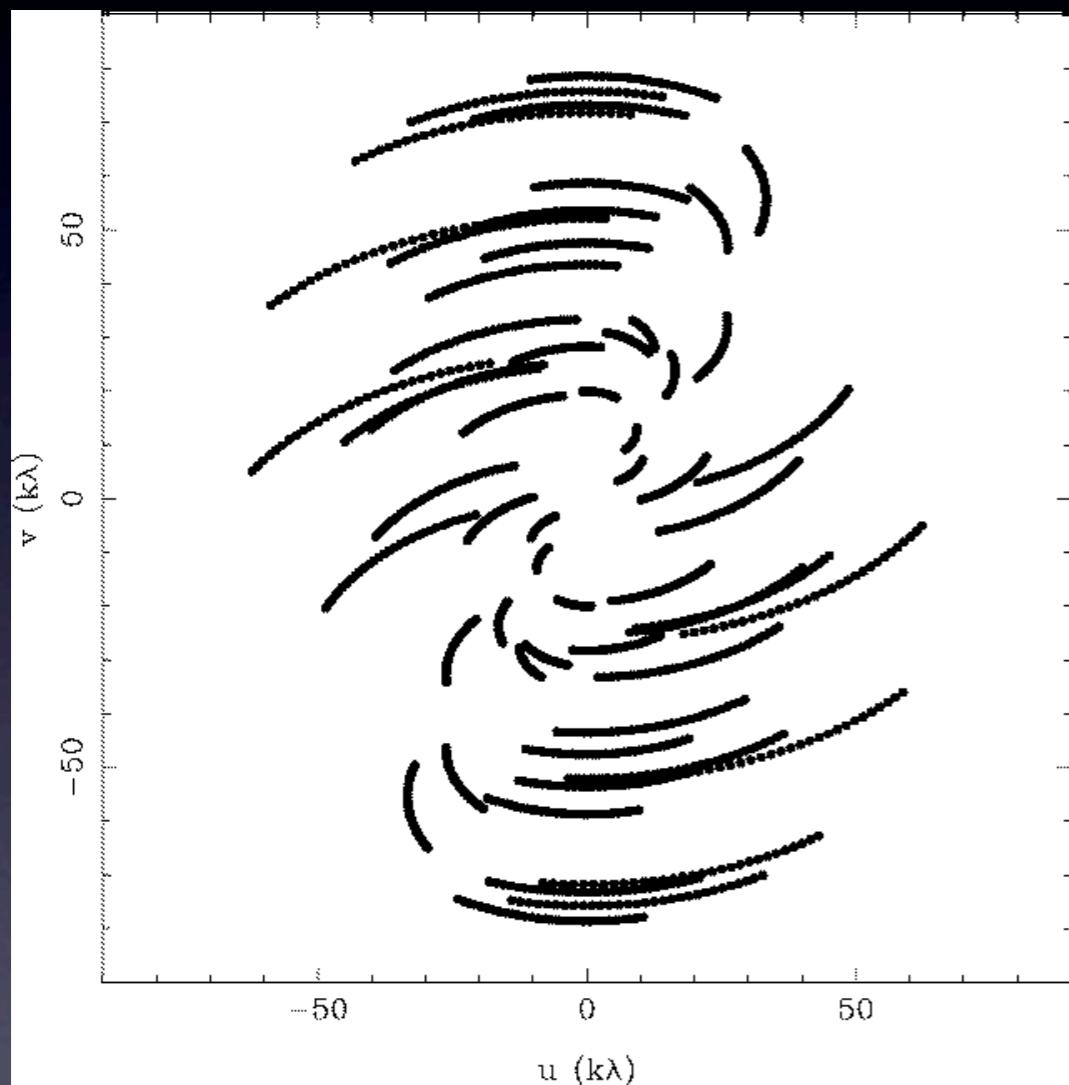
u-v coverage: Dirty beam

8 antennas \times 6 samples



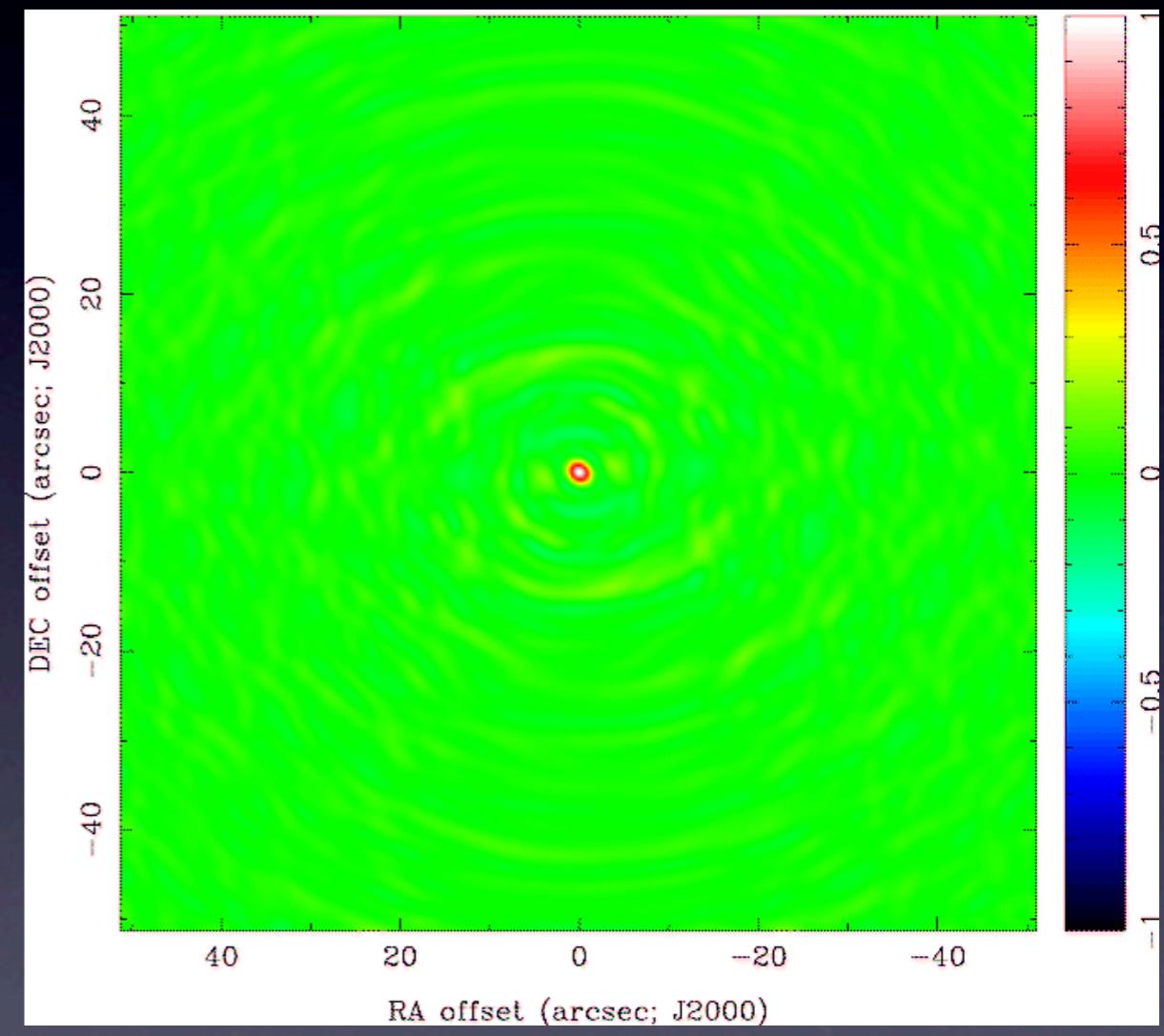
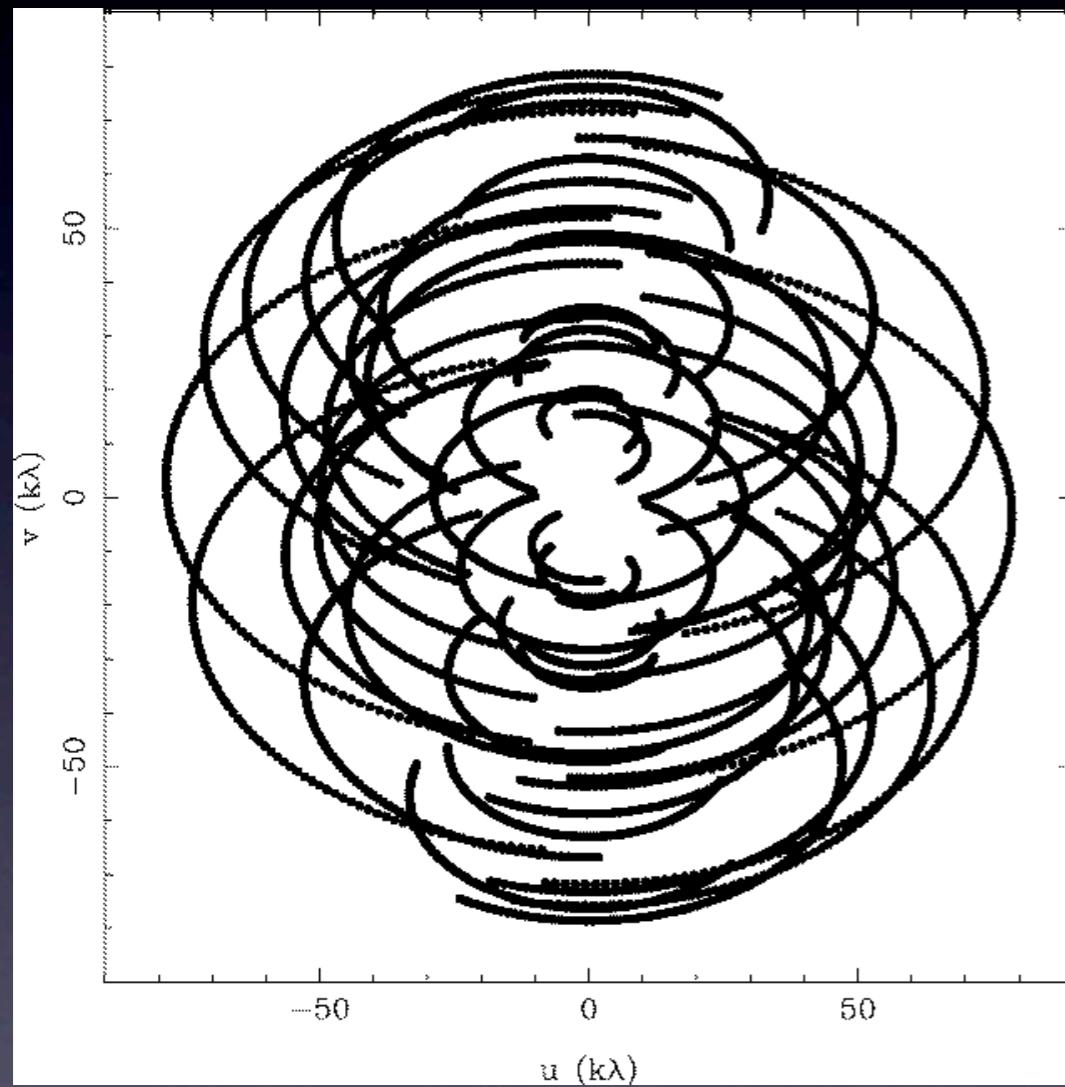
u-v coverage: Dirty beam

8 antennas \times 30 samples



u-v coverage: Dirty beam

8 antennas \times 107 samples



Observations

Main ideas

- Flux calibrator.
- Phase calibrators.
- Best u-v coverage.
- Real time VS Tapes.

Interferometers

Interferometers



EVLA

Interferometers



EVLA



eMERLIN

Interferometers



EVLA



eMERLIN



Plateau de Bure

Interferometers



ALMA

Interferometers



ALMA



LOFAR

Interferometers



ALMA

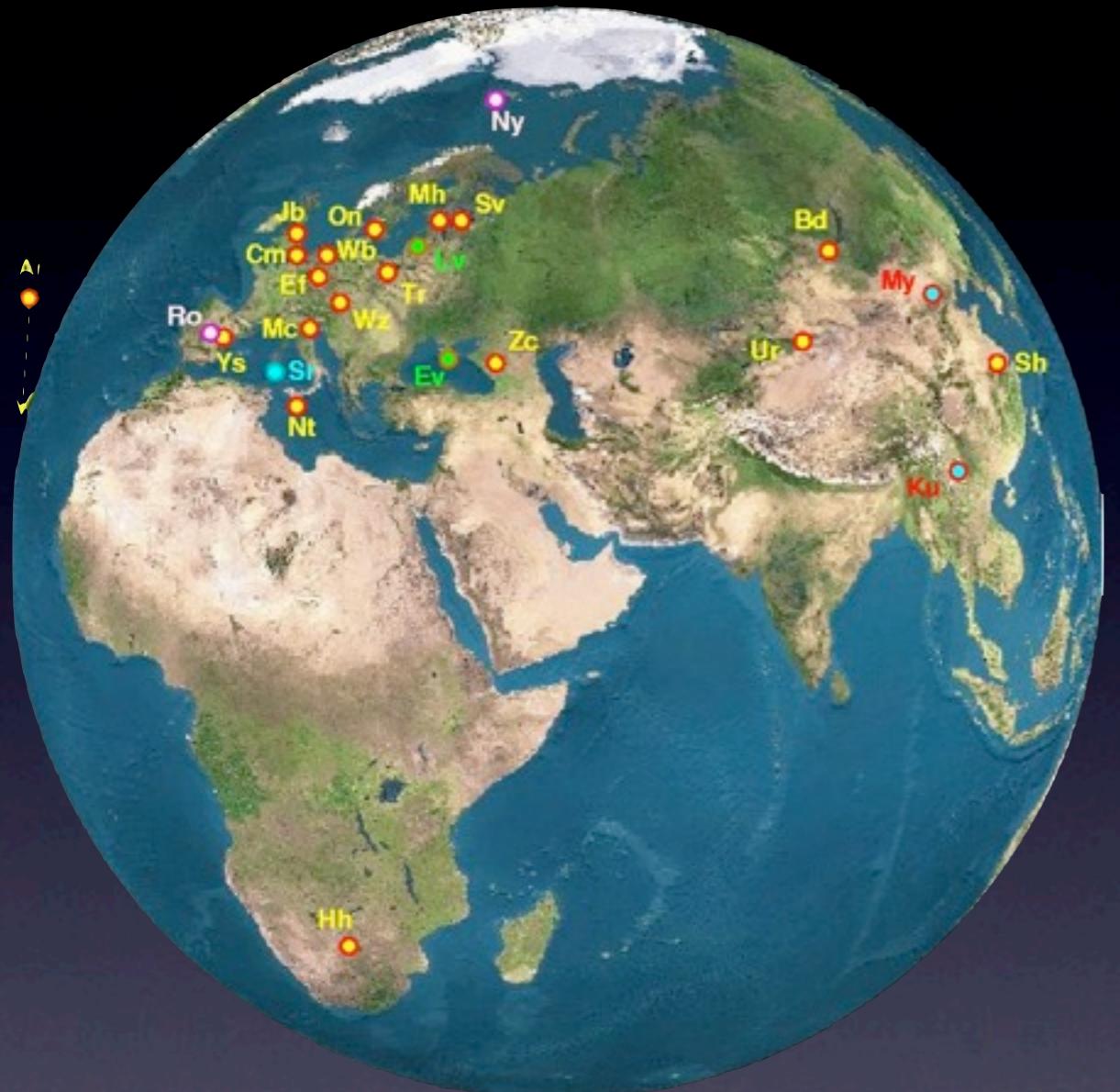


SKA

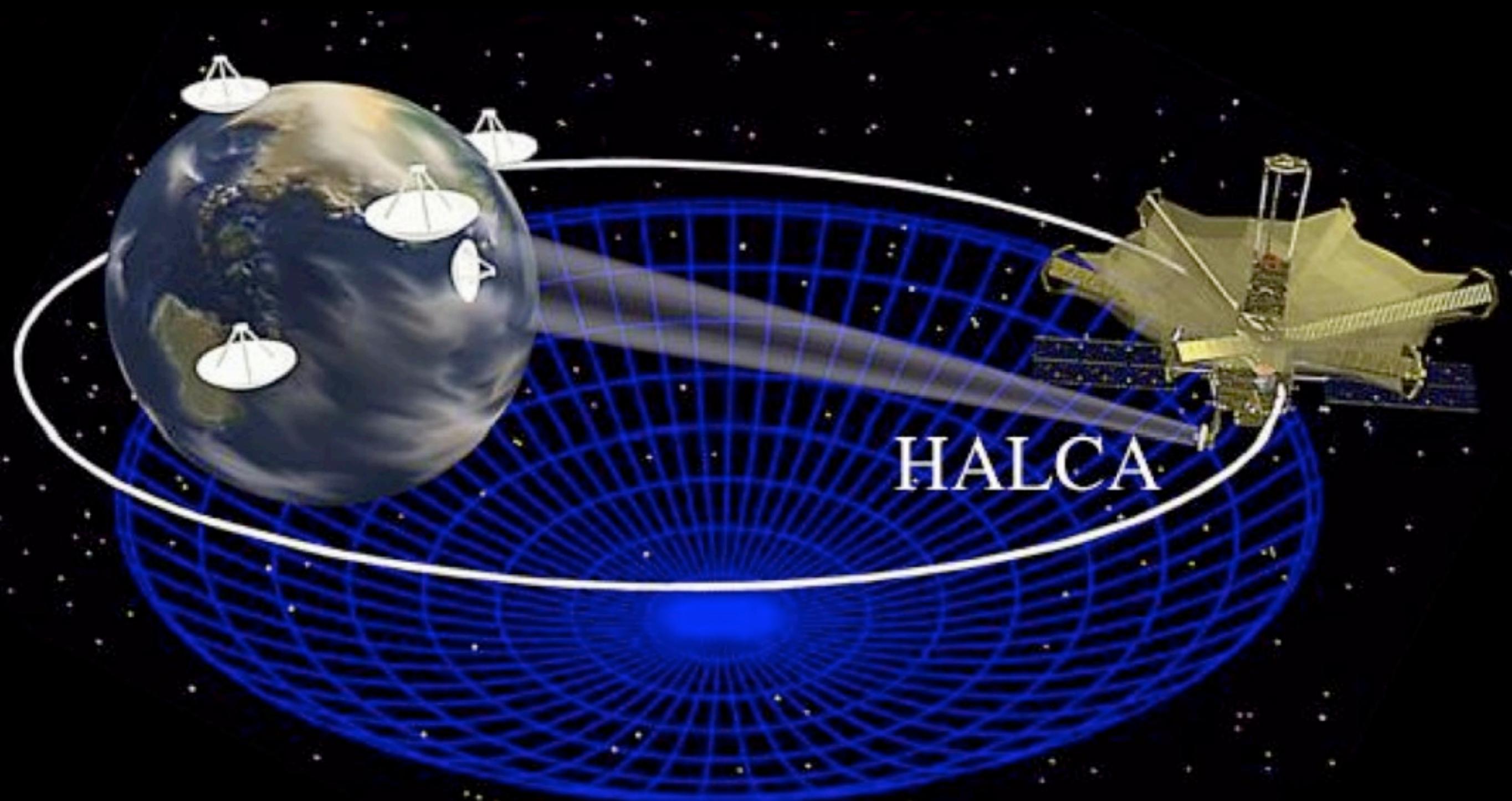


LOFAR

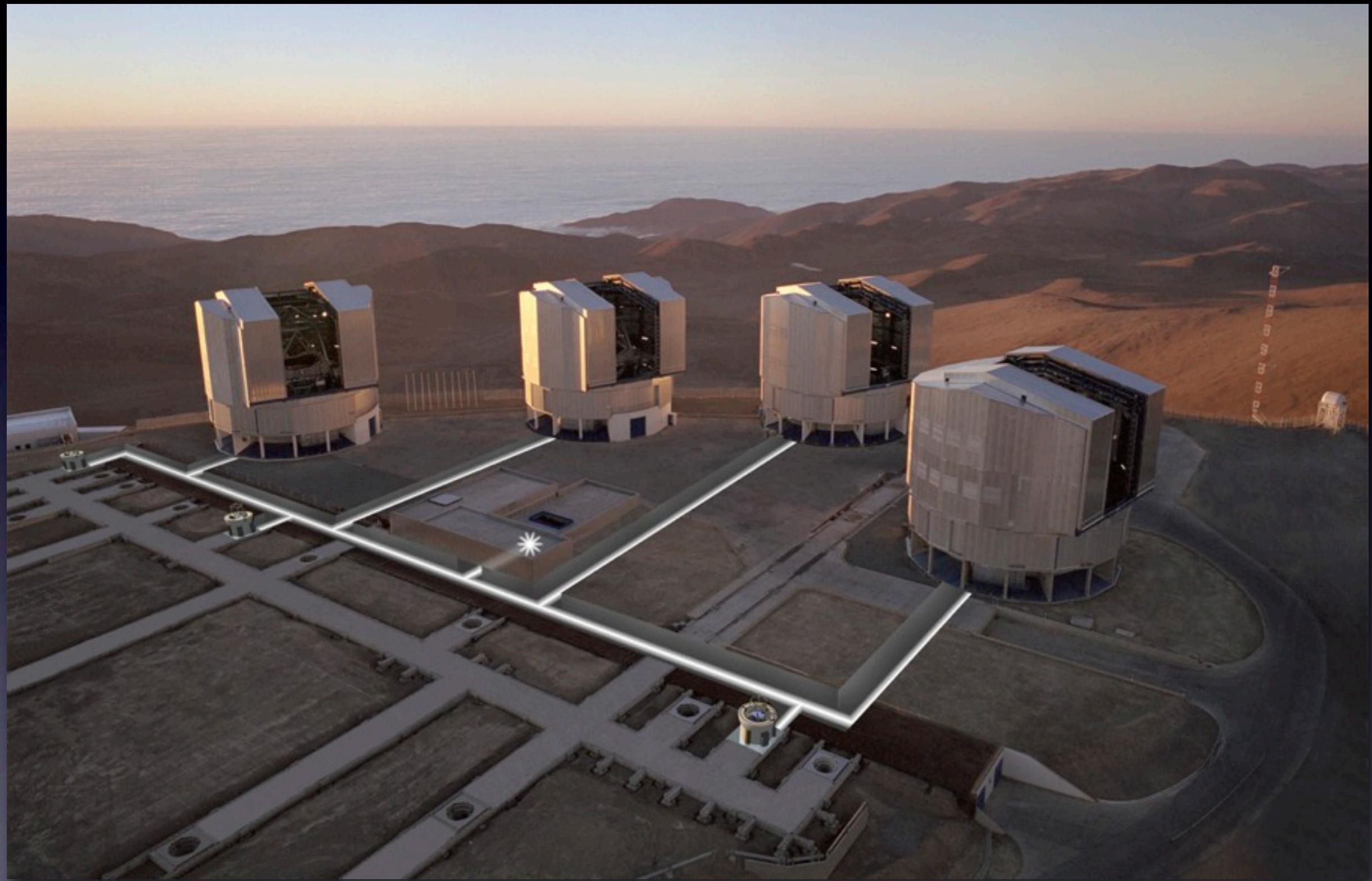
VLBI



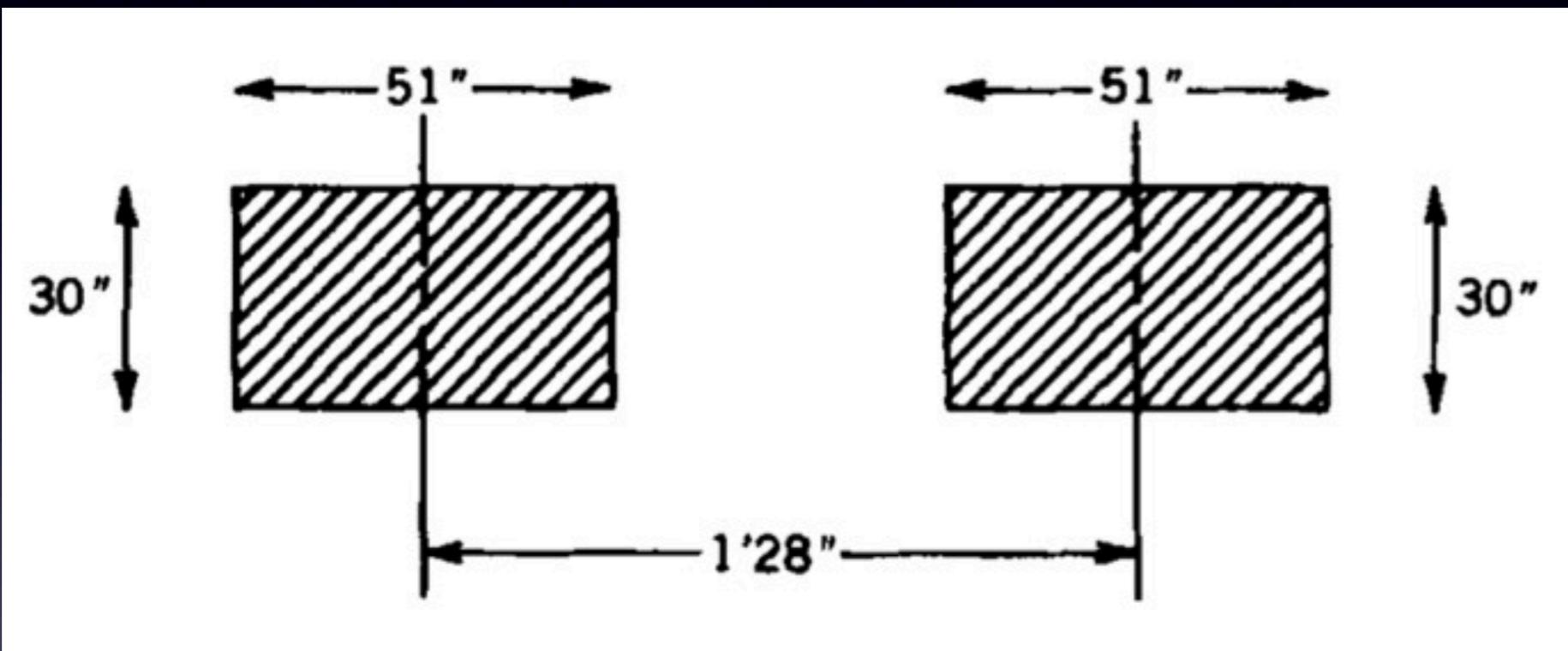
Space - VLBI



IR and beyond...

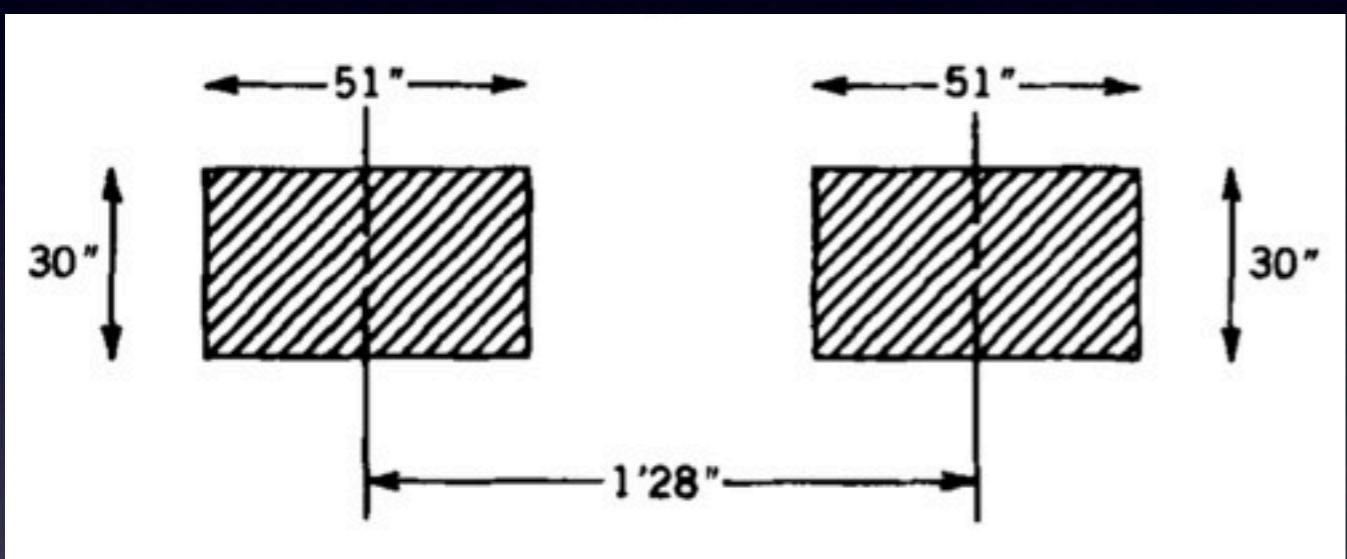


Cygnus A

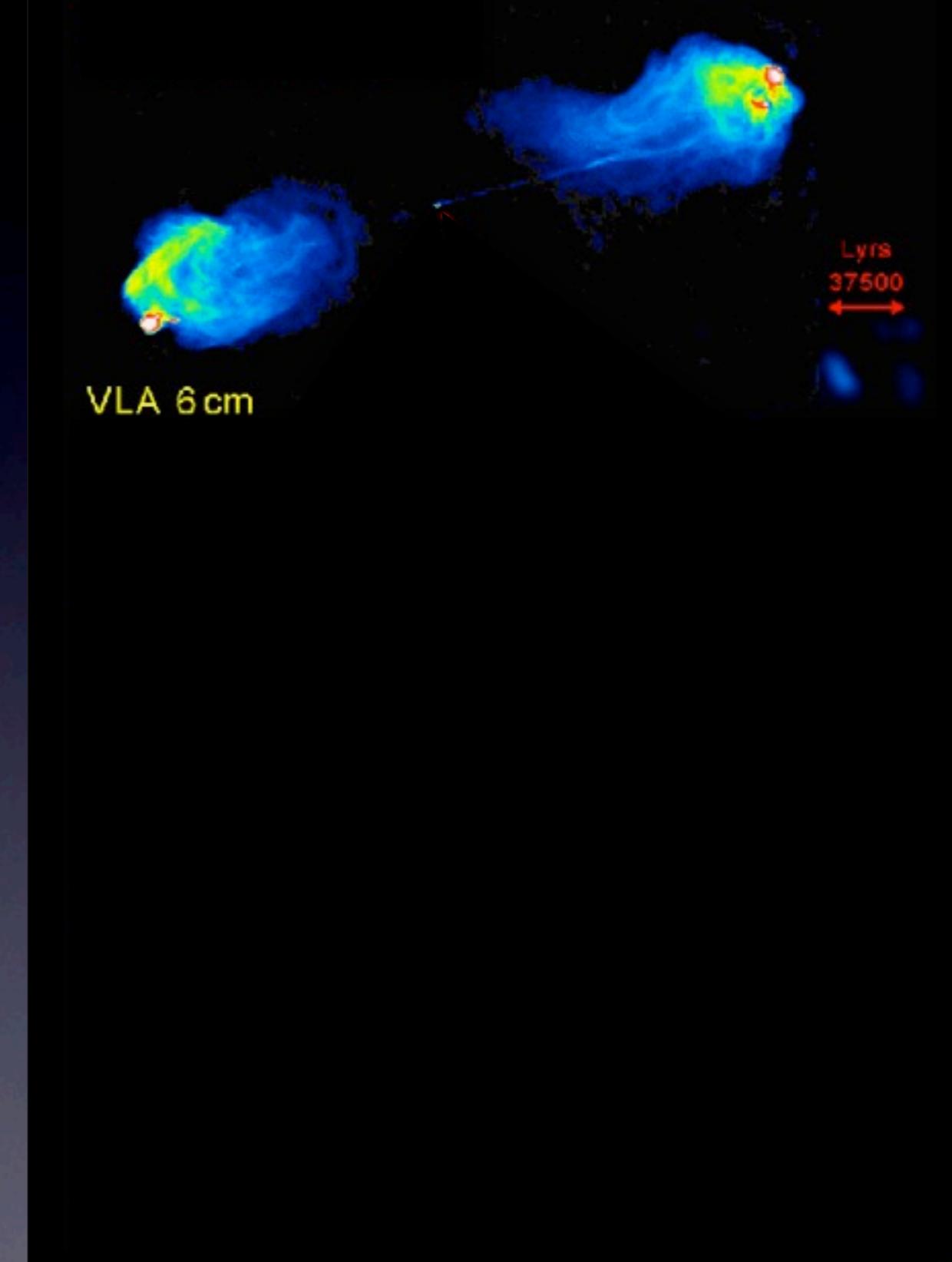


Jennison & Das Gupta (1953)

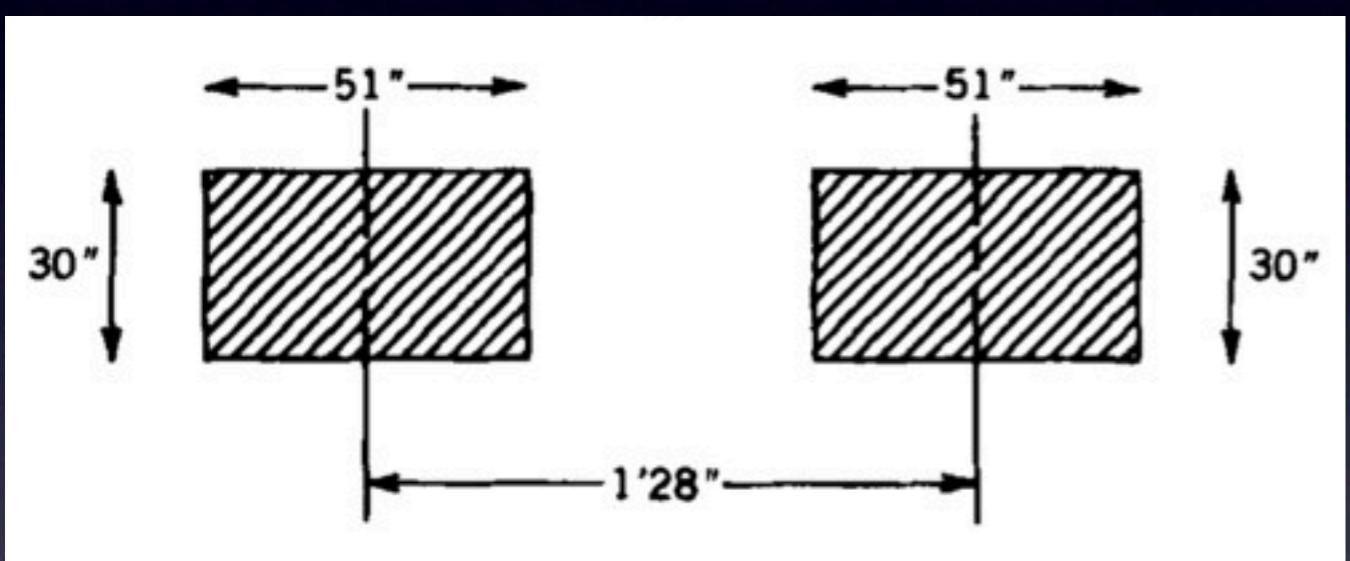
Cygnus A



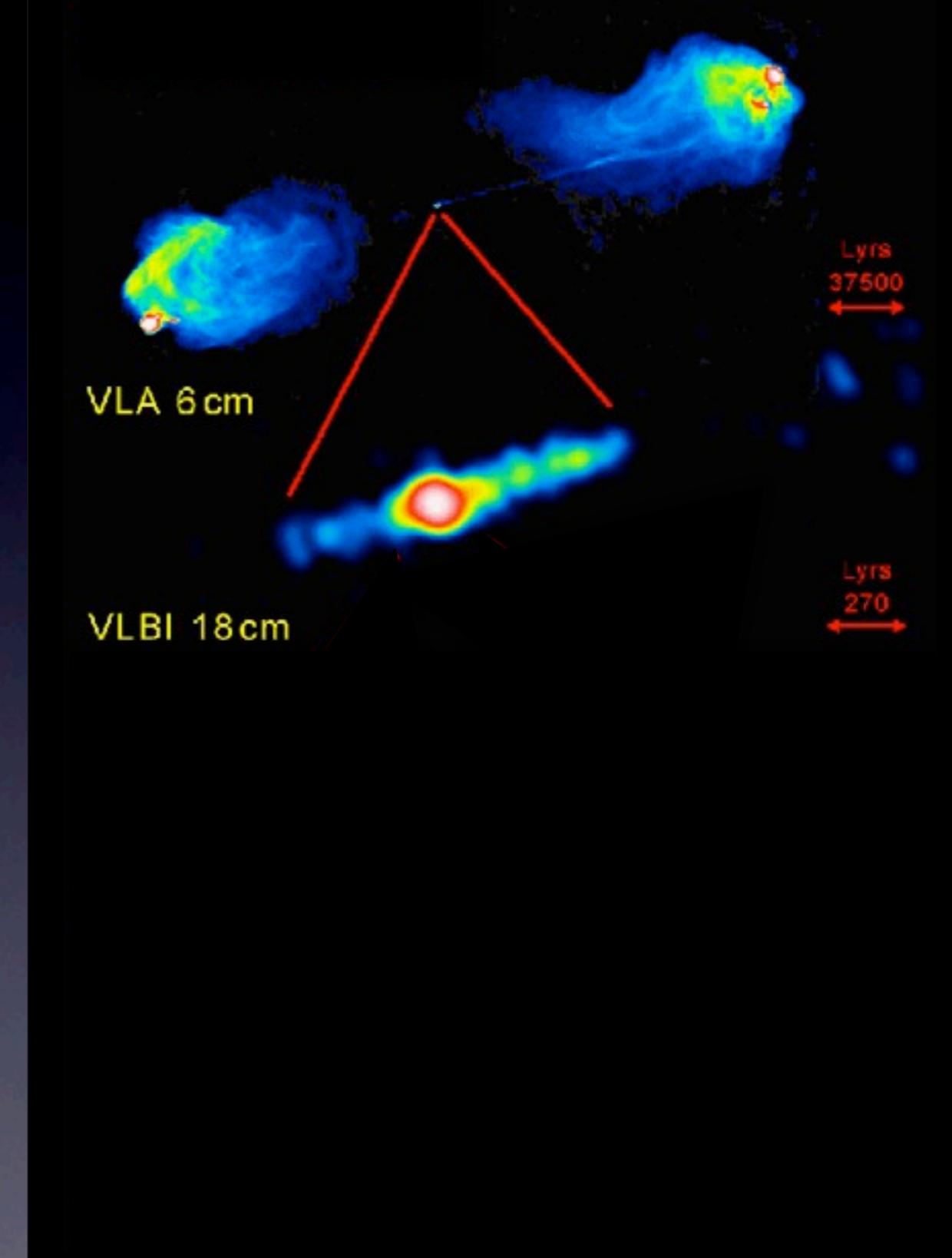
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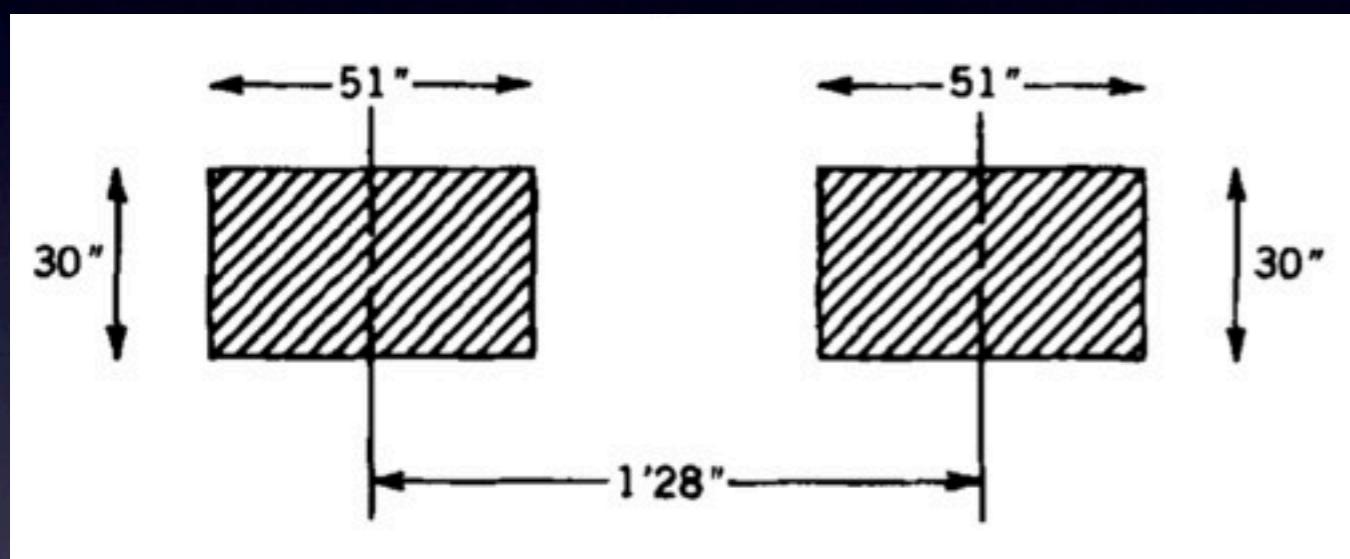
Cygnus A



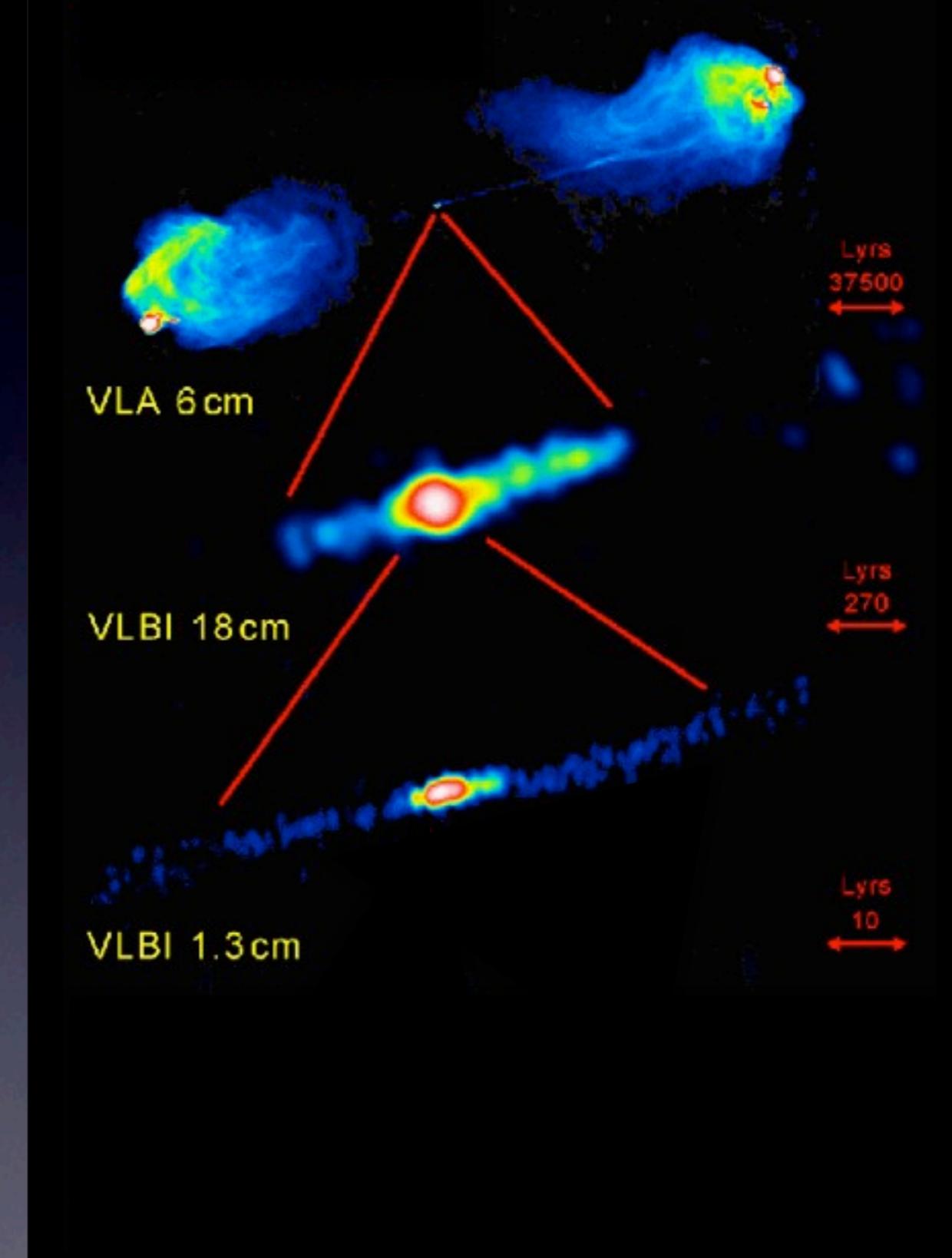
Jennison & Das Gupta (1953)



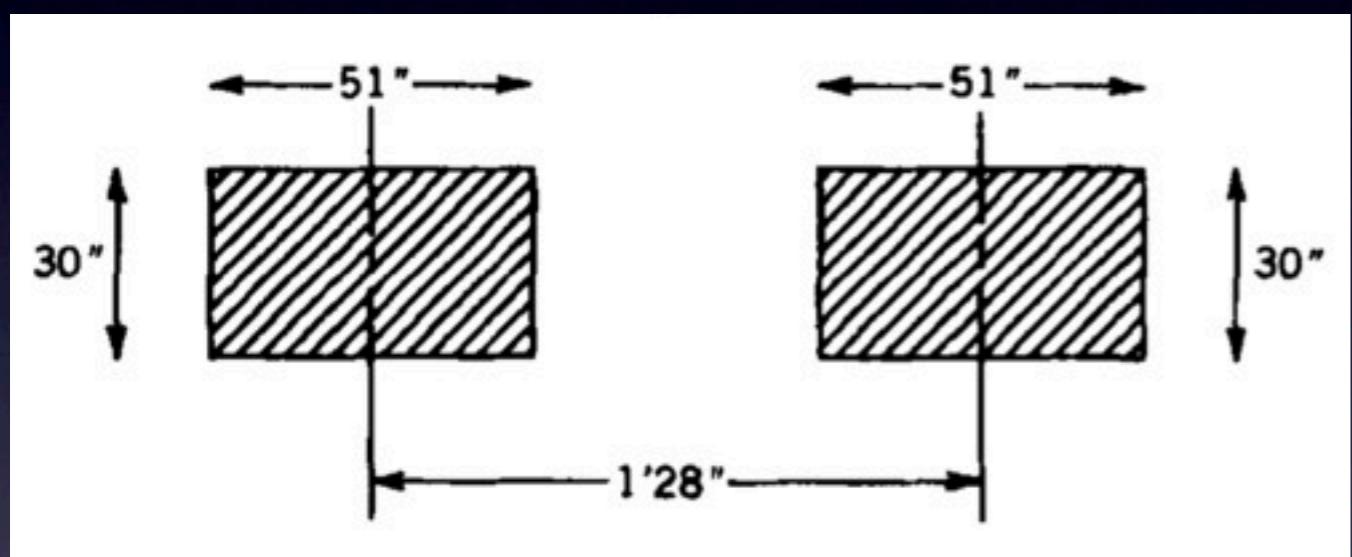
Cygnus A



Jennison & Das Gupta (1953)

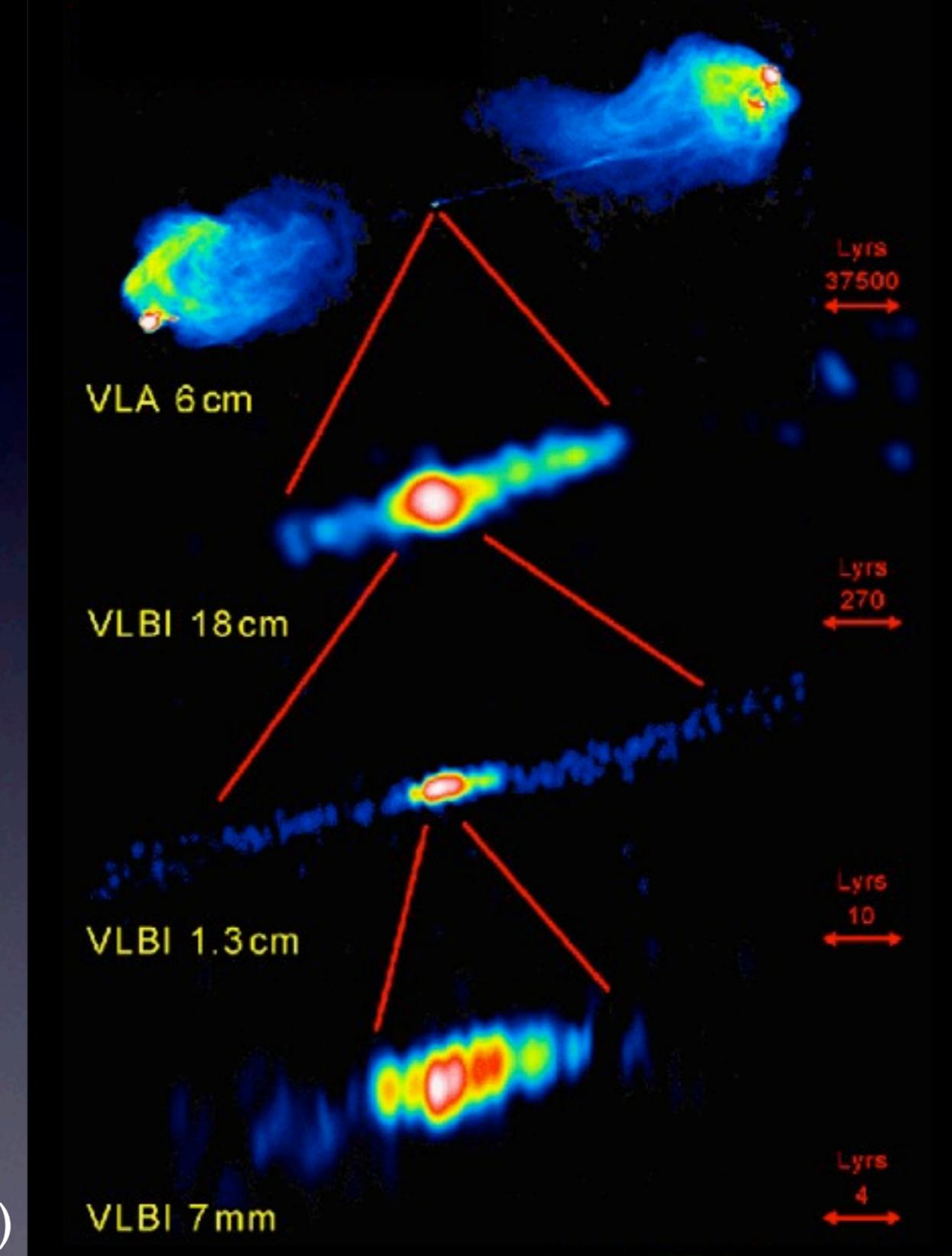


Cygnus A



Jennison & Das Gupta (1953)

Krichbaum et al. (1998)



Data reduction

Software

AIPS



Most powerful software.

IRAF ‘like’.

Created in 1970s (ancient FORTRAN).

Software

AIPS



Most powerful software.

IRAF ‘like’.

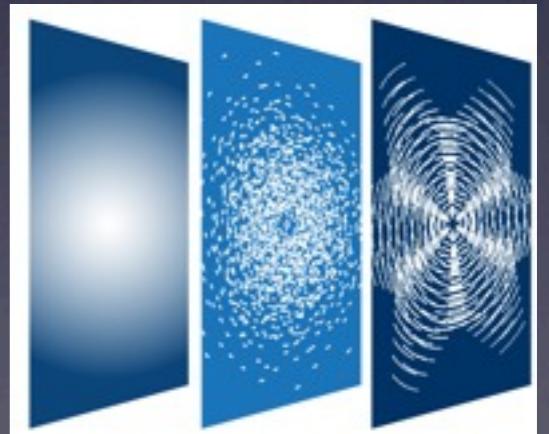
Created in 1970s (ancient FORTRAN).

CASA

Developed for ALMA.

C/C++ kernel / Python interface.

Nearly completed.



Step by step

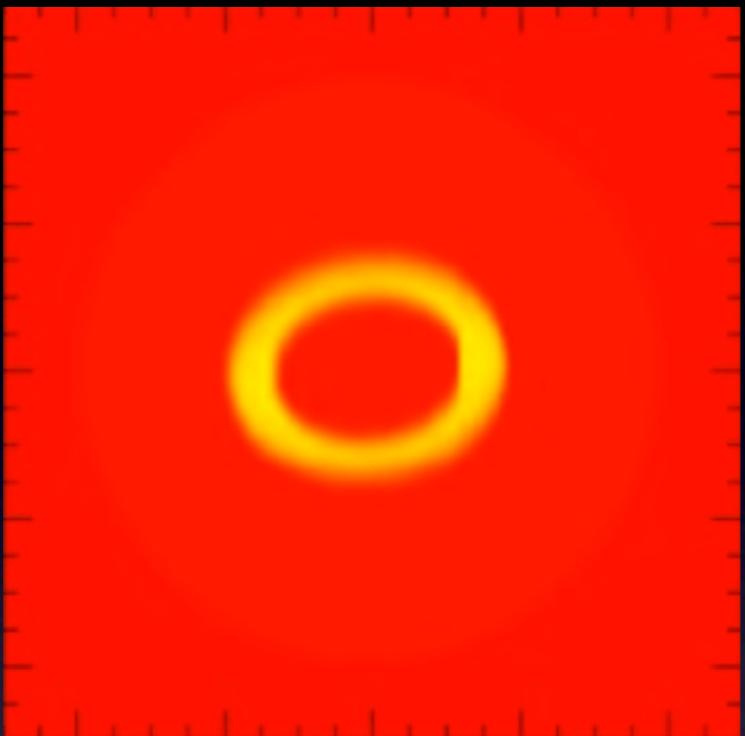
- Data examination.
- Amplitude calibration.
- Phase calibration.
- Imaging.

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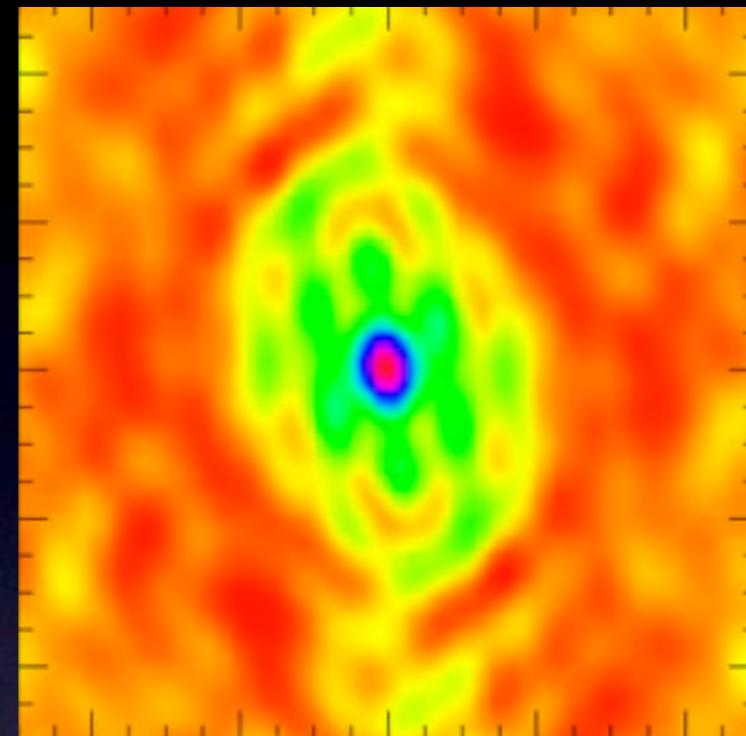
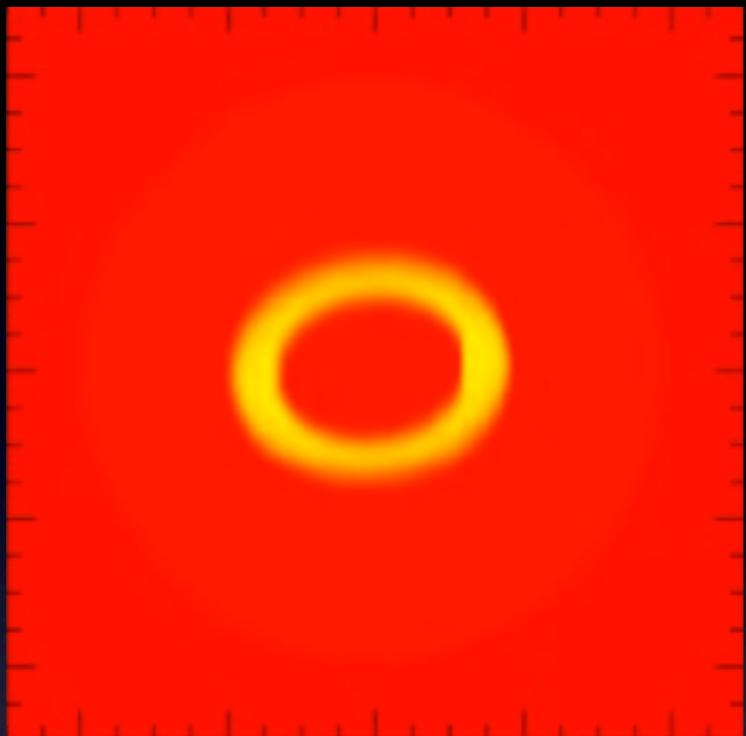
Imaging

Source



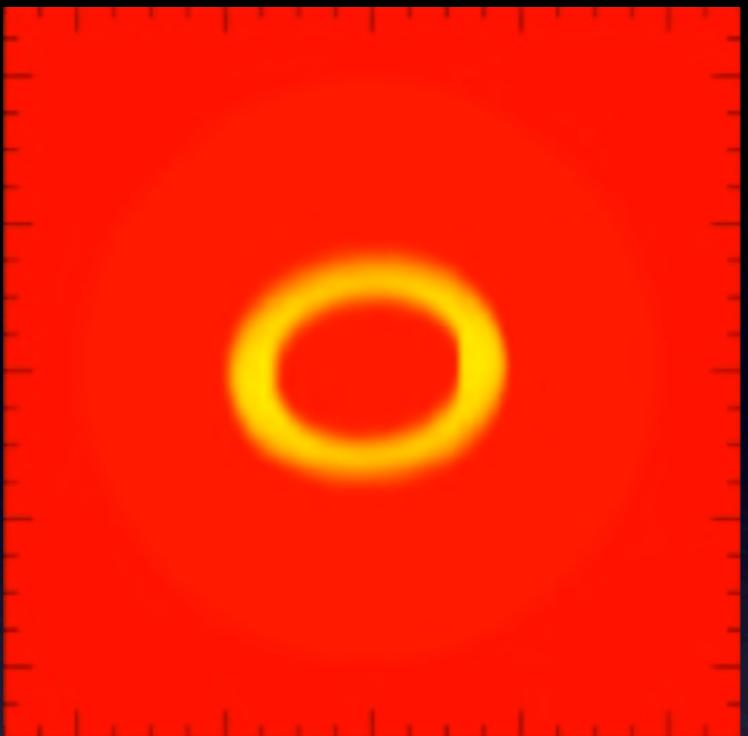
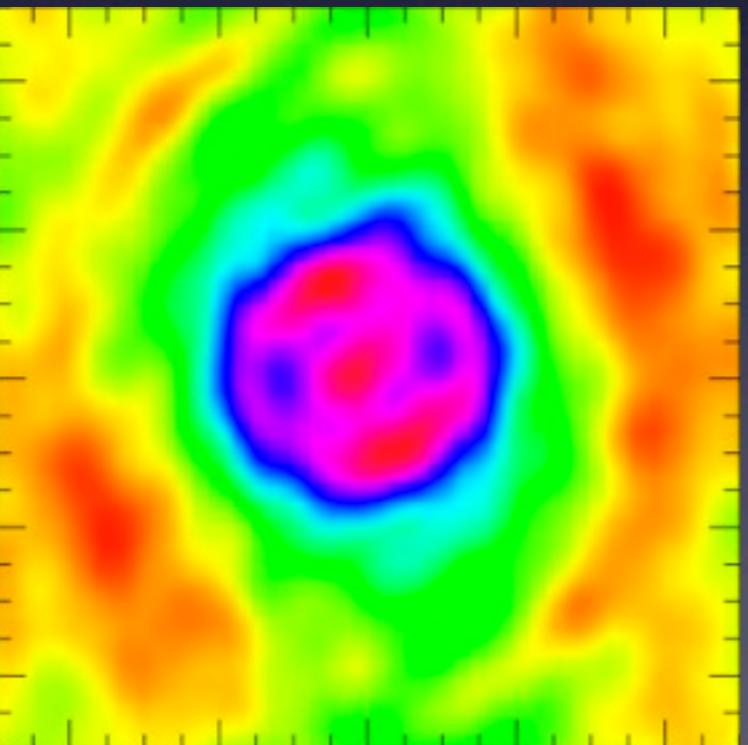
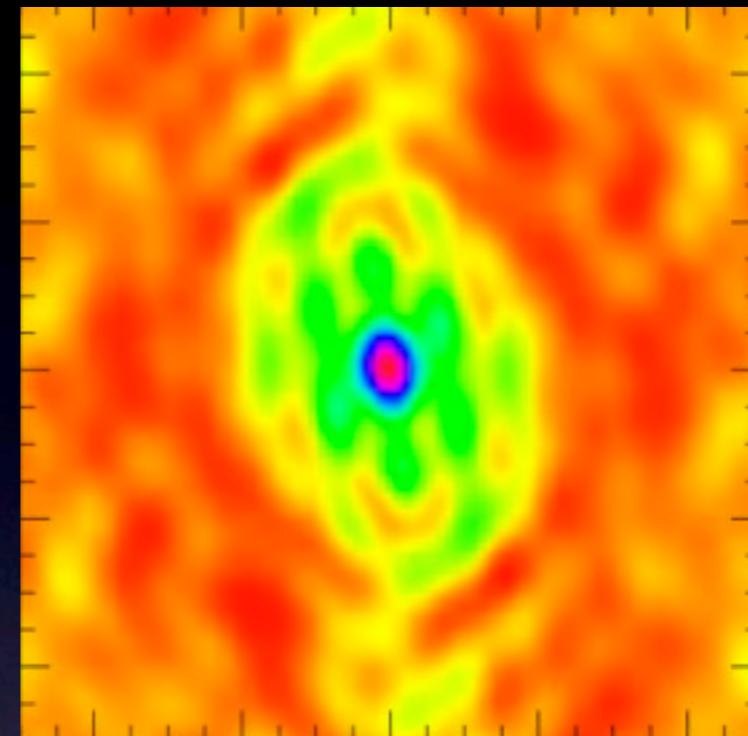
Imaging

Source

Dirty beam
(PSF)

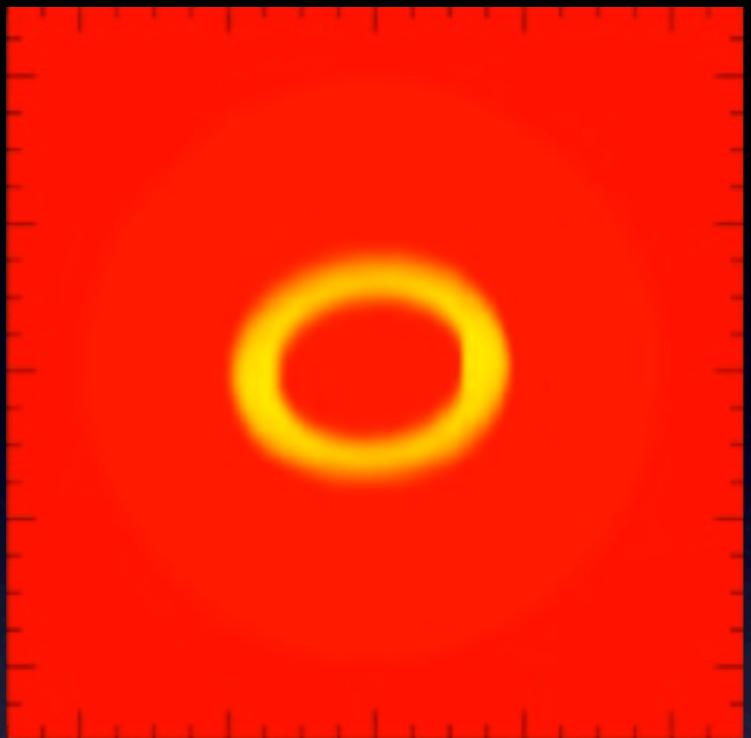
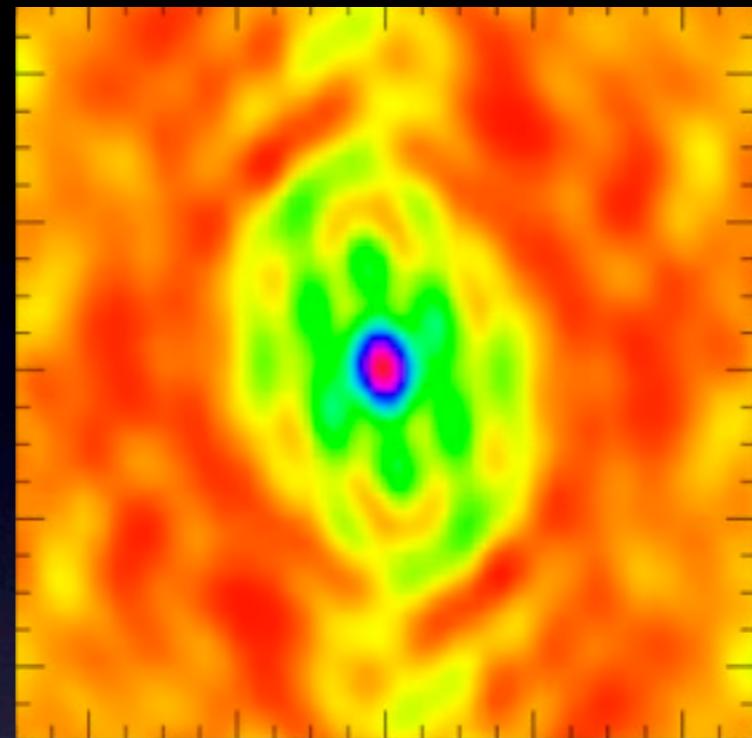
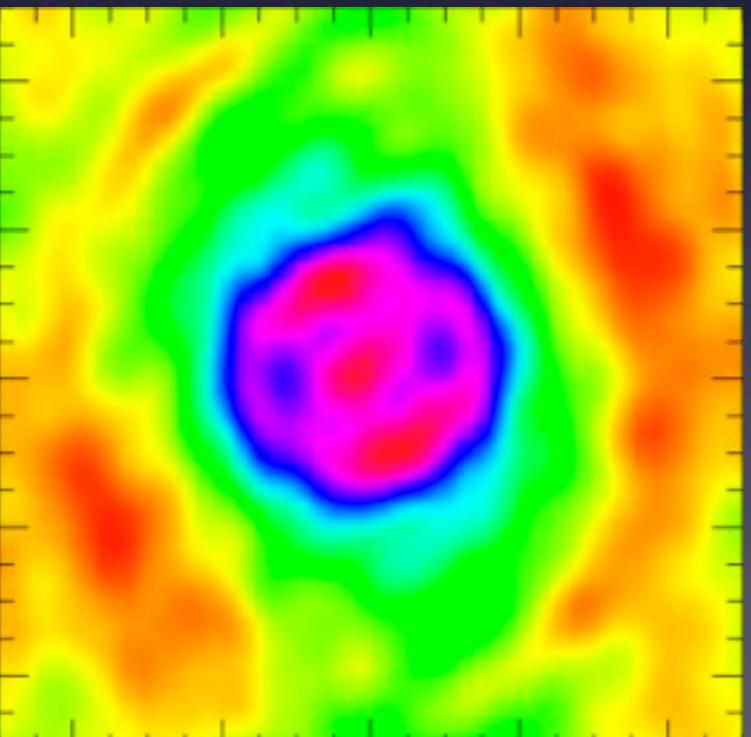
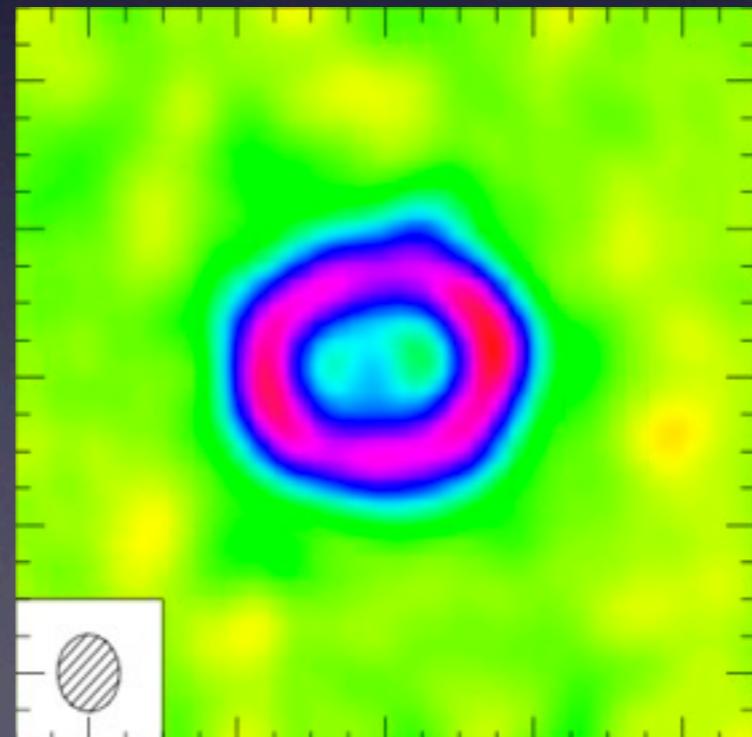
Imaging

Source

Dirty
imageDirty beam
(PSF)

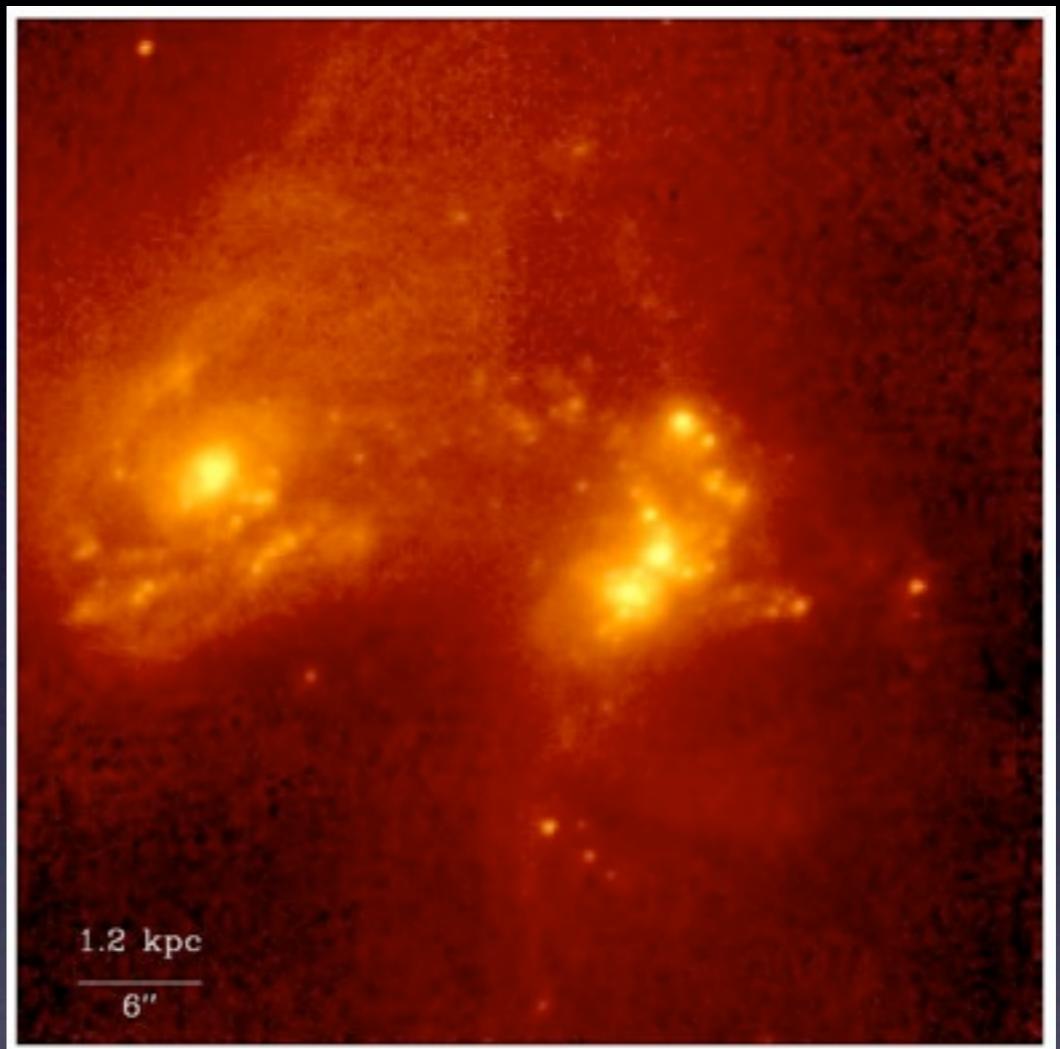
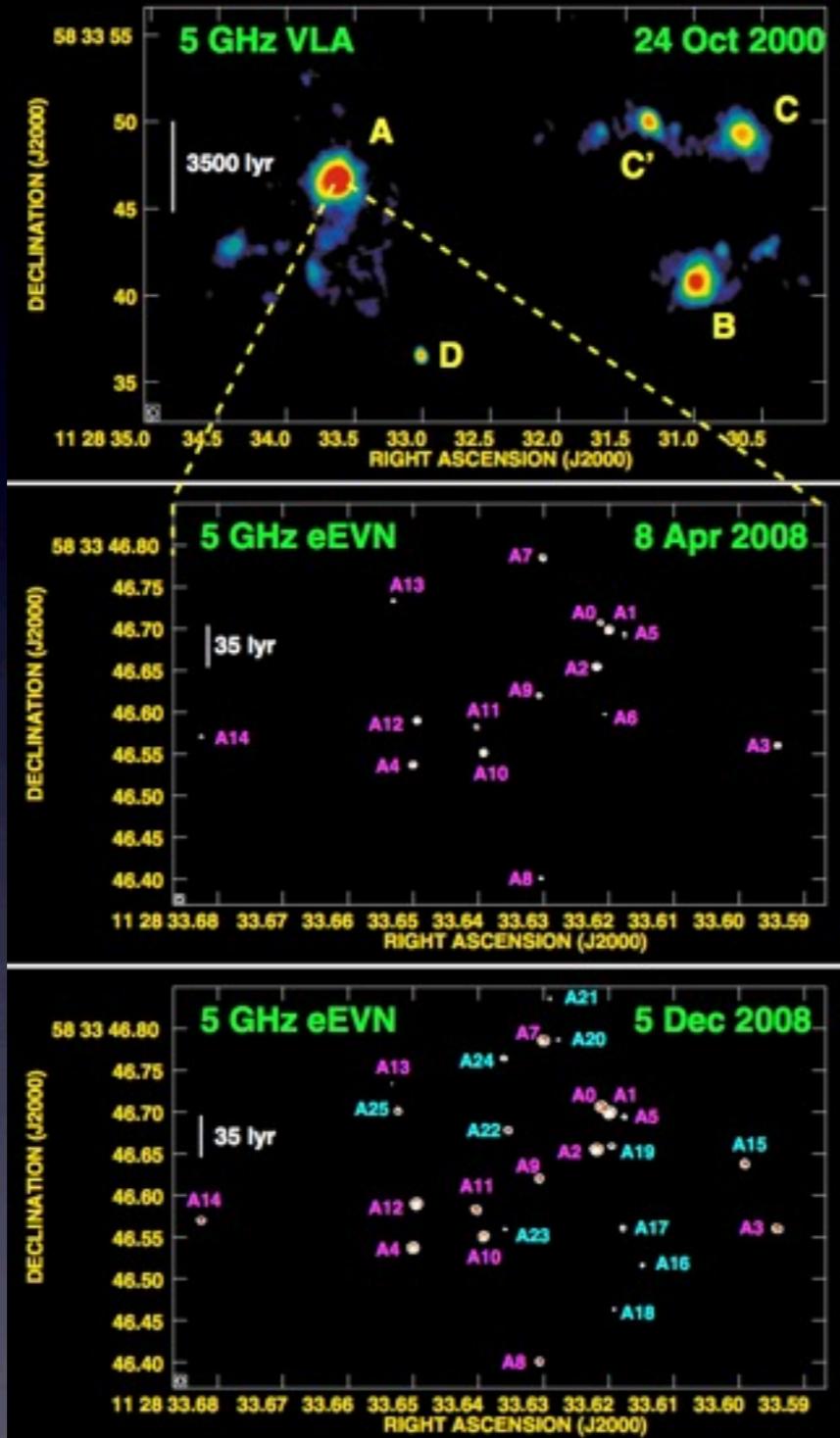
Imaging

Source

Dirty beam
(PSF)Dirty
imageClean
image

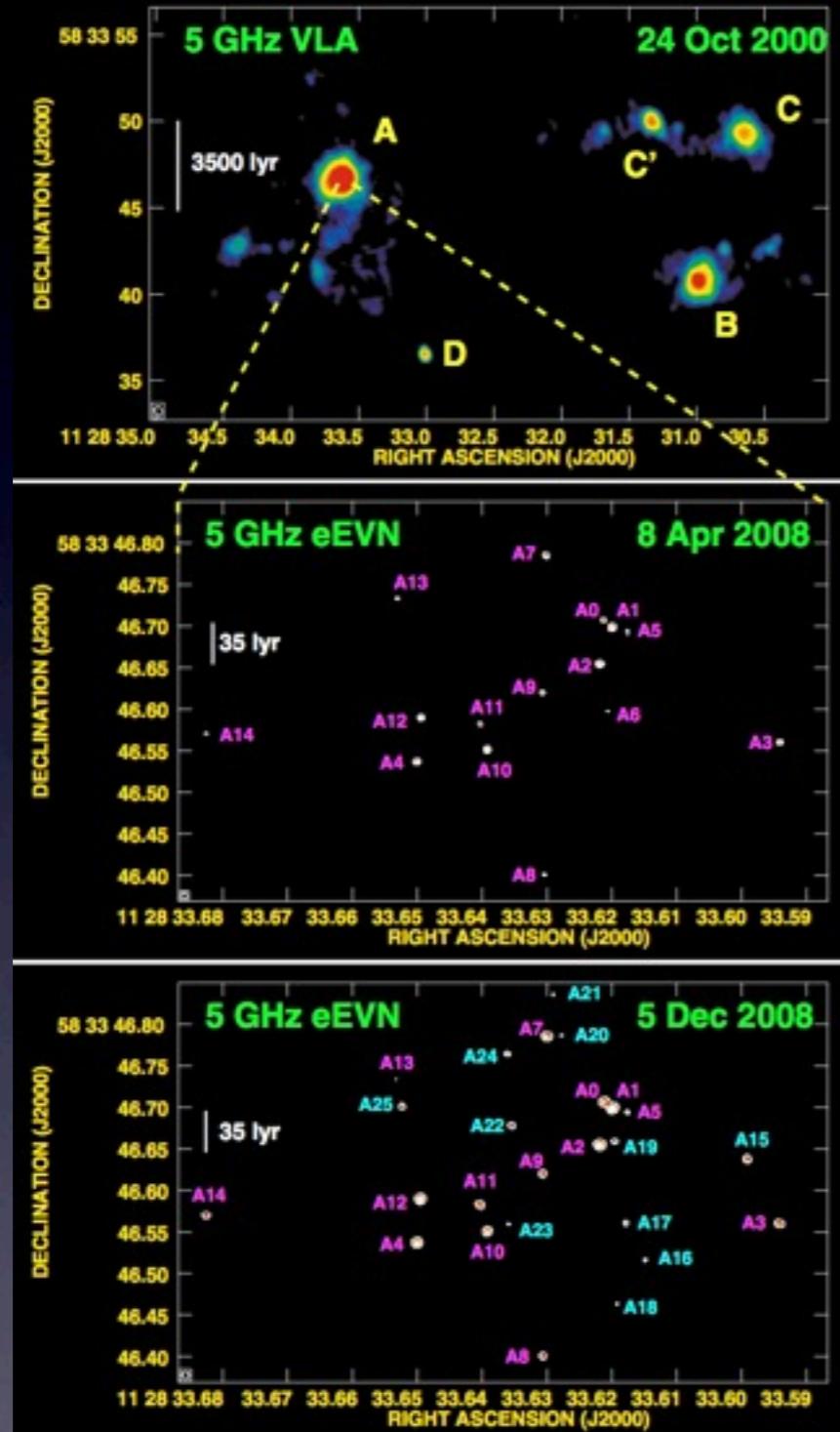
Some examples (LIRG biased!)

Arp 299: Hunting SNe & AGN

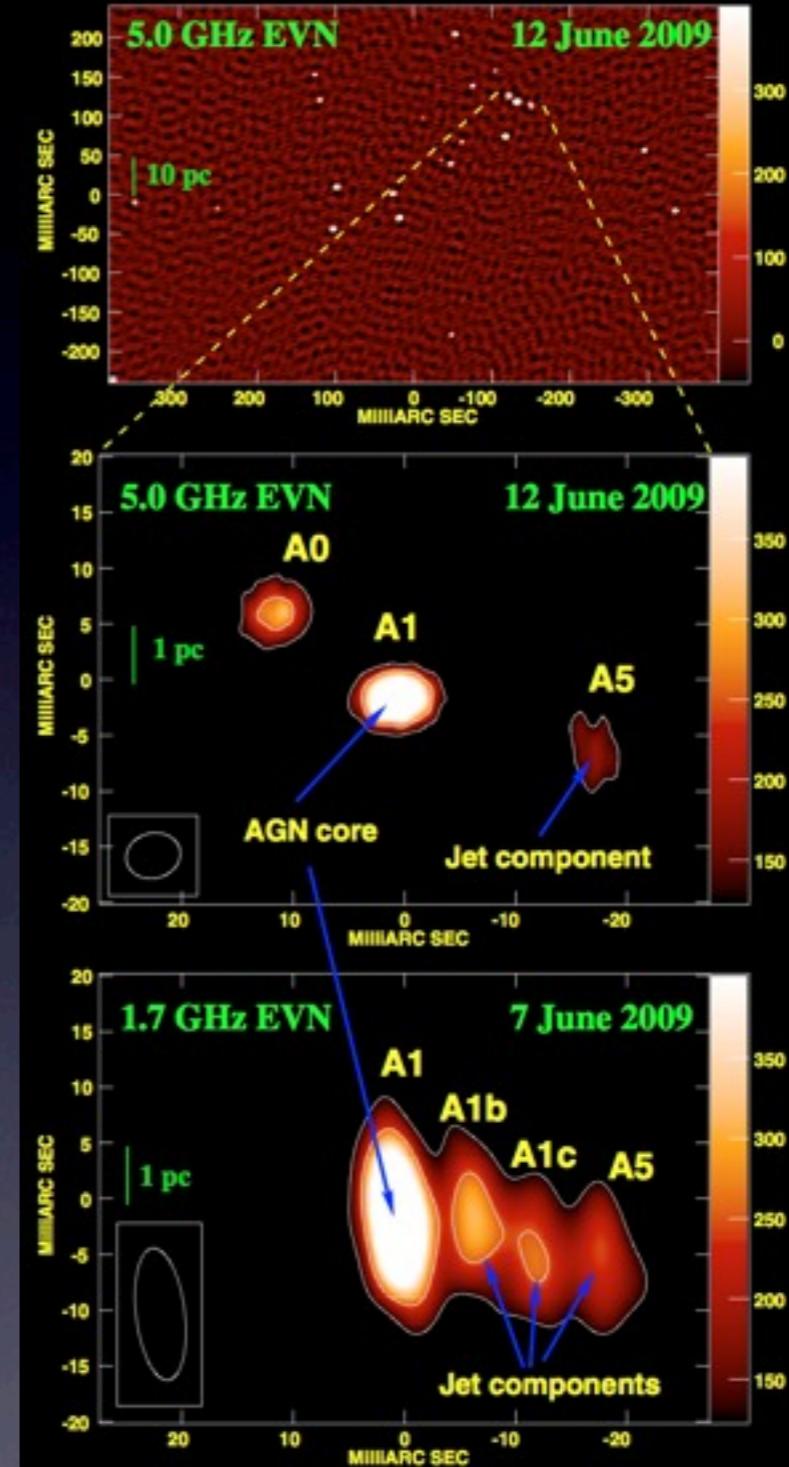


Pérez-Torres et al. 2009

Arp 299: Hunting SNe & AGN



Pérez-Torres et al. 2009



Pérez-Torres et al. 2010

IC 1623



3.6cm

814nm

435nm

¡Muchas gracias!



